

Comprehensive Evaluation and Compilation of H_3^+ Spectroscopy

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Since its initial detection in 1980, there have been 17 laboratory studies of the H_3^+ infrared spectrum, reporting 895 transitions from a variety of fundamental, overtone, combination, and hot bands. The results of these two decades of labor, however, are difficult to utilize. There is no comprehensive list of the observed H_3^+ transitions, and the literature contains errors in frequency measurement and assignment due to the inherent difficulty of the measurements and the complexity of the spectrum. This paper resolves these problems while assembling all of the data into a single source. We have reviewed all reported transitions of H_3^+ for reliability in frequency measurement and have reassigned them based on a comparison with recent theoretical calculations. We have also developed a complete labeling scheme for all energy levels below 9000 cm^{-1} , which alleviates the confusion in assigning H_3^+ transitions that results from the difficulty of labeling the rovibrational energy levels of a molecule with such strong mixing. Our comprehensive linelist was then used to produce a set of 526 experimentally determined energy levels, which facilitates direct comparison with theoretical calculations and prediction of the “forbidden” pure rotation spectrum of H_3^+ . © 2001 Elsevier Science

I. INTRODUCTION

H_3^+ plays important roles in many fields (1), including interstellar chemistry, the study of planetary ionospheres, and the theoretical calculation of rovibrational energy levels of polyatomic molecules. The detailed study of H_3^+ in these fields has only been possible because of the observation of infrared transitions of H_3^+ in laboratory discharges.

The laboratory detection (2) of the fundamental band $\nu_2 \leftarrow 0$ of H_3^+ opened the door to the detailed study of H_3^+ in astronomical sources. In dense interstellar clouds, spectra of H_3^+ have not only confirmed the general picture of ion-neutral chemistry but also allowed measurement of the physical conditions in the clouds (3, 4). In diffuse clouds, H_3^+ has been observed (5) to be far more abundant than predicted by chemical models. H_3^+ has also been observed in emission from several planetary ionospheres (6–9) and has been used to image the plasma activity of the Jovian ionosphere (10).

Continued lab work on other vibrational bands of H_3^+ has allowed a detailed comparison with theoretical predictions of rovibrational energy levels from variational calculations. The variational approach is particularly well suited to H_3^+ because this simple system (consisting of only three protons and two electrons) is amenable to detailed calculations.

Both astronomical spectroscopy and theoretical calculations of H_3^+ have advanced to the point where the quality of the existing laboratory database may soon hinder their progress. The state-of-the-art infrared spectrometers on large telescopes (11) have now

achieved resolving powers of $R \approx 75000$ (corresponding to a resolution of $\approx 0.03\text{ cm}^{-1}$ at $4\text{ }\mu\text{m}$). Soon, this resolution may approach the precision of the existing laboratory data, and the ability of astronomers to accurately measure Doppler shifts (which measure the velocities of molecular clouds and the motions of planetary ionospheres) will be impeded. On the computational side, *ab initio* calculations have produced highly accurate potential energy surfaces for H_3^+ which take into account adiabatic and nonadiabatic corrections to the Born–Oppenheimer approximation, as well as relativistic corrections (12–14). Variational calculations of energy levels using these potentials are said to have an accuracy of a few hundredths of a wavenumber (15). Increasingly accurate laboratory frequencies (as well as reliable assignments of spectral lines) are essential to evaluate the quality of the newest theoretical calculations.

In order to provide transition frequencies of H_3^+ to theorists and astronomers, 17 laboratory spectroscopic studies (2, 16–31) have been performed over the past two decades, resulting in the observation of over 800 different transitions. These experiments have probed a wide range of rotational and vibrational states in both emission and absorption using several different experimental techniques. The job of the laboratory spectroscopists has been a difficult one—many of the observations have been performed at the limits of sensitivity, making frequency measurements difficult. The assignment of H_3^+ transitions also poses a formidable task, due to strong mixing between rovibrational levels. Despite the best efforts of the spectroscopists, the literature still contains errors in frequencies as well as assignments.

Because the precision of theoretical and astronomical work is approaching that of the laboratory work, it is now important to correct these problems and produce a convenient and reliable

Supplementary data for this article are available on IDEAL (<http://www.idealibrary.com>) and as part of the Ohio State University Molecular Spectroscopy Archives (http://msa.lib.ohio-state.edu/jmsa_hp.htm).

collection of the laboratory data. In this work, we have reassigned all of the observed transitions, scrutinized the frequency measurements, and compiled a comprehensive list of transitions. Using this linelist, we have also calculated a set of experimentally determined energy levels for direct comparison with theory. This paper is intended to provide a convenient summary of H₃⁺ laboratory spectroscopy and replaces the outdated lists of Kao *et al.* (32), Majewski *et al.* (27), and Dinelli *et al.* (33).

II. BACKGROUND

II.1. Theoretical Background

The quantum numbers, symmetry restrictions, energy level structure, and selection rules for H₃⁺ have been discussed in detail elsewhere (34, 35). Here we provide a brief discussion of the basic concepts needed to understand the rovibrational spectroscopy of the ground electronic state of H₃⁺.

II.1.1. Quantum Numbers

The total angular momentum (F) and the parity (\pm) are the only completely rigorous quantum numbers of any molecule, as a consequence of the isotropy and inversion symmetry of free space. For H₃⁺, the total angular momentum F is the vector sum of the total nuclear spin angular momentum I and the angular momentum associated with the motion of the nuclei J . H₃⁺ contains three spin 1/2 protons, so I is either 1/2 (referred to as *para*) or 3/2 (*ortho*). Because the interaction between the nuclear spin and nuclear motion is extremely weak, I and J can be considered good quantum numbers, along with \pm .

In addition to these good quantum numbers, there are several approximately good quantum numbers which help us understand the behavior of H₃⁺ at low energies. These include ν_1 and ν_2 , which specify the number of quanta in the ν_1 and ν_2 vibrational modes, as well as the vibrational angular momentum ℓ (associated with the degenerate ν_2 mode), which takes values of $\nu_2, \nu_2 - 2, \dots, -\nu_2 + 2, \text{ or } -\nu_2$.

For most molecules, the projection of J onto the molecular symmetry axis (k) is a good quantum number. In H₃⁺, however, there is near degeneracy for levels with the same $|k - \ell|$ as a result of the form of the Coriolis interaction and the values of the B and C rotational constants.¹ These levels mix strongly by the l -resonance term, and it becomes useful to define a new quantum number $g \equiv k - \ell$ (37), which can be thought of as the portion of the projection of J onto the molecular axis that is due to the rotation of the molecular frame. Because the energy

¹ This near degeneracy is evident from the first three terms in the rotational energy expression, $E_{rot} \approx BJ(J+1) + (C-B)k^2 - 2C\zeta k\ell$. Consider the levels $|J, k, \ell\rangle = |J, g + \ell, \ell\rangle$ and $|J, g - \ell, -\ell\rangle$. In this approximation, the energy separation between these levels is $4g\ell(C - B - \zeta C)$, which is nearly zero because $B \approx 2C$ (due to the planarity of the molecule) and $\zeta = -1$ (for the triatomic equilateral triangle (36)). Because these levels have the same symmetry and a small energy difference, they will be strongly mixed (the mixing is particularly strong for $|\ell| = 1$ due to the ℓ -resonance term $q(q_+^2 J_+^2 + q_-^2 J_-^2)/4$).

does not depend on the sign of g , $G \equiv |g|$ is usually used. G is a reasonably good quantum number at low energies.

II.1.2. Symmetry Restrictions and Energy Level Structure

The Pauli principle requires the total wavefunction to be antisymmetric with respect to (12) permutation of any two protons and symmetric upon cyclic permutation (123) of all three protons (i.e., the total wavefunction must belong to the A_2 representation). This requirement imposes a relationship between the nuclear spin modification and the quantum number G : when $I = 3/2$, only $G = 3n$ levels (and when $I = 1/2$, only $G = 3n \pm 1$ levels) have the proper symmetry. Additionally, certain $G = 0$ levels (most notably $J = \text{even}$ and $G = 0$ in the ground vibrational state) do not satisfy the symmetry requirement and therefore do not exist.

The energy level structure of H₃⁺ is similar to that of a normal oblate symmetric top (when plotted versus G rather than k) except that certain levels come in pairs. These pairs are due to the two ways of forming the same G with different values of k and ℓ . Energy level diagrams for the ground and $\nu_2 = 2, \ell = 2$ vibrational states are plotted in Fig. 1.

II.1.3. Selection Rules

We first consider the electric dipole selection rules for the good quantum numbers I, J , and \pm . Because the dipole operator $\hat{\mu}$ does not operate on the nuclear spin wavefunctions, the nuclear spin must not change during a radiative transition, and thus the selection rule for I is $\Delta I = 0$. The total angular momentum F must obey the ‘‘triangle rule’’ for angular momentum addition [as $\hat{\mu}$ is a tensor of rank one and transforms as the spherical harmonics $Y_{1,0}$ and $Y_{1,\pm 1}$ (38)], and thus $\Delta F = 0$ or ± 1 , and $0 \not\leftrightarrow 0$. Since $\Delta I = 0$, the ‘‘triangle rule’’ also applies to J : $\Delta J = 0$ or ± 1 , and $0 \not\leftrightarrow 0$. The selection rule for the parity can be obtained by considering that the matrix elements of the dipole operator, $\langle \psi_f | \hat{\mu} | \psi_i \rangle$, must be totally symmetric. Since $\hat{\mu}$ changes sign with the inversion operation ($\vec{r} \rightarrow -\vec{r}$), the parity of the initial and final wavefunctions must be different ($+\leftrightarrow -$).

The selection rule for g can be found by examining the symmetry of the wavefunctions (34) with respect to the cyclic permutation (123):

$$(123) |J, k, \ell\rangle = e^{\frac{2\pi i}{3}(k-\ell)} |J, k, \ell\rangle. \quad [1]$$

Combining this with the required invariance of the transition dipole moment matrix elements with respect to (123), we see that

$$(123) \langle J', k', \ell' | \hat{\mu} | J'', k'', \ell'' \rangle = e^{\frac{2\pi i}{3} \{ (k'' - \ell'') - (k' - \ell') \}} \langle J', k', \ell' | \hat{\mu} | J'', k'', \ell'' \rangle \quad [2]$$

which is only invariant when $\Delta g = (k'' - \ell'') - (k' - \ell') = 3n$.

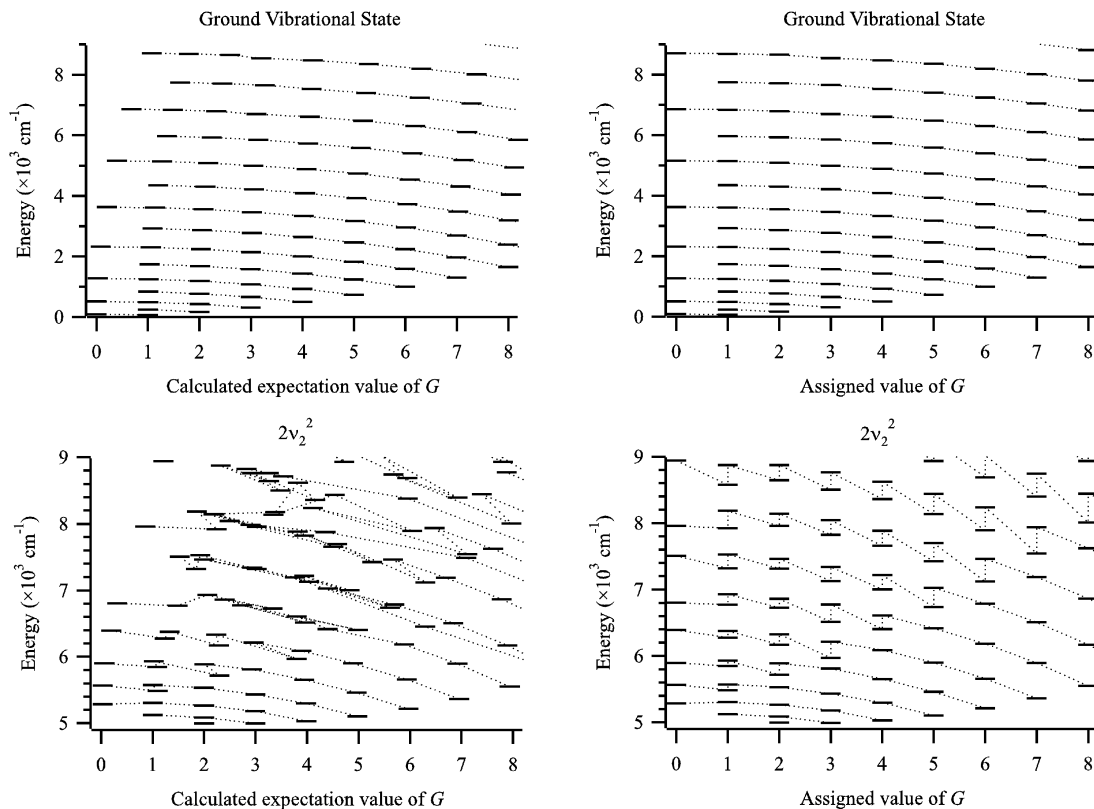


FIG. 1. Energy level diagrams of H_3^+ for the ground vibrational state (top two plots) and the $v_2 = 2, \ell = 2$ vibrational state (lower plots). (Dotted lines connect levels with the same J .) On the two left plots, Watson's calculated energy levels are plotted against the expectation values of G (see Section III.1 for comments on this calculation). On the right plots, the same energy levels are plotted versus our assigned values of G , presented in Table 3. Before being relabeled, the levels in the ground state look much like a classic oblate symmetric top with a small distortion in G at higher energy. The excited vibrational states are highly perturbed, and this mixing is the reason for many of the mislabeled transitions the literature. Once the G values are assigned, the energy level diagram looks relatively well behaved. Similar figures for every vibrational state below 9000 cm^{-1} are available online (39).

The possible selection rules for k (the projection of J) can be derived from the "triangle rule" to be $\Delta k = 0$ or ± 1 . Because the parity is linked to k by the symmetry of the wavefunctions (34) with respect to the inversion (E^*) operation

$$E^*|J, k, \ell\rangle = (-1)^k|J, k, \ell\rangle \quad [3]$$

and because the parity selection rule is $+\leftrightarrow -$, Δk must be odd, restricting its selection rule to $\Delta k = \pm 1$.

The selection rule for ℓ depends on those of g and k :

$$\begin{aligned} \Delta g &= g' - g'' = (k' - \ell') - (k'' - \ell'') = \Delta k - \Delta \ell \\ \Delta \ell &= \Delta k - \Delta g = (\pm 1) - (\pm 3n) \\ \Delta \ell &\neq 3n. \end{aligned} \quad [4]$$

For transitions with $\Delta \ell = \pm 1$ (e.g., the $\Delta v_2 = 1$ fundamental and hot bands), Δg must be 0, and for transitions with $\Delta \ell = \pm 2$ ($\Delta v_2 = 2$ overtone bands), Δg must be ± 3 .

It should be kept in mind that the selection rules for g , k , and ℓ are not rigorous, because these quantum numbers are not rigorous. For example, the $\Delta k = \pm 1$ selection rule can break

down due to mixing, but $\Delta k = \text{odd}$ is maintained because it is based on the parity, which is a rigorously good quantum number.

Finally, we consider the selection rules (which are only approximate) for the vibrational quantum numbers. The ν_1 normal mode is totally symmetric and therefore has the selection rule² $\Delta \nu_1 = 0$. For the symmetry-allowed ν_2 mode, the selection rule $\Delta \nu_2 = \pm 1$ holds in the approximation of using harmonic oscillator wavefunctions and only the first order term in the Taylor series expansion of $\hat{\mu}$. Because the H_3^+ potential and dipole operator are very anharmonic, this is a poor approximation, and transitions with $\Delta \nu_2 > 1$ have reasonable intensity.

²The selection rule $\Delta \nu_1 = 0$ requires some qualification. The band $\nu_1 \leftarrow 0$ has the vibrational symmetry $A_1 \leftarrow A_1$, which is forbidden, but can become allowed via the rotation-vibration interaction. Bands such as $\nu_1 + \nu_2 \leftarrow \nu_2$ are qualitatively different in that they have vibrational symmetry $E \leftarrow E$, which is allowed through a vibrational interaction alone (24). All of these bands are very weak because the change in the dipole moment is small upon excitation of totally symmetric mode, but $\nu_1 \leftarrow 0$ is by far the weakest, since it relies on an accidental degeneracy for the rotation-vibration interaction to be effective.

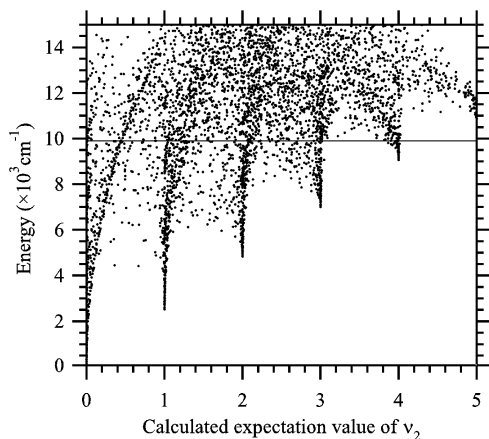


FIG. 2. Rovibrational energy of H₃⁺ plotted against the expectation value of the approximate quantum number ν_2 . Energies and expectation values (see Section III.1 for comments on this calculation) are from the calculations of Watson (27, 54). The solid line is drawn at the barrier to linearity. This plot shows that at lower energies, ν_2 describes the system quite well, but at higher energies, the amount of mixing increases to the point where the value of this approximate quantum number has little meaning.

II.1.4. Effects of Mixing

Levels with the same good quantum numbers [J , I , and \pm] can mix with one another. The strength of the mixing is inversely proportional to the energy separation between the two unmixed levels, so that strong mixing is increasingly common at higher energies where the levels tend to be more closely spaced. When mixing occurs, the energy levels are shifted from the oblate symmetric top energy pattern and the approximate quantum numbers break down. The mixed states can no longer be described by single integral values of g , ν_1 , ν_2 , and ℓ , but can be described by their expectation values, which are linear combinations of the quantum numbers of the unmixed values. The extent of the mixing can be visualized by plotting the energy versus the expectation values of the quantum numbers (see Figs. 1 and 2). When the energy levels are not significantly mixed, the expectation values of their quantum numbers will be nearly integral.

The selection rules for mixed levels incorporate the selection rules of each of the levels involved in the mixing. A consequence of this mixing is the appearance of additional lines in the spectrum—forbidden transitions effectively borrow intensity from allowed ones. One example of these forbidden transitions is the pure rotational transitions, which obey the selection rule $\Delta\ell = 0$. These transitions are weak, but should be observable experimentally (see Section IV.1). Each transition's intensity will depend on the magnitude of mixing, and must be treated on a line by line basis. The topic of rovibrational level mixing and the breakdown of the approximate quantum numbers is discussed further in Section III.1.

II.2. Previous Laboratory Work

Many infrared absorption and emission studies of H₃⁺ have been performed in the laboratory over the past two decades to characterize the rovibrational spectrum and energy levels of H₃⁺

(see (40) for a recent review). These works were considered in our analysis and in this section we briefly summarize each of them (see Table 1). It should be noted that the predissociation spectrum of H₃⁺ has also been measured in the laboratory (41–43) and is the subject of a recent review (44). This subject, however, lies outside the scope of this paper.

The infrared spectrum of H₃⁺ was initially sought after by Herzberg in the mid 1960s when it became clear that H₃⁺ did not possess a stable excited electronic state (45). In the course of this work he observed a group of emission lines near 3600 cm⁻¹ in hydrogen hollow-cathode discharges. These were eventually identified by Watson as emission lines of neutral H₃, which are produced in excited states after dissociative recombination of H₅⁺ with electrons. It was not until 1980 that Oka (2) observed in absorption the first 15 lines of the ν_2 fundamental of H₃⁺ between 2450 and 2950 cm⁻¹. His success was made possible by the development of the broadly tunable difference frequency (DF) spectrometer by Pine (46) and the use of the long positive column discharge by Woods (47). In his search he scanned roughly 500 cm⁻¹, a feat only possible with the DF laser. To increase the sensitivity, he frequency modulated the radiation and achieved a signal-to-noise ratio of ~ 30 for the strongest H₃⁺ lines. In the first studies, a liquid-nitrogen-cooled positive column discharge was used to produce H₃⁺ in a pure hydrogen discharge.

In the year that followed, Oka (16) was able to extend his observations to higher J levels (to a total of 30 lines) by studying a warmer, ice-cooled discharge. By 1984, the frequency coverage was expanded by making adjustments to the DF laser and by using diode lasers. Two major advances in sensitivity also occurred. It was found that modulating the discharge current (concentration modulation (48)) or applying an AC field across a positive column discharge (velocity modulation (49)) could substantially improve the sensitivity. The combination of these improvements enabled Watson, Oka, and co-workers (17) to observe an additional 16 $\nu_2 \leftarrow 0$ transitions, bringing the total up to 46.

All of the observed lines up to this point were from levels with $J \leq 5$, and it was expected that large perturbations would occur at higher J between the ν_1 and ν_2 states. With this in mind, Majewski *et al.* (18) in Ottawa constructed a high-pressure hollow-cathode discharge coupled to a Fourier transform infrared (FTIR) spectrometer to observe the emission of H₃⁺ in a hydrogen discharge. With an ingeniously designed hollow-cathode discharge and a pressure discrimination method, this technique turned out to be very effective, nearly tripling the number of observed lines and probing levels up to $J' = 10$. Many perturbations were indeed observed, and they provided a new test for theoretical calculations. Many additional emission features were recorded around 2 μm . At the time, the authors could not rule out the possibility that the 2- μm lines were due to Rydberg H₂ or H₃ neutral transitions, and consequently the lines were not reported.

In 1987 Trafton *et al.* (6) and in 1988 Drossart *et al.* (7) stumbled upon a rich set of unidentified emission features at 2 μm while observing H₂ emission in Jupiter. This prompted

TABLE 1
Summary of the Laboratory Spectroscopic Studies of H₃⁺

Label	cm ⁻¹	Observed	Assignment	Technique ^a	Reference
Oka80	2450–2950 [†]	15 lines	$\nu_2 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, DF laser, FM detection	T. Oka, <i>Phys. Rev. Lett.</i> 45 , 531–534 (1980).
Oka81	2450–3030 [†]	30 lines (15 new)	$\nu_2 \leftarrow 0$	<i>l</i> -N ₂ and ice-water cooled positive column, DF laser, FM detection	T. Oka, <i>Phil. Trans. R. Soc. Lond. A</i> 303 , 543–549 (1981).
Wat84	2210–3030 [†]	46 lines (16 new)	$\nu_2 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, DF and diode lasers, VM and CM detection	J. K. G. Watson, S. C. Foster, A. R. W. McKellar, P. Bernath, T. Amano, F. S. Pan, M. W. Crofton, R. S. Altman, and T. Oka, <i>Can. J. Phys.</i> 62 , 1875–1885 (1984).
Maj87	1800–3300	113 lines (67 new)	$\nu_2 \leftrightarrow 0$	Water cooled, high-pressure hollow cathode, FTIR emission, and diode laser absorption	W. A. Majewski, M. D. Marshall, A. R. W. McKellar, J. W. C. Johns, and J. K. G. Watson, <i>J. Mol. Spectrosc.</i> 122 , 341–355 (1987).
Maj89	4500–5100	47 new lines	$2\nu_2^2 \rightarrow 0$	Water cooled, high-pressure hollow cathode, FTIR emission	W. A. Majewski, P. A. Feldman, J. K. G. Watson, S. Miller, and J. Tennyson, <i>Astrophys. J.</i> 347 , L51–L54 (1989).
Nak90	2400–2800	12 re-measured	$\nu_2 \leftarrow 0$	Water cooled hollow cathode, FTIR absorption	T. Nakanaga, F. Ito, K. Sugawara, H. Takeo, and C. Matsumura, <i>Chem. Phys. Lett.</i> 169 , 269–273 (1990).
Baw90	2080–2950 [†]	14 new lines 70 new lines 14 new lines 21 new lines 136 new lines	$\nu_2 \leftarrow 0$ $2\nu_2^2 \leftarrow \nu_2$ $2\nu_2^0 \leftarrow \nu_2$ $\nu_1 + \nu_2 \leftarrow \nu_1$ unassigned	<i>l</i> -N ₂ cooled positive column, DF laser, VM detection	M. G. Bawendi, B. D. Rehfuss, and T. Oka, <i>J. Chem. Phys.</i> 93 , 6200–6209 (1990).
Xu90	4550–6000 [†]	34 lines (7 new)	$2\nu_2^2 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, DF laser, VM detection	L.-W. Xu, C. M. Gabrys, and T. Oka, <i>J. Chem. Phys.</i> 93 , 6210–6215 (1990).
Lee91	6860–6900 [†]	4 new lines	$3\nu_2^1 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, diode laser, VM detection	S. S. Lee, B. F. Ventrudo, D. T. Cassidy, T. Oka, S. Miller, J. Tennyson, <i>J. Mol. Spectrosc.</i> 145 , 222–224 (1991).
Xu92	2400–3300 [†]	9 new lines 21 new lines 30 new lines 13 new lines 89 new lines	$\nu_1 \leftarrow 0$ $\nu_1 + \nu_2 \leftarrow \nu_2$ $\nu_2 \leftarrow 0$ $2\nu_2^2 \leftarrow \nu_2$ unassigned	<i>l</i> -N ₂ cooled positive column, DF laser, VM detection	L.-W. Xu, M. Rösslein, C. M. Gabrys, and T. Oka, <i>J. Mol. Spectrosc.</i> 153 , 726–737 (1992).
Ven94	6800–7270 [†]	15 lines (11 new)	$3\nu_2^1 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, diode laser, VM detection	B. F. Ventrudo, D. T. Cassidy, Z. Y. Guo, S. Joo, S. S. Lee, and T. Oka, <i>J. Chem. Phys.</i> 100 , 6263–6266 (1994).
Uy94	2690–3580 [†]	75 lines (37 new)	$\nu_2 \leftarrow 0$	Water cooled positive column, DF laser, VM detection	D. Uy, C. M. Gabrys, M.-F. Jagod, and T. Oka, <i>J. Chem. Phys.</i> 100 , 6267–6274 (1994).
Maj94	1800–2550 2900–5000	52 new lines 9 new lines 12 new lines 31 new lines 16 new lines 2 new lines 1 new line	$\nu_2 \rightarrow 0$ $2\nu_2^2 \rightarrow 0$ $\nu_1 + \nu_2 \rightarrow \nu_1$ $2\nu_2^2 \rightarrow \nu_2$ $2\nu_2^0 \rightarrow \nu_2$ $3\nu_2^3 \rightarrow \nu_2$ $3\nu_2^3 \rightarrow 2\nu_2^2$	Water cooled, high-pressure hollow cathode, FTIR emission	W. A. Majewski, A. R. W. McKellar, D. Sadovskii, and J. K. G. Watson, <i>Can. J. Phys.</i> 72 , 1016–1027 (1994).
McK98	2450–2850	27 re-measured	$\nu_2 \leftarrow 0$	Refrig. methanol cooled hollow cathode, FTIR absorption	A. R. W. McKellar and J. K. G. Watson, <i>J. Mol. Spectrosc.</i> 191 , 215–217 (1998).
Joo00	~1550	1 new line	$\nu_2 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, diode laser, VM detection	S. Joo, F. Kühnemann, M.-F. Jagod, and T. Oka, <i>The Royal Society Discussion Meeting on Astronomy, Physics and Chemistry of H₃⁺</i> , London, February 9–10 (2000) (poster).
McC00	7850–8170	28 new lines 2 new lines	$\nu_1 + 2\nu_2^2 \leftarrow 0$ $2\nu_1 + \nu_2 \leftarrow 0$	<i>l</i> -N ₂ cooled positive column, diode laser, VM detection	B. J. McCall and T. Oka, <i>J. Chem. Phys.</i> 113 , 3104–3110 (2000).
Lin01	3000–4200	6 lines (5 new) 22 lines (10 new) 5 lines (4 new) 76 lines (44 new) 4 lines (3 new) 1 new line 25 lines (9 new) 14 lines (7 new) 2 new lines 1 re-measured 3 lines (2 new) 3 lines (1 new) 6 lines (5 new)	$\nu_2 \leftarrow 0$ $\nu_1 \leftarrow 0$ $2\nu_2^0 \leftarrow \nu_2$ $2\nu_2^2 \leftarrow \nu_2$ $2\nu_2^2 \leftarrow \nu_1$ $2\nu_2^0 \leftarrow \nu_1$ $\nu_1 + \nu_2 \leftarrow \nu_2$ $\nu_1 + \nu_2 \leftarrow \nu_1$ $2\nu_1 \leftarrow \nu_1$ $\nu_1 + 2\nu_2^2 \leftarrow \nu_1 + \nu_2$ $3\nu_2^3 \leftarrow 2\nu_2^2$ $3\nu_2^2 \leftarrow 2\nu_2^0$ unassigned	<i>l</i> -N ₂ cooled positive column, CCL, VM detection	C. M. Lindsay, R. M. Rade, Jr., and T. Oka, <i>J. Mol. Spectrosc.</i> 210 , 51–59 (2001).

^a Abbreviations used in this column are defined as follows: FM = frequency modulation; VM = velocity modulation; DF = difference frequency; CM = concentration modulation; CCL = color center laser.

[†] Region was not scanned continuously.

the Ottawa group to revisit the 2 μm lines observed with the FTIR emission apparatus, and after a month, Watson assigned many of the FTIR and Jovian features to the 2ν₂² ← 0 band of H₃⁺. The new-found confidence in these assignments was based upon the latest *ab initio* calculations of Miller and Tennyson (50) as well as the yet-to-be published work on the 2ν₂² ← ν₂ hot band by Bawendi *et al.* in Chicago (see below). Once assigned, 47 lines of the 2ν₂² ← 0 band were reported from the FTIR studies (19).

Quite apart from the work in Chicago and Ottawa, Nakanaga and co-workers in 1990 successfully performed the first FTIR absorption spectroscopy of molecular ions, including 12 fundamental transitions of H₃⁺ (20). While all of these lines had been observed initially 9 years earlier, this work represented the first accurate measurement of the relative absorption intensities.

After several years of refining their technique of DF laser/velocity modulation spectroscopy of carbocations (51) and introducing the helium-dominated positive column discharge, the Chicago group revisited H₃⁺ with a tremendous increase in sensitivity. The next 5 years brought seven experiments which substantially increased the number of probed levels. Initially, Bawendi *et al.* (21) observed lines of the 2ν₂² ← ν₂, 2ν₂⁰ ← ν₂, and ν₁ + ν₂ ← ν₁ hot bands as well as 14 new fundamental lines and 136 lines which they could not confidently assign. Shortly after, Xu *et al.* (22) observed the 2ν₂² ← 0 band, though they only observed seven lines not covered in Majewski's work. Two years later, Xu *et al.* (24) reported the ν₁ ← 0 and ν₁ + ν₂ ← ν₂ forbidden bands, as well as more transitions from the ν₂ ← 0 and 2ν₂² ← ν₂ bands, and additional unassignable lines. Advances in external cavity near infrared diode lasers enabled the scanning to be extended to higher frequency allowing the second overtone, 3ν₂¹ ← 0, to be observed (23, 25). During this period a diode laser was also used to measure the lowest frequency line to date (the ν₂ ← 0 P(12, 12) at 1546.901 cm⁻¹), though this was only reported recently (29). Finally, Uy and co-workers (26) recorded an H₃⁺ spectrum with a water-cooled cell and observed highly excited rotational levels of the ν₂ fundamental, up to J' = 16.

With the large amount of experimental data made available, Watson (52, 27) and, independently, Dinelli *et al.* (53) produced empirically fitted potentials which were used to calculate more accurate transition frequencies (see Section 4.2 for more details). All of the experimental data available at the time, as well as some newly measured FTIR emission lines (27), were collected and included in these calculations. These calculations proved to be essential to understanding the unassigned lines in Bawendi's and Xu's (1992) data, which were assigned in subsequent papers (27, 33).

Four years later, McKellar and Watson recorded a clean broadband absorption spectrum of H₃⁺ with an FTIR spectrometer. Their work was similar to that of Nakanaga *et al.* 8 years earlier, but achieved about a factor of 10 improvement in signal-to-noise ratio enabling them to observe 27 lines of the ν₂ fundamental. This beautiful spectrum has been published in its entirety in a letter (28).

No new data were reported until the year 2000 when McCall and Oka (30) recorded lines from the ν₁ + 2ν₂² ← 0 and 2ν₁ + ν₂ ← 0 combination bands using an external cavity diode laser and the velocity modulation/positive column discharge technique. Thirty lines were observed from these two bands, probing the highest vibrational states observed to date.

Most recently, Lindsay and co-workers (31) used a computer-controlled color center laser (CCL) spectrometer to continuously scan H₃⁺ from 3000–4200 cm⁻¹. The improved sensitivity of their spectrometer and the hottest discharge to date enabled them to study very high rovibrational levels. A total of 96 new transitions were observed from a variety of hot, overtone, fundamental, and forbidden bands—some probing rovibrational levels in the vicinity of the barrier to linearity.

A total of 895 unique transitions of H₃⁺ have been reported over the past 21 years, probing every vibrational state below the barrier to linearity (except 3ν₁ and 4ν₂). Many of these transitions have been recorded multiple times by multiple techniques with multiple sensitivities. This work was only possible with substantial advancements to the sensitivity of the experiments, which improved dramatically over the last 21 years. It is interesting to note that if the sensitivity of the latest studies were applied to the transitions observed by Oka in 1980, the signal-to-noise ratio would be 3000–6000—a two-orders-of-magnitude improvement over the initial spectrum!

III. ANALYSIS AND RESULTS

In this section we explain our efforts to assign approximate quantum numbers to every energy level below 9000 cm⁻¹, evaluate the assignment and quality of every reported laboratory absorption and emission transition, and determine energy levels from the experimental transitions. Most of the results of this work are tabulated here in print, but an electronic version of the complete work (tables and figures) is available online (39).

III.1. Labeling of Rovibrational Levels below 9000 cm⁻¹

Much of the confusion in “assigning” transitions in the literature is based upon energy level labeling and not the actual assignment of the transitions. This distinction is important—most of the assignments (that is the identification of an observed transition based on a particular calculated transition between two levels with a similar frequency and intensity) were correctly made, but there has been confusion in the naming of the transition and energy levels which were involved in the transition. Before each transition can be labeled, every energy level must be given a unique label. Below the barrier to linearity, the approximate quantum numbers *G*, ν₁, ν₂, and *ℓ* are reasonably good and can be used to label rovibrational energy levels. Many of these levels mix, and the resulting levels have character of two or more levels with different values of the approximate quantum numbers. In most cases this mixing is not complete, and

each mixed level can be labeled by the quantum numbers of the dominant unmixed level.

Theoretical calculations of energy levels provide only the quantum numbers J , parity, and (in most cases) I —the assignment of the approximate quantum numbers to each level must be done manually. This task would be nearly impossible without the help of the expectation values calculated by Watson (54). Even with the expectation values, this task is not easy. One can appreciate the difficulty in applying labels by examining the energy dependence of various expectation values. In Fig. 2 the energy is plotted versus the expectation value of v_2 . At energies close to each band origin, the values of $\langle v_2 \rangle$ are very close to integers. As the energy approaches $10\,000\text{ cm}^{-1}$ however, many levels have expectation values in between the integer values, as a result of mixing. Likewise, we can look at the expectation values of G (Fig. 1). While well behaved in the ground vibrational state (top left), G becomes extremely mixed at higher vibrational states (bottom left). By carefully considering the energy, the expectation values of G , v_1 , v_2 , and ℓ , and the values of J , I , and parity of all of the levels simultaneously, it is possible to assign integral values of G , v_1 , v_2 , and ℓ for each energy level (Fig. 1 right, top and bottom). We have produced energy level diagrams similar to those in Fig. 1 for all vibrational states below 9000 cm^{-1} , and these are available online (39). Please note that these calculated expectation values are only approximate and were performed with the intention of forming a qualitative picture of the nature of the energy levels (54).

The five quantum numbers J , G , v_1 , v_2 , and ℓ are not sufficient to uniquely label each level. For levels with $\ell \neq 0$ and $(J - |\ell|) \geq G \geq 1$ there are two ways to form the same G for different values of k . Take as an example a level where $\ell = 1$ and $J = 2$. Since $G \equiv |k - \ell|$, both $k = 0$ and $k = 2$ make $G = 1$. These levels always differ in energy, and we have distinguished them by assigning a “ u ” and an “ l ” to the upper and lower energy level, respectively (30). In the earliest papers, these levels were designated by “I” or “II” (2, 16) or later with “+” or “−”. Also used was the U notation initially defined by Watson (34) as +1 for “ u ” levels and −1 for “ l ” levels of the v_2 vibrational state. Later, Miller and Tennyson (50) extended the notation to arbitrary v_2 by redefining $U = +|\ell|$ for “ u ” and $U = -|\ell|$ for “ l ” levels. We have abandoned these other notations due to the confusion with the value of the real quantum numbers ℓ and parity. We instead use the symbol (40)

$$(J, G)\{u | l\} \quad [5]$$

to label individual rotational levels within the vibrational state,

$$v_1 v_1 + v_2 v_2^{|\ell|} \quad \text{or} \quad v_1 v_2^{|\ell|}. \quad [6]$$

A small number of levels are so badly mixed that the assignment of their approximate quantum numbers is almost arbitrary. In some cases the expectation value of one quantum number suggests an assignment to one vibrational state, while another

TABLE 2
Heavily Mixed Rovibrational Levels of H_3^+ below 9000 cm^{-1}
(Each Row Corresponds to a Set of Mixed Levels)

$v_1 + 2v_2^0(5,4)$	$v_1 + 2v_2^0(5,2)$	$v_1 + 2v_2^2(5,2)l$	$v_1 + 2v_2^2(5,2)u$
$3v_2^3(6,2)l$	$3v_2^1(6,2)u$	$v_1 + 2v_2^2(6,3)$	
$2v_2^2(7,0)$	$v_1 + v_2(7,3)l$		
$v_1 + v_2(7,3)u$	$v_1(7,6)$		
$v_1 + v_2(7,5)l$	$2v_2^2(7,4)u$		
$v_1(8,6)$	$v_1 + v_2(8,3)u$		
$2v_2^0(8,2)$	$2v_2^2(8,2)u$		
$v_1 + v_2(8,5)l$	$v_1(8,8)$		
$v_2(9,2)u$	$v_1(9,5)$		
$2v_2^2(9,5)l$	$2v_2^0(9,1)$		
$2v_2^2(9,1)u$	$v_1 + v_2(9,2)l$		
$2v_2^0(9,3)$	$2v_2^2(9,3)u$		
$v_1(10,5)$	$v_2(10,2)u$		
$2v_2^0(10,2)$	$2v_2^2(10,4)$		
$v_1 + v_2(11,9)l$	$2v_2^2(11,6)$		
$v_2(12,6)l$	$v_1(12,9)$		

suggests a different vibrational state. Table 2 lists all of the badly mixed levels below 9000 cm^{-1} . Above 9000 cm^{-1} , the density of states becomes quite large and severe mixing occurs for many levels with $J \gtrsim 5$. For these levels, it probably is not useful to assign approximate quantum numbers. It should still be possible, however, to label low J levels with approximate quantum numbers because their density of states is much lower and the mixing will be less complete. At very high energy, even the low J levels will mix and their approximate quantum numbers will eventually fail. When this occurs, levels will have to be labeled by the only good quantum numbers (J , I , \pm) and the energy-ordering index n .

The results of our energy level labeling scheme are listed in Table 3. Every rovibrational level below 9000 cm^{-1} has been labeled with J , G , v_1 , v_2 , ℓ , and (where appropriate) u or l . A few levels have been labeled above 9000 cm^{-1} , corresponding to the upper states of some of the experimentally observed transitions. We have related these labels to the quantum numbers usually used in the theoretical calculations, J , I , parity, and n . Note that our assignment of n , the index ordering levels with the same J , I , and parity by energy, is not necessarily final. We based n on the ordering of the energy levels in the calculations of Watson. It is possible, though unlikely, that for very closely spaced energy levels, the ordering of these levels may change in more accurate calculations, thus changing the assigned values of n . Also included in this table are the experimentally determined energy levels, which are the subject of Section III.3.

Our work is not the first attempt at labeling the energy levels. In 1994, Majewski *et al.* (which we refer to as Maj94 for the remainder of the paper) labeled each of the experimentally determined levels available at the time. This list was later expanded by Dinelli *et al.* (Din97) (33) to include many of the levels below 9000 cm^{-1} . The list of Din97 is not complete, however; 273 levels were left unlabeled. Of the roughly 720 levels that our work and Din97 have in common, 79 levels differ in assignment. Ten

TABLE 3
Energy Level Labels and Experimentally Determined Energy Levels

Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d	Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d	Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d			
J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.
1	p-1	64.123	64.121(00)*	(1,1)	00 ⁰	1	p-3	3240.678	3240.739(18)*	(1,1)	10 ⁰	10	p-3	4348.219	4348.435(64)	(10,1)	00 ⁰
1	o+1	86.959	86.960(00)*	(1,0)	00 ⁰	4	p-3	3260.197	3260.219(07)*	(4,2)	10 ¹	8	p+4	4371.260	4371.318(09)*	(8,7)	01 ¹
2	p+1	169.296	169.295(04)*	(2,2)	00 ⁰	1	o+2	3263.054	3263.145(16)	(1,0)	10 ⁰	6	p+7	4378.282	4378.380(10)*	(6,1)	01 ¹
2	p-1	237.350	237.356(05)*	(2,1)	00 ⁰	6	p+3	3269.582	3269.591(09)*	(6,7)	01 ¹	6	p-7	4389.279	4389.287(09)	(6,5)	10 ⁰
3	o-1	315.342	315.349(04)*	(3,3)	00 ⁰	5	p+3	3300.113	3300.141(08)*	(5,5)	01 ¹	5	p-8	4398.694		(5,1)	10 ⁰
3	p+1	428.009	428.018(07)*	(3,2)	00 ⁰	4	p+4	3326.091	3326.118(08)*	(4,1)	10 ¹	6	o-3	4400.982	4401.056(10)*	(6,0)	01 ¹
3	p-1	494.753	494.775(07)*	(3,1)	00 ⁰	9	p+2	3335.438	3335.559(19)*	(9,4)	00 ⁰	5	o+4	4419.137		(5,0)	10 ⁰
4	p+1	502.023	502.032(06)*	(4,4)	00 ⁰	2	p+4	3343.086	3343.147(14)*	(2,2)	10 ⁰	7	p-5	4420.158	4420.218(14)*	(7,4)	01 ¹
3	o+1	516.867	516.873(07)*	(3,0)	00 ⁰	4	p-4	3351.353	3351.385(08)*	(4,2)	10 ¹	7	p+5	4431.609	4431.693(08)*	(7,5)	01 ¹
4	o-1	658.698	658.720(06)*	(4,3)	00 ⁰	11	p+1	3352.780		(11,10)	00 ⁰	13	o+1	4449.473		(13,12)	00 ⁰
5	p-1	728.991	729.022(07)*	(5,5)	00 ⁰	5	p-3	3396.514	3396.519(09)*	(5,4)	10 ¹	7	p-6	4456.873	4456.867(09)*	(7,7)	10 ⁰
4	p+2	768.451	768.476(09)*	(4,2)	00 ⁰	12	o+1	3402.858		(12,12)	00 ⁰	14	p+1	4494.966		(14,14)	00 ⁰
4	p-1	833.555	833.583(08)*	(4,1)	00 ⁰	2	p-3	3409.771	3409.825(15)*	(2,1)	10 ⁰	10	p+5	4539.221		(10,11)	01 ¹
5	p+1	928.943	928.965(10)*	(5,4)	00 ⁰	4	p+5	3423.085	3423.125(08)*	(4,1)	10 ¹	11	o+1	4544.240	4544.410(22)	(11,6)	00 ⁰
6	o+1	995.842	995.884(05)*	(6,6)	00 ⁰	4	o-2	3447.011	3447.031(09)*	(4,0)	01 ¹	7	o+3	4562.728	4562.825(10)*	(7,3)	01 ¹
5	o-1	1080.453	1080.490(08)*	(5,3)	00 ⁰	9	o-2	3460.941	3461.058(17)*	(9,3)	00 ⁰	8	p+5	4567.212	4567.277(08)*	(8,7)	01 ¹
5	p+2	1187.067	1187.115(10)*	(5,2)	00 ⁰	10	p-1	3484.646	3484.761(16)*	(10,7)	00 ⁰	6	p+8	4575.977	4575.987(14)	(6,4)	10 ⁰
6	p-1	1238.409	1238.462(11)*	(6,5)	00 ⁰	3	o-3	3485.258	3485.306(12)*	(3,3)	10 ⁰	9	o+3	4605.661	4605.735(33)*	(9,9)	01 ¹
5	p-2	1250.267	1250.313(10)*	(5,1)	00 ⁰	5	p-4	3510.119	3510.142(07)*	(5,4)	10 ¹	12	o-1	4634.047		(12,9)	00 ⁰
5	o+1	1271.225	1271.245(10)*	(5,0)	00 ⁰	7	p-4	3530.235	3530.252(16)*	(7,8)	01 ¹	7	p-7	4635.928	4636.020(09)*	(7,4)	01 ¹
7	p-1	1302.095	1302.141(09)*	(7,7)	00 ⁰	5	o+2	3553.304	3553.333(09)*	(5,3)	10 ¹	8	o-2	4650.861	4650.945(08)*	(8,6)	01 ¹
6	p+1	1430.667	1430.706(11)*	(6,4)	00 ⁰	9	p+3	3555.305	3555.438(35)	(9,2)	00 ⁰	7	p-8	4663.773	4663.887(14)*	(7,2)	01 ¹
6	o-1	1577.279	1577.334(09)*	(6,3)	00 ⁰	6	o-2	3569.436	3569.467(07)*	(6,6)	01 ¹	6	o-4	4719.294	4719.259(11)	(6,3)	10 ⁰
7	o+1	1586.535	1586.594(08)*	(7,6)	00 ⁰	3	p+4	3595.694	3595.739(20)	(3,2)	10 ⁰	7	p+6	4720.296	4720.421(17)	(7,1)	01 ¹
8	p+1	1647.199	1647.267(12)	(8,8)	00 ⁰	9	p-3	3609.462	3609.643(52)	(9,1)	00 ⁰	7	o+4	4721.787	4721.794(07)	(7,6)	10 ⁰
6	p+2	1679.734	1679.805(14)*	(6,2)	00 ⁰	9	o+2	3627.453	3627.578(19)	(9,0)	00 ⁰	11	p-3	4733.919		(11,5)	00 ⁰
6	p-2	1740.834	1740.906(14)*	(6,1)	00 ⁰	5	p-5	3660.307	3660.348(10)*	(5,2)	10 ¹	7	o-4	4739.173	4739.271(18)	(7,0)	01 ¹
7	p-2	1818.077	1818.155(13)*	(7,5)	00 ⁰	3	p-5	3661.043	3661.081(21)	(3,1)	10 ⁰	9	p-5	4767.501	4767.585(11)	(9,8)	01 ¹
8	p-1	1972.727	1972.800(11)*	(8,7)	00 ⁰	4	p+6	3667.082	3667.126(16)*	(4,4)	10 ⁰	8	p+6	4774.975	4774.998(33)	(8,8)	10 ⁰
7	p+1	2002.387	2002.456(14)*	(7,4)	00 ⁰	5	o+3	3673.918	3673.958(06)*	(5,3)	10 ¹	7	o+5	4793.598	4793.695(09)*	(7,3)	01 ¹
9	o-1	2030.535	2030.623(13)*	(9,9)	00 ⁰	3	o+3	3682.683	3682.750(16)	(3,0)	10 ⁰	6	p+9	4818.401		(6,2)	10 ⁰
7	o-1	2142.025	2142.094(11)*	(7,3)	00 ⁰	6	p+4	3685.067	3685.094(10)*	(6,5)	10 ¹	1	p-4	4842.455	4842.607(71)	(1,1)	02 ⁰
7	p+2	2241.910	2241.999(20)*	(7,2)	00 ⁰	5	p+4	3722.593	3722.636(10)*	(5,1)	10 ¹	8	o-3	4862.697	4862.793(07)*	(8,6)	01 ¹
8	o+1	2242.117	2242.206(10)*	(8,6)	00 ⁰	11	o-1	3725.471	3725.625(19)	(11,9)	00 ⁰	1	o+3	4870.187	4870.309(08)*	(1,0)	02 ⁰
7	p-3	2300.773	2300.879(19)*	(7,1)	00 ⁰	10	o+1	3726.430	3726.566(16)*	(10,6)	00 ⁰	8	p+7	4874.318	4874.407(11)*	(8,5)	01 ¹
7	o+2	2320.309	2320.372(15)	(7,0)	00 ⁰	5	o-3	3743.140	3743.168(14)*	(5,0)	01 ¹	6	p-8	4877.837		(6,1)	10 ⁰
9	p+1	2396.323	2396.426(15)	(9,8)	00 ⁰	5	p-6	3792.977	3793.038(08)*	(5,2)	10 ¹	13	p-2	4879.901		(13,11)	00 ⁰
10	p+1	2451.425		(10,10)	00 ⁰	4	o-3	3820.769	3820.805(12)*	(4,3)	10 ⁰	11	p+3	4886.315	4886.494(31)	(11,4)	00 ⁰
8	p-2	2462.786	2462.889(15)*	(8,5)	00 ⁰	6	p+5	3825.386	3825.442(07)*	(6,5)	10 ¹	7	p-9	4891.925	4892.057(14)	(7,2)	01 ¹
0	p+1	2521.416	2521.411(05)*	(0,1)	01 ¹	8	o+2	3828.991	3829.019(13)*	(8,9)	01 ¹	12	p+2	4932.993		(12,8)	00 ⁰
1	p-2	2548.171	2548.164(11)*	(1,2)	01 ¹	5	p+5	3863.351	3863.417(09)*	(5,1)	10 ¹	2	p+5	4942.656	4942.720(15)	(2,2)	02 ⁰
1	p+1	2609.542	2609.541(05)*	(1,1)	01 ¹	7	p+3	3877.008	3877.036(10)*	(7,7)	01 ¹	11	o-2	4949.854		(11,12)	01 ¹
2	o+1	2614.279	2614.279(11)*	(2,3)	01 ¹	12	p-1	3884.031		(12,11)	00 ⁰	7	p+7	4961.582	4961.729(16)	(7,1)	01 ¹
1	o-1	2616.686	2616.684(05)*	(1,0)	01 ¹	6	p-3	3884.088	3884.117(10)*	(6,4)	10 ¹	7	p-10	4962.125	4962.118(11)	(7,5)	10 ⁰
8	p+2	2639.046	2639.135(17)*	(8,4)	00 ⁰	5	p-7	3888.654	3888.682(08)*	(5,5)	10 ⁰	9	p-6	4992.888	4992.978(13)	(9,8)	01 ¹
9	p-1	2701.980	2702.076(13)*	(9,7)	00 ⁰	10	p-2	3926.036	3926.180(23)	(10,5)	00 ⁰	1	o-2	4994.698	4994.833(08)*	(1,3)	02 ²
3	p-2	2719.486	2719.482(12)*	(3,4)	01 ¹	4	p+7	3928.143		(4,2)	10 ⁰	11	o-3	4994.803		(11,3)	00 ⁰
2	p-2	2723.958	2723.962(06)*	(2,2)	01 ¹	13	p-1	3931.766		(13,13)	00 ⁰	0	p+2	4997.920	4998.049(15)	(0,2)	02 ²
2	p+2	2755.565	2755.565(04)*	(2,1)	10 ¹	4	p-5	3991.803	3991.806(25)	(4,1)	10 ⁰	2	p-4	5023.366	5023.458(13)*	(2,1)	02 ⁰
8	o-1	2775.568	2775.667(13)*	(8,3)	00 ⁰	7	o-2	4010.200	4010.245(07)*	(7,6)	10 ¹	10	p-4	5026.026		(10,10)	01 ¹
2	p+3	2790.335	2790.344(04)*	(2,1)	10 ¹	6	o+2	4029.988	4030.048(09)*	(6,3)	10 ¹	8	p-5	5028.265	5028.395(12)*	(8,4)	01 ¹
2	o-1	2812.850	2812.857(05)*	(2,0)	01 ¹	6	p-4	4035.720	4035.770(08)*	(6,4)	10 ¹	2	p+6	5032.288	5032.393(07)*	(2,4)	02 ²
10	o-1	2856.600	2856.725(15)	(10,9)	00 ⁰	11	p+2	4044.000		(11,8)	00 ⁰	14	p-1	5048.185		(14,13)	00 ⁰
4	p+3	2863.938	2863.944(12)*	(4,5)	01 ¹	5	p+6	4084.701	4084.730(14)*	(5,4)	10 ⁰	3	o-4	5078.915	5078.930(09)*	(3,3)	02 ⁰
8	p+3	2868.766	2868.892(27)	(8,2)	00 ⁰	10	p+3	4086.290	4086.425(25)	(10,4)	00 ⁰	9	p+4	5086.222	5086.331(10)*	(9,7)	01 ¹
3	o+2	2876.835	2876.847(06)*	(3,3)	01 ¹	6	p-5	4129.260	4129.331(11)*	(6,2)	10 ¹	11	p+4	5087.285		(11,2)	00 ⁰
11	p-1	2909.130		(11,11)	00 ⁰	6	o+3	4147.034	4147.057(07)*	(6,6)	10 ⁰	1	p+2	5087.485	5087.617(08)*	(1,2)	02 ²
8	p-3	2925.302	2925.456(39)	(8,1)	00 ⁰	9	p-4	4165.459	4165.479(17)*	(9,10)	01 ¹	15	o-1	5091.529		(15,15)	00 ⁰
3	p-3	2931.365	2931.366(05)*	(3,2)	01 ¹	7	o-3	4177.864	4177.920(06)*	(7,6)	10 ¹	3	p-6	5105.206	5105.292(10)*	(3,5)	02 ²
9	o+1	2957.195	2957.306(13)*	(9,6)	00 ⁰	6	p+6	4188.726	4188.806(11)*	(6,1)	10 ¹	8	p+8	5107.141	5107.271(11)*	(8,5)	01 ¹
3	p-4	2992.421	2992.436(05)*	(3,2)	10 ¹	6	o+4	4202.235	4202.300(07)*	(6,3)	10 ¹	8	p-6	5109.777	5109.740(19)	(8,7)	10 ⁰
3	p+2	3002.888	3002.905(05)*	(3,1)	10 ¹ </												

TABLE 3—Continued

Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d	
J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.
3	p + 5	5210.738	5210.797(12)*	(3,2)	02 ⁰	4	o - 5	5810.924	5811.003(06)*	(4,3)	02 ²	6	o - 5	6301.641	6301.446(09)*	(6,3)	02 ⁰
4	o + 3	5215.696	5215.742(08)*	(4,6)	02 ²	2	p + 9	5815.659	5815.854(12)*	(2,1) <i>u</i>	11 ¹	13	o + 2	6304.993	6306.607	(13,6)	00 ⁰
4	p + 8	5251.799	5251.736(16)	(4,4)	02 ⁰	9	p - 8	5819.792	5819.799	(9,7)	10 ⁰	9	p + 10	6306.607	6306.854(34)	(9,1) <i>u</i>	01 ¹
13	p + 1	5253.446		(13,10)	00 ⁰	10	p - 6	5827.581	5827.721(13)	(10,8) <i>u</i>	01 ¹	9	p - 12	6310.290	6310.308(27)	(9,5)	10 ⁰
8	p - 7	5257.167	5257.344(29)	(8,2) <i>l</i>	01 ¹	5	o - 5	5830.536	5830.435(07)*	(5,3)	02 ⁰	7	o + 7	6312.444	6312.163(08)*	(7,6)	02 ⁰
2	p + 7	5266.301	5266.427(08)*	(2,2)	02 ²	2	o - 3	5835.140	5835.365(12)*	(2,0)	11 ¹	1	p - 7	6323.167		(1,1)	20 ⁰
7	o - 5	5269.950		(7,3)	10 ⁰	10	p + 7	5842.601	5842.715(11)*	(10,7) <i>l</i>	01 ¹	10	p + 10	6326.372		(10,8)	10 ⁰
3	p - 7	5282.255	5282.318(11)*	(3,1)	02 ⁰	9	p + 8	5842.747	5842.897(14)	(9,5) <i>u</i>	01 ¹	5	p + 12	6327.871	6327.954(06)*	(5,2) <i>u</i>	02 ²
2	o + 2	5286.801	5286.913(06)*	(2,0)	02 ²	4	p - 8	5846.716	5846.800(08)*	(4,1) <i>l</i>	02 ²	9	p - 13	6333.126	6332.831(16)	(9,11)	02 ²
3	p + 6	5299.131	5299.227(09)*	(3,4)	02 ²	12	o - 2	5856.687		(12,3)	00 ⁰	16	o - 1	6341.543		(16,15)	00 ⁰
8	p - 8	5304.760	5304.879(13)*	(8,4) <i>u</i>	01 ¹	13	p + 2	5858.836		(13,8)	00 ⁰	4	p + 14	6342.605	6342.581(23)	(4,1) <i>l</i>	11 ¹
2	p - 5	5304.836	5304.960(07)*	(2,1)	02 ²	13	p - 3	5880.157		(13,14)	01 ¹	1	o + 4	6345.170		(1,0)	20 ⁰
3	o + 4	5305.521	5305.584(09)*	(3,0)	02 ⁰	4	p + 12	5888.230	5888.310(08)*	(4,2) <i>u</i>	02 ²	5	p + 13	6346.262	6346.291(06)*	(5,5)	11 ¹
8	p + 9	5312.854	5313.058(40)	(8,1) <i>l</i>	01 ¹	8	o - 5	5895.174	5895.122(21)	(8,3)	10 ⁰	11	o + 4	6359.874	6360.031(18)	(11,9) <i>u</i>	01 ¹
9	p + 5	5328.204	5328.318(10)*	(9,7) <i>u</i>	01 ¹	6	p - 9	5895.861	5895.803(08)*	(6,7)	02 ²	4	p - 12	6363.351	6363.417(22)	(4,2) <i>u</i>	11 ¹
9	o - 4	5341.999	5342.110(09)*	(9,6) <i>l</i>	01 ¹	4	o + 4	5896.787	5896.838(08)*	(4,0)	02 ²	11	p - 8	6373.929	6374.075(19)	(11,8) <i>l</i>	01 ¹
8	o + 4	5361.226	5361.203(12)	(8,6)	10 ⁰	5	p - 11	5899.394	5899.405(07)*	(5,5)	02 ²	5	p - 14	6376.427	6376.531(13)*	(5,1) <i>u</i>	02 ²
5	p - 9	5363.833	5363.825(09)*	(5,7)	02 ²	14	p - 2	5900.751		(14,11)	00 ⁰	17	p - 1	6380.839		(17,17)	00 ⁰
7	p + 9	5368.025		(7,2)	10 ⁰	9	p - 9	5908.498	5908.688(37)	(9,2) <i>l</i>	01 ¹	5	o + 7	6391.743	6391.860(08)*	(5,0)	02 ²
12	p + 3	5396.949		(12,13)	01 ¹	3	o + 6	5909.950	5910.110(06)*	(3,3)	11 ¹	6	p + 13	6395.046	6394.877(10)	(6,2)	02 ⁰
12	o + 2	5406.041		(12,6)	00 ⁰	4	p + 13	5920.770	5920.863(09)*	(4,5)	11 ¹	14	o + 2	6399.234		(14,15)	01 ¹
7	p - 11	5424.988		(7,1)	10 ⁰	12	p + 5	5922.883		(12,2)	00 ⁰	10	p - 7	6400.921	6401.106(27)	(10,4) <i>l</i>	01 ¹
3	o - 5	5431.017	5431.122(06)*	(3,3)	02 ²	4	p - 9	5931.782	5931.881(06)*	(4,1) <i>u</i>	02 ²	6	p + 14	6403.624	6403.513(08)*	(6,4) <i>l</i>	02 ²
4	o - 4	5434.341	5434.331(12)	(4,3)	02 ⁰	5	p + 9	5939.804	5939.707(12)	(5,2)	02 ⁰	5	p - 15	6410.567	6410.544(19)	(5,4) <i>l</i>	11 ¹
7	o + 6	5443.899		(7,0)	10 ⁰	10	o - 3	5944.621		(10,9)	10 ⁰	10	o - 5	6412.156	6412.314(12)	(10,6) <i>u</i>	01 ¹
10	o + 3	5454.327	5454.430(11)	(10,9) <i>u</i>	01 ¹	3	p - 11	5949.328	5949.443(17)	(3,2) <i>l</i>	11 ¹	6	p - 11	6415.774	6415.757(07)*	(6,5)	02 ²
4	p - 6	5460.400	5460.463(10)*	(4,5)	02 ²	11	p - 6	5950.826		(11,10) <i>u</i>	01 ¹	2	p + 10	6422.880		(2,2)	20 ⁰
5	p - 10	5460.611	5460.464(24)	(5,5)	02 ⁰	9	p + 9	5962.005		(9,1) <i>l</i>	01 ¹	11	p + 6	6429.599		(11,10)	10 ⁰
8	o + 5	5463.022	5463.104(10)*	(8,3) <i>u</i>	01 ¹	12	p - 4	5969.587		(12,1)	00 ⁰	4	p + 15	6430.890	6430.944(23)	(4,1) <i>u</i>	11 ¹
11	p + 5	5483.252		(11,11)	01 ¹	5	o - 6	5971.155	5971.228(08)*	(5,3) <i>l</i>	02 ²	9	p + 11	6449.215	6449.198(39)	(9,4)	10 ⁰
3	p - 8	5486.357	5486.457(06)*	(3,1) <i>l</i>	02 ²	12	o - 3	5976.967		(12,12)	01 ¹	7	p - 13	6451.317	6451.126(14)*	(7,8)	11 ¹
9	p + 6	5487.233	5487.329(18)	(9,8)	10 ⁰	9	o - 6	5979.029	5979.217(21)	(9,0)	01 ¹	4	o - 6	6453.613	6453.690(19)	(4,0)	11 ¹
14	o + 1	5502.846		(14,12)	00 ⁰	8	p + 12	5981.430		(8,2)	10 ⁰	6	p - 12	6461.276		(6,1)	02 ⁰
8	p - 9	5532.618	5532.751(20)	(8,2) <i>u</i>	01 ¹	6	p - 10	5984.075	5983.896(13)	(6,5)	02 ⁰	13	p - 5	6476.011		(13,5)	00 ⁰
3	p + 7	5533.626	5533.730(06)*	(3,2)	02 ²	7	p - 12	5985.536	5985.149(13)	(7,7)	02 ⁰	12	p + 7	6481.539		(12,11) <i>u</i>	01 ¹
4	p + 9	5544.226	5544.213(08)*	(4,2)	02 ⁰	5	p - 12	6003.275	6003.183(14)	(5,1)	02 ⁰	8	o - 6	6482.308	6482.118(19)*	(8,9)	02 ²
6	p + 10	5549.695	5549.624(11)*	(6,8)	02 ²	11	p - 7	6003.418		(11,11)	10 ⁰	12	o + 3	6483.120		(12,12)	10 ⁰
0	p + 3	5554.029		(0,1)	11 ¹	3	p - 12	6015.800	6015.946(17)	(3,2) <i>u</i>	11 ¹	2	p - 7	6488.458		(2,1)	20 ⁰
10	p - 5	5555.295	5555.440(16)	(10,8) <i>l</i>	01 ¹	5	o + 6	6023.187	6023.081(17)	(5,0)	02 ⁰	7	p - 14	6505.291	6505.157(08)*	(7,7)	02 ²
10	p + 6	5558.547		(10,10)	10 ⁰	3	p + 8	6023.657	6023.757(18)	(3,1) <i>l</i>	11 ¹	13	p + 3	6506.815		(13,13)	01 ¹
9	p + 7	5565.214	5565.255(12)*	(9,5) <i>l</i>	01 ¹	9	p - 10	6031.544	6031.681(15)	(9,4) <i>u</i>	01 ¹	6	o - 6	6516.118	6516.152(09)*	(6,3) <i>l</i>	02 ²
3	o + 5	5567.276	5567.389(07)*	(3,0)	02 ²	8	p + 13	6034.410	6034.182(13)*	(8,10)	02 ²	5	p - 16	6529.268	6529.276(11)*	(5,4) <i>u</i>	11 ¹
3	p - 9	5573.651	5573.764(05)*	(3,1) <i>u</i>	02 ²	8	p - 11	6035.672		(8,1)	10 ⁰	10	o + 4	6539.707	6539.950(14)	(10,3) <i>l</i>	01 ¹
13	o - 1	5577.736		(13,9)	00 ⁰	3	o - 6	6047.437	6047.564(19)	(3,0)	11 ¹	14	o - 1	6552.138		(14,9)	00 ⁰
1	p - 6	5584.000	5584.224(10)*	(1,2)	11 ¹	9	o + 5	6053.084	6053.096(14)*	(9,6)	10 ⁰	9	o - 7	6559.077		(9,3)	10 ⁰
12	p - 3	5585.628		(12,5)	00 ⁰	11	o + 3	6057.273	6057.448(13)	(11,9) <i>l</i>	01 ¹	3	o - 7	6561.242		(3,3)	20 ⁰
8	p - 10	5604.263	5604.254(22)	(8,5)	10 ⁰	3	p + 9	6080.829	6080.967(18)	(3,1) <i>u</i>	11 ¹	5	o + 8	6568.335	6568.247(10)*	(5,3) <i>l</i>	11 ¹
8	p + 10	5606.619	5606.814(20)	(8,1) <i>u</i>	01 ¹	10	o - 4	6087.435	6087.522(11)*	(10,6) <i>l</i>	01 ¹	7	p - 15	6571.908		(7,5)	02 ⁰
9	o - 5	5610.185	5610.323(10)*	(9,6) <i>u</i>	01 ¹	5	p + 10	6089.800	6089.815(06)*	(5,4)	02 ²	15	o + 1	6574.418		(15,12)	00 ⁰
4	p - 7	5610.453	5610.451(13)	(4,1)	02 ⁰	13	p - 4	6100.528		(13,7)	00 ⁰	10	p - 8	6579.738		(10,7)	10 ⁰
8	o - 4	5628.911	5629.057(13)	(8,0)	01 ¹	4	p - 10	6105.534	6105.639(06)*	(4,4)	11 ¹	12	p - 5	6591.347		(12,10) <i>l</i>	01 ¹
1	p + 3	5640.267	5640.488(15)*	(1,1)	11 ¹	5	o - 7	6129.541	6129.539(07)*	(5,6)	11 ¹	6	p + 15	6608.128	6608.127(07)*	(6,4) <i>u</i>	02 ²
1	o - 3	5644.521	5644.739(15)	(1,0)	11 ¹	6	p + 11	6141.434	6141.238(13)	(6,4)	02 ⁰	10	p - 9	6612.267		(10,2) <i>l</i>	01 ¹
4	p + 10	5652.415	5652.479(07)*	(4,4)	02 ²	10	p + 8	6145.084	6145.224(14)	(10,7) <i>u</i>	01 ¹	13	p + 4	6612.793		(13,4)	00 ⁰
2	o + 3	5653.801	5654.004(06)*	(2,3)	11 ¹	15	p - 1	6154.036		(15,13)	00 ⁰	10	p + 11	6628.474	6628.649(18)	(10,5) <i>u</i>	01 ¹
5	o + 5	5659.211	5659.227(07)*	(5,6)	02 ²	4	o + 5	6158.228	6158.271(15)*	(4,3) <i>l</i>	11 ¹	6	o - 7	6639.008	6638.903(05)*	(6,6)	11 ¹
11	p - 5	5662.198		(11,10) <i>l</i>	01 ¹	12	p + 6	6158.807		(12,11) <i>l</i>	01 ¹	11	p + 7	6644.910	6644.997(17)	(11,7) <i>l</i>	01 ¹
15	p + 1	5679.206		(15,14)	00 ⁰	5	p + 11	6169.392	6169.455(08)*	(5,2) <i>l</i>	02 ²	6	p + 16	6650.995	6650.933(10)*	(6,5) <i>l</i>	11 ¹
9	p - 7	5689.513	5689.686(14)*	(9,4) <i>l</i>	01 ¹	7	p + 10	6170.176	6170.055(11)*	(7,8)	02 ²	9	o - 8	6651.211	6650.536(16)* [†]	(9,9)	02 ⁰
5	p + 8	5690.933	5690.831(12)	(5,4)	02 ⁰	9	o + 6	6175.031	6175.176(14)	(9,3) <i>u</i>	01 ¹	9	p + 12	6654.121		(9,2)	10 ⁰
6	o + 5																

TABLE 3—Continued

Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d	
J	I P N	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P N	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P N	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.
11	p-9	6710.274	6710.449(19)	(11,8) <i>u</i>	01 ¹	1	o-4	7082.860		(1,0)	03 ¹	3	p+11	7460.169		(3,1) <i>l</i>	03 ¹
9	o+7	6722.382		(9,0)	10 ⁰	1	p+4	7102.718	7103.087(70)*	(1,1)	03 ¹	7	p+19	7462.158	7462.319(12)	(7,2) <i>u</i>	02 ²
6	p+17	6724.776	6724.704(11)*	(6,2) <i>l</i>	02 ²	9	p-15	7105.227	7104.956(13)*	(9,10)	11 ¹	8	o+9	7463.737	7463.686(08)*	(8,6) <i>u</i>	02 ²
5	p+14	6733.406		(5,1) <i>l</i>	11 ¹	8	o+8	7119.292	7119.112(07)*	(8,6) <i>l</i>	02 ²	11	p+12	7472.560	7472.755(24)	(11,5) <i>u</i>	01 ¹
3	p-13	6734.090		(3,1)	20 ⁰	2	o+4	7122.615	7122.646(70)*	(2,3)	03 ¹	10	p+17	7474.940		(10,11)	11 ¹
7	p-16	6736.644	6736.501(11)	(7,5) <i>l</i>	02 ²	7	o-9	7126.799	7126.714(08)*	(7,3) <i>l</i>	02 ²	5	o+10	7475.145		(5,0)	20 ⁰
4	p+16	6738.201		(4,4)	20 ⁰	6	p-17	7137.200		(6,2) <i>l</i>	11 ¹	8	p+18	7478.072		(8,2)	02 ⁰
5	o-9	6753.295		(5,0)	11 ¹	11	p+10	7143.390		(11,8)	10 ⁰	9	o+8	7487.746		(9,6)	02 ⁰
3	o+7	6755.339		(3,0)	20 ⁰	5	p+16	7147.109		(5,4)	20 ⁰	0	o+1	7492.558	7492.912(13) [†]	(0,3)	03 ³
11	o-4	6765.378	6765.479(18)	(11,9)	10 ⁰	13	p+6	7156.697		(13,11) <i>l</i>	01 ¹	12	p+10	7494.243		(12,7) <i>l</i>	01 ¹
8	o+6	6766.598	6766.387(12)	(8,9)	11 ¹	11	p-11	7157.738		(11,4) <i>l</i>	01 ¹	13	o+4	7496.871		(13,12)	10 ⁰
6	p-13	6768.699	6768.675(14)*	(6,1) <i>l</i>	02 ²	6	o+9	7184.268	7184.132(24)	(6,3) <i>u</i>	11 ¹	3	o-9	7498.146		(3,0)	03 ¹
6	o-8	6775.243	6775.358(05)*	(6,3) <i>u</i>	02 ²	8	p-14	7187.096	7186.960(11)	(8,7)	02 ²	7	p-21	7499.704		(7,7)	20 ⁰
7	o+8	6784.144	6784.078(06)*	(7,6)	02 ²	10	p+15	7191.871		(10,4)	10 ⁰	7	o+10	7504.956	7504.928(11)*	(7,0)	02 ²
5	p-18	6792.527		(5,2) <i>u</i>	11 ¹	6	p+20	7192.784		(6,1) <i>l</i>	11 ¹	10	o-8	7507.678	7507.250(34)	(10,9)	02 ⁰
13	p+5	6798.691		(13,2)	00 ⁰	7	p+15	7193.350	7193.266(12)	(7,5) <i>l</i>	11 ¹	2	p-9	7514.456		(2,4)	03 ³
7	o-7	6801.893	6801.634(14)	(7,3)	02 ⁰	2	p+11	7208.359	7208.334(70)*	(2,1) <i>l</i>	03 ¹	4	p+19	7514.707		(4,7)	03 ³
6	o+7	6803.721	6803.674(08)*	(6,0)	02 ²	7	o-10	7209.275	7209.165(07)*	(7,6) <i>u</i>	11 ¹	13	p+7	7518.884		(13,11) <i>u</i>	01 ¹
10	o+6	6804.407	6804.385(17)	(10,6)	10 ⁰	6	o+10	7214.088		(6,6)	20 ⁰	13	p-8	7522.856		(13,10) <i>l</i>	01 ¹
10	p-10	6811.587	6811.748(17)	(10,4) <i>u</i>	01 ¹	7	p+16	7215.802	7215.768(09)*	(7,4) <i>u</i>	02 ²	3	p-16	7525.632		(3,2) <i>u</i>	03 ¹
14	p+3	6816.687		(14,8)	00 ⁰	10	p-13	7220.112		(10,11)	02 ²	14	p+5	7528.745		(14,4)	00 ⁰
16	p+2	6831.384		(16,14)	00 ⁰	3	p-14	7230.283		(3,4)	03 ¹	7	p-22	7529.091	7529.104(12)*	(7,1) <i>u</i>	02 ²
9	p+13	6832.272	6832.012(23)*	(9,10)	02 ²	12	p-8	7231.551		(12,8) <i>l</i>	01 ¹	17	o-1	7532.405		(17,15)	00 ⁰
13	p-6	6842.607		(13,1)	00 ⁰	2	p-8	7235.508	7235.742(70)*	(2,2)	03 ¹	8	p-17	7533.994		(8,1)	02 ⁰
6	p+18	6842.645	6842.628(11)*	(6,5) <i>u</i>	11 ¹	8	p-15	7237.698		(8,8)	11 ¹	9	p-17	7543.585	7543.360(08)	(9,7) <i>l</i>	02 ²
13	o+3	6857.353		(13,0)	00 ⁰	14	o+3	7240.093		(14,6)	00 ⁰	14	p+6	7543.859		(14,14)	10 ⁰
6	p+19	6858.525	6858.614(10)*	(6,2) <i>u</i>	02 ²	8	p+16	7245.339	7244.972(17)	(8,4)	02 ⁰	16	p+3	7544.997		(16,17)	01 ¹
5	p+15	6859.884	6859.846(27)	(5,1) <i>u</i>	11 ¹	14	p+4	7245.506		(14,13) <i>l</i>	01 ¹	8	p+19	7550.745	7550.536(10)	(8,7) <i>u</i>	11 ¹
8	p+15	6862.974	6862.763(11)*	(8,8)	02 ²	11	o+5	7255.270		(11,3) <i>l</i>	01 ¹	4	o+7	7551.194	7550.316(17)*	(4,3) <i>l</i>	03 ¹
7	p+12	6863.533	6863.442(14)	(7,7)	11 ¹	9	o-9	7256.963	7256.689(34)*	(9,9)	02 ²	15	o-2	7552.649		(15,9)	00 ⁰
6	p-14	6885.871	6885.838(07)*	(6,4) <i>l</i>	11 ¹	11	o-6	7257.663	7257.845(16)	(11,6) <i>u</i>	01 ¹	5	o-11	7554.025		(5,6)	03 ¹
11	o-5	6889.006	6888.997(16)	(11,6) <i>l</i>	01 ¹	15	p+2	7267.339		(15,10)	00 ⁰	1	p-10	7571.716		(1,2)	03 ³
4	o-7	6889.507		(4,3)	20 ⁰	16	p-1	7272.831		(16,13)	00 ⁰	11	o+6	7592.288	7592.299(21)	(11,6)	10 ⁰
7	p+13	6910.479		(7,2)	02 ⁰	5	o-10	7292.388		(5,3)	20 ⁰	7	o+11	7596.047	7595.852(11)*	(7,3) <i>l</i>	11 ¹
12	p-6	6924.053		(12,10) <i>u</i>	01 ¹	6	p-18	7295.826		(6,2) <i>u</i>	11 ¹	3	p+12	7597.008		(3,1) <i>u</i>	03 ¹
6	p-15	6929.171	6929.194(10)*	(6,1) <i>u</i>	02 ²	2	p+12	7301.181	7301.423(50)*	(2,1) <i>u</i>	03 ¹	4	p-14	7598.521		(4,4)	03 ¹
12	o+4	6934.573	6934.696(19)	(12,9) <i>l</i>	01 ¹	12	o+5	7304.749	7304.925(24)	(12,9) <i>u</i>	01 ¹	12	o-4	7607.102	7607.488(24)	(12,6) <i>l</i>	11 ¹
15	p-2	6942.487		(15,11)	00 ⁰	10	o-7	7316.361		(10,3)	10 ⁰	7	p-23	7618.630	7618.564(17)	(7,4) <i>u</i>	11 ¹
8	o+7	6942.604	6942.162(31)	(8,6)	02 ⁰	7	p+17	7317.914	7317.753(12)	(7,2) <i>l</i>	02 ²	8	o-8	7620.729	7620.408(12)	(8,6) <i>l</i>	11 ¹
12	p-7	6946.958		(12,11)	10 ⁰	7	p-19	7318.407	7318.322(15)	(7,1) <i>l</i>	02 ²	9	p+15	7622.656	7622.455(12)	(9,8)	02 ²
5	p-19	6953.844		(5,5)	20 ⁰	12	p+8	7319.294		(12,10)	10 ⁰	6	p+22	7628.078		(6,4)	20 ⁰
15	p-3	6954.114		(15,16)	01 ¹	1	p-9	7325.092		(1,4)	03 ³	12	p-9	7632.355		(12,8) <i>u</i>	01 ¹
10	o+7	6958.960	6959.029(17)	(10,3) <i>u</i>	01 ¹	2	o-4	7327.984	7328.209(18)*	(2,0)	03 ¹	11	p-15	7638.575	7638.738(29)	(11,4) <i>u</i>	01 ¹
7	p-17	6961.325		(7,1)	02 ⁰	7	o-11	7340.062	7340.114(08)*	(7,3) <i>u</i>	02 ²	5	p-21	7639.368		(5,8)	03 ³
10	p-11	6967.179	6967.296(36)	(10,5)	10 ⁰	9	p-16	7349.528	7348.806(13) [†]	(9,7)	02 ⁰	14	p+7	7641.001		(14,13) <i>u</i>	01 ¹
7	o-8	6985.302	6985.098(12)	(7,6) <i>l</i>	11 ¹	8	p+17	7352.822	7352.467(13)	(8,7) <i>l</i>	11 ¹	11	o+7	7645.892		(11,12)	02 ²
7	o+9	6990.002	6989.796(20)	(7,0)	02 ⁰	11	p-12	7357.593		(11,2) <i>l</i>	01 ¹	9	o+9	7651.520	7651.050(13)	(9,9)	11 ¹
4	p+17	6994.960		(4,2)	20 ⁰	6	p+21	7361.961		(6,1) <i>u</i>	11 ¹	8	p+20	7656.631	7656.488(15)*	(8,4) <i>l</i>	02 ²
13	p-7	6996.883		(13,13)	10 ⁰	3	p-15	7362.607	7362.203(70)*	(3,2) <i>l</i>	03 ¹	16	o+1	7658.990		(16,12)	00 ⁰
11	p+8	6999.062		(11,5) <i>l</i>	01 ¹	2	p+13	7368.759		(2,5)	03 ³	3	p+13	7659.448		(3,5)	03 ³
7	p+14	7002.735	7002.630(09)	(7,4) <i>l</i>	02 ²	4	p+18	7374.555		(4,5)	03 ¹	14	o-2	7662.640		(14,3)	00 ⁰
0	p+4	7005.822		(0,1)	03 ¹	1	o+5	7380.868		(1,3)	03 ³	7	p-24	7667.810		(7,2) <i>l</i>	11 ¹
11	p+9	7008.607	7008.759(17)	(11,7) <i>u</i>	01 ¹	10	p+16	7381.747		(10,2)	10 ⁰	15	o+2	7673.607		(15,15)	01 ¹
7	p-18	7027.151	7027.114(09)	(7,5) <i>u</i>	02 ²	6	o-9	7383.902		(6,0)	11 ¹	9	p-18	7676.894	7676.650(34)	(9,8) <i>u</i>	11 ¹
17	p+1	7034.173		(17,16)	00 ⁰	11	p-13	7391.054		(11,7)	10 ⁰	10	p+18	7686.921	7686.522(17)	(10,10)	02 ²
6	p-16	7034.347		(6,4) <i>u</i>	11 ¹	3	o+8	7394.164		(3,3)	03 ¹	8	p-18	7697.713	7697.709(10)	(8,5) <i>u</i>	02 ²
10	p+13	7035.883	7035.014(19)	(10,10)	02 ⁰	5	p+17	7394.743		(5,2)	20 ⁰	4	p-15	7701.393			

TABLE 3—Continued

Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d		Q. N. ^a		E _{calc} ^b	E _{exp} ^c	Label ^d	
J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.	J	I P n	(cm ⁻¹)	(cm ⁻¹)	Rot.	Vib.
6	o-10	7768.674		(6,3)	20 ⁰	3	p-19	8017.654		(3,2)	03 ³	6	p-22	8275.953		(6,4)l	03 ¹
6	p+23	7769.017		(6,7)	03 ¹	7	p-28	8019.418		(7,8)	03 ¹	5	o-13	8277.627	8277.033(09)*	(5,6)	03 ³
5	p-22	7769.343	7767.914(70)*	(5,4)l	03 ¹	6	p+25	8020.895		(6,5)l	03 ¹	13	o+6	8281.674		(13,9)u	01 ¹
4	p+20	7773.696		(4,1)l	03 ¹	4	o-9	8031.115	8030.925(13)*	(4,0)	03 ¹	8	p-24	8288.656	8288.481(16)	(8,4)u	11 ¹
9	p-19	7778.214	7777.748(34)	(9,5)	02 ⁰	4	p+22	8036.056		(4,5)	03 ³	9	o-12	8294.533		(9,6)l	11 ¹
8	p+21	7785.048		(8,5)l	11 ¹	8	o-11	8043.282	8043.489(10)	(8,3)u	02 ²	9	p+19	8296.092		(9,7)u	11 ¹
4	o-8	7786.995	7786.722(09)*	(4,6)	03 ³	8	p-20	8045.233		(8,4)l	11 ¹	3	o-12	8302.131	8302.108(12)*	(3,3)	12 ²
10	p+19	7787.226	7786.171(12)	(10,8)	02 ⁰	5	p-23	8054.048		(5,4)u	03 ¹	8	p+28	8303.162		(8,1)l	11 ¹
19	p-1	7791.596		(19,19)	00 ⁰	9	p+17	8055.975		(9,7)l	11 ¹	7	o-13	8303.281	8300.927(18)†	(7,6)l	03 ¹
4	o+8	7795.070	7794.757(13)	(4,3)u	03 ¹	2	o-5	8057.083	8057.354(21)*	(2,3)	12 ²	8	o+12	8305.143		(8,9)	03 ¹
3	p-17	7796.561		(3,4)	03 ³	11	o-9	8058.834		(11,3)	10 ⁰	7	o-14	8306.730		(7,3)	20 ⁰
7	o+13	7797.199	7796.716(15)†	(7,6)	20 ⁰	8	p+25	8071.099		(8,5)u	11 ¹	16	p+4	8306.809		(16,10)	00 ⁰
6	o+11	7798.639		(6,9)	03 ³	4	p+23	8074.027		(4,1)l	03 ³	16	p-3	8309.933		(16,16)	01 ¹
15	p+3	7801.887		(15,8)	00 ⁰	14	p-7	8078.755		(14,13)	10 ⁰	12	p+15	8315.622		(12,13)	11 ¹
12	p+11	7808.089		(12,5)l	01 ¹	5	p-24	8090.997		(5,2)l	03 ¹	4	p+24	8332.152		(4,1)u	03 ³
8	o-9	7822.890	7822.667(08)*	(8,3)l	02 ²	10	o-9	8092.072	8091.784(33)	(10,9)	02 ²	3	p-21	8335.467	8335.280(21)	(3,1)l	12 ²
5	p+18	7831.448		(5,5)	03 ¹	6	o-11	8099.519		(6,6)	03 ¹	4	p-18	8340.205	8340.064(12)	(4,5)	12 ²
15	p-4	7833.506		(15,14)l	01 ¹	12	o+7	8100.211		(12,3)l	01 ¹	12	p+16	8348.569		(12,5)u	01 ¹
8	p+22	7837.460		(8,8)	20 ⁰	9	p+18	8104.543		(9,2)	02 ⁰	15	p-7	8352.686		(15,5)	00 ⁰
1	p-11	7839.758		(1,1)	12 ⁰	4	o+9	8105.366	8105.227(11)	(4,6)	12 ²	9	p+20	8359.938	8359.769(13)	(9,4)l	02 ²
13	o+5	7844.719		(13,9)l	01 ¹	10	o+8	8108.521		(10,9)l	11 ¹	4	p+25	8365.683		(4,4)	12 ⁰
11	p+13	7846.486		(11,1)u	01 ¹	10	p-16	8109.690		(10,7)	02 ⁰	4	o+10	8365.752	8365.478(08)*†	(4,3)	03 ³
8	o-10	7851.386	7851.267(11)	(8,6)u	11 ¹	12	p-11	8109.750		(12,13)	02 ²	4	p-19	8366.382	8366.107(13)†	(4,2)	03 ³
3	o+9	7854.403		(3,3)	03 ³	15	o-3	8123.151		(15,15)	10 ⁰	8	o+13	8370.655		(8,6)	20 ⁰
1	o+6	7857.588		(1,0)	12 ⁰	13	p-11	8128.161		(13,8)l	01 ¹	6	p+26	8377.251		(6,5)u	03 ¹
6	p+24	7865.379		(6,2)	20 ⁰	2	p+17	8135.537	8135.727(11)*	(2,2)	12 ²	15	p+4	8378.098		(15,13)l	01 ¹
7	p-25	7866.075		(7,2)u	11 ¹	9	p-22	8135.925	8135.743(12)	(9,5)l	02 ²	11	p+17	8380.507		(11,11)	11 ¹
3	o-10	7866.482	7866.300(07)*	(3,0)	03 ³	14	p+9	8136.791		(14,11)l	01 ¹	10	p-17	8380.721		(10,5)	02 ⁰
0	p+5	7869.974	7870.015(10)†	(0,2)	12 ²	5	o-12	8138.695	8137.585(71)*	(5,0)	03 ¹	14	p+10	8391.253		(14,16)	02 ²
1	o-5	7872.300	7872.661(10)	(1,3)	12 ²	3	o-11	8139.528	8139.068(12)	(3,3)	12 ⁰	13	p+9	8391.710		(13,7)l	01 ¹
9	p+16	7875.817		(9,4)	02 ⁰	8	p+26	8139.636	8139.751(15)	(8,2)u	02 ²	13	p-12	8394.156		(13,13)	02 ⁰
11	o-8	7878.132		(11,12)	11 ¹	2	o+6	8142.021	8142.088(11)	(2,0)	12 ²	10	p-18	8395.134	8394.981(14)	(10,7)l	02 ²
11	p-17	7882.766		(11,2)u	01 ¹	14	o-4	8142.969		(14,12)u	01 ¹	17	p-2	8395.571		(17,13)	00 ⁰
8	p+23	7883.920	7883.910(10)	(8,4)u	02 ²	8	p-21	8143.854		(8,7)	20 ⁰	5	p-27	8397.251		(5,2)u	03 ¹
9	o+10	7891.450	7891.333(07)*	(9,6)l	02 ²	9	o+11	8145.964	8145.790(09)	(9,0)	02 ⁰	3	p+17	8400.644	8400.492(12)*	(3,2)	12 ²
13	p-9	7897.671		(13,11)	10 ⁰	12	o-6	8149.413		(12,6)u	01 ¹	7	p+23	8401.355		(7,7)	03 ¹
13	o-5	7903.526		(13,15)	02 ²	12	p-12	8151.257		(12,2)l	01 ¹	7	p+24	8402.733		(7,2)	20 ⁰
12	o+6	7907.409		(12,12)	02 ⁰	11	p-18	8152.184		(11,11)	02 ²	5	p+21	8405.601		(5,1)l	03 ³
12	p+12	7911.595		(12,7)u	01 ¹	5	p+20	8153.168		(5,1)l	03 ¹	6	o+12	8413.795		(6,3)l	03 ¹
2	p+15	7914.800	7915.081(10)	(2,4)	12 ¹	11	p+16	8160.296		(11,2)	10 ⁰	3	o+11	8425.544	8425.436(16)*	(3,0)	12 ²
4	p-16	7915.419	7915.179(16)	(4,2)u	03 ¹	4	p-17	8167.805		(4,4)	03 ³	7	o+14	8432.257		(7,9)	03 ³
10	p-15	7921.531		(10,10)	11 ¹	2	p-12	8168.050	8168.185(11)	(2,1)	12 ²	12	o+8	8435.511		(12,6)	10 ⁰
8	p-19	7921.988	7921.807(17)	(8,1)l	02 ²	9	o-11	8169.582		(9,9)	20 ⁰	3	p-22	8435.637	8435.428(12)	(3,1)u	12 ²
6	p-20	7923.430		(6,1)	20 ⁰	17	o-2	8172.423		(17,18)	01 ¹	9	p-24	8438.879	8438.843(12)	(9,5)u	02 ²
13	p-10	7927.073		(13,10)u	01 ¹	9	p-23	8176.283	8176.101(12)	(9,1)	02 ⁰	10	p+21	8443.972	8443.688(14)	(10,8)u	02 ²
7	p+21	7931.270		(7,1)u	11 ¹	3	p+15	8176.849	8176.975(11)	(3,4)	12 ²	9	p+21	8444.698		(9,5)l	11 ¹
9	p-20	7935.986	7935.837(09)*	(9,7)u	02 ²	7	p+22	8176.919		(7,4)	20 ⁰	10	o+10	8445.678		(10,9)u	11 ¹
12	p-10	7954.901		(12,4)l	01 ¹	6	p-21	8181.377		(6,8)	03 ³	6	p+27	8447.143		(6,8)	12 ²
8	p+24	7956.627	7956.409(15)	(8,2)l	02 ²	8	p-22	8182.966	8182.943(16)	(8,1)u	02 ²	8	o+14	8447.999		(8,3)u	11 ¹
1	p+5	7958.502	7958.833(10)	(1,2)	12 ²	8	o+11	8183.324	8182.737(20)	(8,3)l	11 ¹	16	o+2	8449.967		(16,15)l	01 ¹
8	o+10	7959.597	7959.452(19)	(8,0)	02 ²	12	p+14	8197.456		(12,1)l	01 ¹	7	p-29	8458.186		(7,1)	20 ⁰
2	p+16	7963.525		(2,2)	12 ⁰	15	o+3	8197.917		(15,6)	00 ⁰	13	o-6	8468.234		(13,6)l	01 ¹
5	p+19	7964.453		(5,7)	03 ³	11	p-19	8203.419		(11,1)	10 ⁰	15	p+5	8472.890		(15,4)	00 ⁰
5	o+11	7964.573	7963.581(70)*	(5,3)l	03 ¹	11	o+9	8218.112		(11,0)	10 ⁰	7	o+15	8476.591		(7,0)	20 ⁰
4	p+21	7966.380		(4,1)u	03 ¹	3	p+16	8221.093		(3,2)	12 ⁰	12	p-14	8477.374		(12,11)	02 ⁰
11	p+14	7975.338		(11,10)	02 ⁰	13	p+8	8224.778		(13,10)	10 ⁰	6	p+28	8482.606		(6,7)	03 ³
3	p+14	7977.777		(3,1)	03 ³	8	p+27	8225.035		(8,11)	03 ³	14	p-8	8482.722		(14,10)l	01 ¹
9	p-21	7980.439		(9,8)l	11 ¹	5	o+12	8230.298	8229.545(15)	(5,3)u	03 ¹	4	o-11	8484.009		(4,3)	12 ⁰
9	o-10	7984.334		(9,3)	02 ⁰	12	p-13	8237.880		(12,7)	10 ⁰	5	p+22	8485.982		(5,5)	03 ³
11	p+15	7984.584		(11,4)	10 ⁰	9	o+12	8237.934	8237.796(09)	(9,6)u	02 ²	0	p+6	8488.013		(0,1)	21 ¹
1	p-12	7989.277	7989.534(11)	(1,1)	12 ²	4	o-10	8247.865		(4,0)	03 ³	9	o-13	8492.697		(9,12)	03 ³
3	p-18	7991.546	7991.670(11)	(3,5)	12 ²	11	o-10	8253.293	8251.782(12)	(11,9)	02 ⁰	8	p-25	8496.564		(8,2)u	11 ¹
7	p-26	7993.597		(7,10)	03 ³	8	p-23	8253.742		(8,2)l	11 ¹	9	o-14	8500.247	8499.940(12)	(9,3)l	02 ²
17	p+2	7994.364		(17,14)	00 ⁰	18	p+1	8254.233		(18,16)	00 ⁰	14	o+4	8502.347		(14,12)	10 ⁰
16	p-2	8002.762		(16,11)	00 ⁰	5	p-25	8256.823		(5,7)	12 ²	11	o-11	8505.700	8505.370(12)	(11,9)l	02 ²
12	p+13	8003.329		(12,8)	10 ⁰	10	o+9	8260.234		(10,6)	02 ⁰	19	o+1	8506.101		(19,18)	00 ⁰
7	p-27	8003.851		(7,5)	20 ⁰	3	p-20	8260.436		(3,1)	12 ⁰	12	p-15	8506.621			

TABLE 3—Continued

Q. N. ^a <i>J I P n</i>	E _{calc} ^b (cm ⁻¹)	E _{exp} ^c (cm ⁻¹)	Label ^d Rot. Vib.	Q. N. ^a <i>J I P n</i>	E _{calc} ^b (cm ⁻¹)	E _{exp} ^c (cm ⁻¹)	Label ^d Rot. Vib.	Q. N. ^a <i>J I P n</i>	E _{calc} ^b (cm ⁻¹)	E _{exp} ^c (cm ⁻¹)	Label ^d Rot. Vib.
4 <i>p</i> + 27	8532.944	8532.448(12)	(4,2) <i>l</i> 12 ²	8 <i>p</i> - 27	8719.623		(8,10) 03 ³	13 <i>p</i> + 13	8898.676		(13,8) 10 ⁰
15 <i>o</i> - 4	8538.865		(15,3) 00 ⁰	7 <i>p</i> - 30	8723.372		(7,4) <i>l</i> 03 ¹	14 <i>p</i> + 12	8915.420		(14,14) 02 ⁰
20 <i>p</i> + 1	8539.875		(20,20) 00 ⁰	7 <i>o</i> - 16	8729.497		(7,6) <i>u</i> 03 ¹	15 <i>p</i> - 9	8918.591		(15,17) 02 ²
5 <i>o</i> + 13	8540.141	8539.642(12)	(5,6) 12 ²	16 <i>p</i> + 5	8733.833		(16,16) 10 ⁰	16 <i>o</i> + 3	8919.168		(16,15) <i>u</i> 01 ¹
10 <i>p</i> + 22	8543.308		(10,4) 02 ⁰	12 <i>p</i> - 17	8734.530		(12,2) <i>u</i> 01 ¹	8 <i>o</i> - 13	8919.473		(8,3) 20 ⁰
10 <i>p</i> - 19	8545.624		(10,8) <i>l</i> 11 ¹	18 <i>o</i> - 1	8737.452		(18,15) 00 ⁰	9 <i>p</i> - 30	8924.988	8924.601(21)	(9,2) <i>l</i> 11 ¹
9 <i>p</i> + 22	8548.456		(9,8) 20 ⁰	8 <i>p</i> - 28	8738.991		(8,8) 03 ¹	11 <i>p</i> + 20	8930.701		(11,8) <i>l</i> 02 ²
6 <i>o</i> + 13	8550.179		(6,3) <i>l</i> 03 ³	10 <i>p</i> - 20	8740.025	8739.733(12)	(10,7) <i>u</i> 02 ²	10 <i>o</i> - 11	8931.540		(10,9) 20 ⁰
5 <i>p</i> + 23	8557.172		(5,1) <i>u</i> 03 ¹	2 <i>p</i> + 19	8744.402		(2,1) <i>u</i> 21 ¹	10 <i>p</i> - 23	8935.174	8934.919(13)	(10,5) <i>l</i> 02 ²
11 <i>p</i> - 20	8561.525		(11,10) <i>l</i> 11 ¹	12 <i>p</i> + 18	8746.539		(12,10) <i>l</i> 02 ²	5 <i>p</i> + 26	8938.948		(5,2) 12 ⁰
8 <i>p</i> + 29	8561.972		(8,1) <i>u</i> 11 ¹	4 <i>o</i> + 11	8748.979	8748.137(21)	(4,0) 12 ²	8 <i>p</i> + 32	8939.808		(8,10) 12 ²
9 <i>o</i> - 15	8565.006	8564.715(15)	(9,6) <i>u</i> 11 ¹	4 <i>p</i> + 29	8752.192	8751.462(20)	(4,2) <i>u</i> 12 ²	9 <i>o</i> + 14	8940.205	8940.065(14)	(9,0) 02 ²
1 <i>p</i> + 6	8572.720		(1,1) 21 ¹	17 <i>o</i> + 1	8756.619		(17,12) 00 ⁰	3 <i>p</i> - 25	8942.914		(3,2) <i>u</i> 21 ¹
4 <i>p</i> + 28	8573.375		(4,2) 12 ⁰	6 <i>o</i> - 12	8757.928		(6,6) 03 ³	11 <i>p</i> - 22	8945.891		(11,10) <i>u</i> 11 ¹
1 <i>o</i> - 6	8574.403		(1,0) 21 ¹	5 <i>p</i> + 24	8759.484		(5,1) <i>u</i> 03 ³	3 <i>p</i> + 18	8946.610		(3,1) <i>l</i> 21 ¹
16 <i>o</i> - 2	8575.190		(16,9) 00 ⁰	2 <i>o</i> - 6	8761.535		(2,0) 21 ¹	6 <i>p</i> - 27	8946.897		(6,2) <i>u</i> 03 ¹
14 <i>p</i> + 11	8575.666		(14,11) <i>u</i> 01 ¹	9 <i>o</i> - 16	8762.271	8762.094(11)	(9,3) <i>u</i> 02 ²	6 <i>o</i> + 15	8948.553	8946.888(12)*	(6,6) 12 ⁰
9 <i>p</i> - 25	8575.676		(9,1) <i>l</i> 02 ²	10 <i>p</i> + 24	8763.050		(10,2) 02 ⁰	12 <i>o</i> - 9	8949.021		(12,3) 10 ⁰
10 <i>p</i> + 23	8578.169		(10,10) 20 ⁰	9 <i>o</i> + 13	8765.826		(9,3) <i>l</i> 11 ¹	9 <i>p</i> + 27	8952.113		(9,1) <i>l</i> 11 ¹
4 <i>p</i> - 20	8579.827		(4,1) 12 ⁰	11 <i>p</i> - 21	8772.922		(11,7) 02 ⁰	14 <i>p</i> - 10	8955.250		(14,10) <i>u</i> 01 ¹
12 <i>p</i> - 16	8582.886		(12,5) 10 ⁰	15 <i>o</i> - 5	8773.852		(15,12) <i>l</i> 01 ¹	9 <i>p</i> - 31	8957.405		(9,8) <i>l</i> 03 ¹
8 <i>o</i> - 12	8583.064		(8,0) 11 ¹	5 <i>p</i> - 30	8774.892	8774.058(31)	(5,5) 12 ²	9 <i>o</i> - 17	8959.407		(9,0) 11 ¹
13 <i>p</i> - 13	8587.078		(13,8) <i>u</i> 01 ¹	6 <i>p</i> - 25	8777.967		(6,7) 12 ²	5 <i>p</i> + 27	8962.886	8961.863(13)	(5,4) 12 ²
2 <i>o</i> + 7	8590.380		(2,3) 21 ¹	9 <i>p</i> + 25	8780.294	8780.023(20)	(9,5) <i>u</i> 11 ¹	3 <i>o</i> - 13	8971.474		(3,0) 21 ¹
11 <i>p</i> + 18	8594.359		(11,10) 02 ²	10 <i>p</i> - 21	8786.124		(10,8) <i>u</i> 11 ¹	5 <i>p</i> - 32	8971.972		(5,1) 12 ⁰
11 <i>p</i> + 19	8605.314		(11,8) 02 ⁰	14 <i>o</i> + 5	8786.369		(14,9) <i>l</i> 01 ¹	6 <i>o</i> - 13	8973.388		(6,0) 03 ¹
6 <i>p</i> + 29	8605.757		(6,1) <i>l</i> 03 ¹	13 <i>p</i> - 14	8787.329		(13,4) <i>l</i> 01 ¹	5 <i>o</i> + 15	8979.321		(5,0) 12 ⁰
13 <i>p</i> + 10	8611.877		(13,14) 02 ²	13 <i>p</i> - 15	8787.748		(13,14) 11 ¹	8 <i>o</i> - 14	8980.344	8978.305(13) [†]	(8,6) <i>l</i> 03 ¹
7 <i>p</i> + 25	8611.959		(7,5) <i>l</i> 03 ¹	8 <i>p</i> + 31	8789.389		(8,4) 20 ⁰	17 <i>p</i> + 3	8981.093		(17,17) 01 ¹
8 <i>p</i> + 30	8615.863		(8,7) <i>l</i> 03 ¹	4 <i>p</i> - 22	8791.144	8790.446(21)	(4,1) <i>u</i> 12 ²	11 <i>o</i> - 12	8982.013	8981.477(12)	(11,9) <i>u</i> 02 ²
9 <i>p</i> + 23	8618.836		(9,4) <i>u</i> 02 ²	7 <i>p</i> - 31	8792.745		(7,8) 03 ³	13 <i>p</i> - 16	8983.051		(13,2) <i>l</i> 01 ¹
5 <i>p</i> - 28	8620.874		(5,4) 03 ³	5 <i>p</i> - 31	8793.409		(5,2) <i>u</i> 03 ³	10 <i>p</i> + 27	8983.186		(10,11) 03 ¹
6 <i>p</i> - 24	8624.270		(6,4) <i>u</i> 03 ¹	15 <i>p</i> + 8	8794.725		(15,13) <i>u</i> 01 ¹	6 <i>p</i> + 31	8985.268		(6,1) <i>l</i> 03 ³
9 <i>p</i> - 26	8626.456		(9,10) 03 ¹	5 <i>p</i> + 25	8794.729		(5,4) 12 ⁰	12 <i>p</i> + 20	8986.242		(12,2) 10 ⁰
8 <i>p</i> - 26	8629.365		(8,5) 20 ⁰	10 <i>p</i> + 25	8795.279		(10,13) 03 ³	7 <i>p</i> + 26	8987.697		(7,1) <i>l</i> 03 ¹
5 <i>p</i> - 29	8636.047	8634.652(14)	(5,5) 12 ⁰	6 <i>p</i> - 26	8795.648		(6,2) <i>l</i> 03 ³	19 <i>p</i> - 2	8993.902		(19,17) 00 ⁰
12 <i>o</i> + 9	8641.551		(12,3) <i>u</i> 01 ¹	5 <i>o</i> - 15	8808.177	8807.586(18)	(5,3) <i>l</i> 12 ²	11 <i>p</i> - 23	8997.677		(11,11) 20 ⁰
9 <i>p</i> + 24	8645.701	8645.499(29)	(9,2) <i>l</i> 02 ²	16 <i>p</i> + 6	8809.895		(16,8) 00 ⁰	10 <i>o</i> + 12	9017.962	9017.824(11)	(10,6) <i>u</i> 02 ²
13 <i>o</i> - 7	8650.911		(13,9) 10 ⁰	10 <i>p</i> + 26	8819.788		(10,7) <i>l</i> 11 ¹	5 <i>o</i> - 17	9050.652	9049.264(18)*	(5,6) 21 ¹
13 <i>p</i> + 11	8651.186		(13,5) <i>l</i> 01 ¹	12 <i>p</i> + 19	8820.198		(12,4) 10 ⁰	5 <i>o</i> - 18	9078.301	9077.153(32)	(5,3) <i>u</i> 12 ²
12 <i>o</i> + 10	8651.889		(12,12) 02 ²	10 <i>p</i> - 22	8828.587		(10,1) 02 ⁰	9 <i>o</i> + 17	9165.074	9164.618(22)	(9,3) <i>u</i> 11 ¹
5 <i>o</i> - 14	8652.117		(5,0) 03 ³	6 <i>o</i> + 14	8834.528	8833.688(09)	(6,3) <i>u</i> 03 ¹	5 <i>p</i> + 29	9186.991	9185.774(37)	(5,2) <i>u</i> 12 ²
15 <i>p</i> + 6	8654.498		(15,2) 00 ⁰	9 <i>p</i> - 28	8835.646		(9,7) 20 ⁰	5 <i>o</i> + 16	9241.971	9241.090(21)	(5,0) 12 ²
7 <i>o</i> - 15	8671.382		(7,9) 12 ²	18 <i>p</i> + 2	8837.258		(18,19) 01 ¹	11 <i>p</i> - 26	9252.173	9251.699(15)	(11,7) <i>u</i> 02 ²
12 <i>p</i> + 17	8676.488		(12,1) <i>u</i> 01 ¹	3 <i>o</i> + 12	8841.826		(3,3) 21 ¹	10 <i>o</i> - 13	9268.164	9267.795(13)	(10,3) <i>l</i> 02 ²
4 <i>o</i> - 12	8680.186	8679.526(12)	(4,3) 12 ²	13 <i>p</i> + 12	8847.726		(13,7) <i>u</i> 01 ¹	6 <i>o</i> + 17	9290.552	9289.879(12)	(6,3) <i>u</i> 03 ³
5 <i>o</i> + 14	8683.673	8682.938(12)	(5,3) 03 ³	4 <i>p</i> + 30	8852.976	8852.149(12)	(4,5) 21 ¹	6 <i>p</i> + 36	9314.461	9313.066(21)	(6,4) <i>u</i> 12 ²
2 <i>p</i> - 13	8688.456		(2,2) 21 ¹	14 <i>p</i> - 9	8859.933		(14,11) 10 ⁰	10 <i>o</i> - 14	9347.534	9346.886(15)	(10,6) <i>u</i> 11 ¹
10 <i>o</i> + 11	8691.214	8690.955(11)	(10,6) <i>l</i> 02 ²	13 <i>o</i> + 7	8861.412		(13,3) <i>l</i> 01 ¹	7 <i>o</i> + 18	9366.462	9364.720(14)	(7,3) <i>u</i> 03 ¹
15 <i>p</i> - 8	8691.410		(15,1) 00 ⁰	7 <i>p</i> - 32	8868.239		(7,4) <i>l</i> 03 ³	6 <i>o</i> - 16	9410.563	9409.489(09)	(6,0) 03 ³
15 <i>p</i> + 7	8691.807		(15,14) 10 ⁰	3 <i>p</i> - 24	8869.083		(3,2) <i>l</i> 21 ¹	6 <i>p</i> + 37	9429.922	9428.608(13)	(6,1) <i>u</i> 03 ³
4 <i>p</i> - 21	8696.975		(4,1) <i>l</i> 12 ²	12 <i>o</i> - 8	8869.733		(12,12) 11 ¹	8 <i>o</i> - 16	9448.280	9446.240(10)*	(8,6) <i>u</i> 03 ¹
3 <i>p</i> - 23	8700.697		(3,4) 21 ¹	9 <i>p</i> - 29	8872.506	8872.119(21)	(9,1) <i>u</i> 02 ²	12 <i>o</i> - 11	9497.654	9497.272(14)	(12,9) <i>u</i> 02 ²
2 <i>p</i> + 18	8704.167		(2,1) <i>l</i> 21 ¹	9 <i>p</i> + 26	8873.489	8873.350(23)	(9,2) <i>u</i> 02 ²	10 <i>o</i> - 16	9564.168	9564.295(15)	(10,3) <i>u</i> 02 ²
15 <i>o</i> + 4	8704.195		(15,0) 00 ⁰	7 <i>o</i> + 16	8885.813		(7,3) <i>l</i> 03 ¹	12 <i>p</i> - 21	9644.523	9643.344(22)	(12,7) <i>l</i> 02 ²
12 <i>o</i> - 7	8706.587	8706.844(26)	(12,0) 01 ¹	5 <i>o</i> - 16	8887.790	8886.456(13)	(5,3) 12 ⁰	11 <i>o</i> + 15	9886.631	9886.079(13)	(11,6) <i>u</i> 02 ²
9 <i>p</i> - 27	8707.442		(9,4) <i>l</i> 11 ¹	11 <i>o</i> + 10	8894.796		(11,9) <i>l</i> 11 ¹				
10 <i>o</i> - 10	8712.267		(10,3) 02 ⁰	6 <i>p</i> + 30	8898.674		(6,1) <i>u</i> 03 ¹				

^a Quantum numbers *J, I* (*o* for *I* = 3/2 and *p* for *I* = 1/2), and parity (*P*). The column labeled *n* is an index for levels with the same *J, I*, and *P* ordering them by energy.

^b Calculated energy value from Watson (52).

^c Experimentally determined energy with its uncertainty in the last digits (2σ) in parentheses.

^d Rotational and vibrational labels assigned as described in Section III.1.

[†] Unusually large deviation from *ab initio* calculations.

* Level constructed using only transitions verified by combination differences.

of these discrepancies can be blamed on the arbitrary assignment of the highly mixed levels in Table 2. Most of the remaining disagreements appear to be errors in the assignments of Din97. During our analysis, we found that some of their labels violated parity and symmetry requirements, some levels were labeled as a being part of a pair of G levels when there was only one way to form G , and one label was assigned to two separate levels. Probably most of these misassignments were caused by not considering all of the levels simultaneously, which was essential to our analysis.

III.2. Compilation and Assignment of Laboratory Data

Once the energy levels were given unique labels, the next step was to compile and analyze the frequency, uncertainty, and assignment for every transition reported. Every study in Table 1 was included in our analysis. (Please note that reference to each of these works for the remainder of the paper will be made using the labels assigned in Table 1.)

Instead of reviewing every assignment made (many transitions have been assigned and reassigned more than three times), we decided to consider all of the data simultaneously and make our own assignments independently. To make the assignments, we compared the transition frequencies to the variational calculations of Watson (55) and Neale, Miller, and Tennyson (NMT) (56), which are both based on spectroscopically fitted potential energy surfaces. We found that a combination of both calculations was necessary in our analysis. The NMT calculations were very good, generally differing from experiment by $\sim 0.05 \text{ cm}^{-1}$. There is a serious problem with these predictions, however, for levels with $J > 9$, and the error can be as high as several cm^{-1} (see Section IV.2). Watson's calculations, though not as precise, are very reliable and were used to assign levels of high J .

The intensity predictions³ of both calculations were very similar, with NMT's, on average, lower than Watson's by $\sim 1\%$ (with a standard deviation of 8%). H_3^+ is known to exhibit nonthermal population distributions in laboratory discharges, but can be described effectively as having thermal distributions among vibrational states and rotational levels, individually (31). We adjusted the theoretical intensities accordingly, assuming a vibrational temperature of 1200 K and a rotational temperature of 500 K. While the discharges used in each of the experiments had different temperatures, the values that we chose are roughly the average, and served to predict the order of magnitude of each transition's intensity.

Transition intensities were only reported in the literature for a few of the studies. Fortunately, we had access to all of the previous laser scans performed in Chicago (Baw90, Xu90, Lee91, Xu92, Ven94, Uy94, Joo00, McC00, and Lin01), and it was very

important to our assignments to look at the transition intensities. By comparing the experimental intensities to the theoretically calculated intensities, we were able to determine roughly the sensitivity cutoff, limiting the number of possible lines available for each assignment. For lines that were very close together, it was useful to look at the original scans to see if some features were hidden on the shoulder of other transitions. In several cases we concluded that two calculated transitions were completely overlapped and were observed as a single feature. Studying these scans also enabled us to judge the quality of each line and make an estimate of the uncertainty on a line by line basis. We did not have access to the raw data from the FTIR emission studies and were not able to make such judgements on those lines.

During our analysis, we found that the uncertainties reported in the literature did not account for the discrepancies between different measurements of transition frequencies. This prompted us to re-examine the uncertainty for every experiment, and in most circumstances to increase them. We were rather conservative in our assignment of uncertainties, preferring to overestimate rather than underestimate the error. It is probably safe to consider our values as roughly two times the standard deviation.

A few systematic errors were identified which also have led to an increase in the uncertainty. As recently reported in McC00, it was found that the rate at which a scan needs to be performed is much slower than had previously been thought. In this work, the authors observed small shifts in the transition frequency due to the scan rate and the lock-in detection time constant. We have studied this phenomenon carefully and have concluded that one needs to spend at least 30 time constants on a velocity-modulated transition to avoid a frequency shift in the absorption feature. This requirement was not met in previous laser scans (in Chicago) and must be taken into account by increasing the uncertainty in every transition to 0.01 cm^{-1} . This error will not apply to the FTIR emission and absorption data (Maj87, Maj89, Nak90, Maj94, McK98). Another frequency error was noticed in the work of Uy94, in which several of the reported transitions disagreed with other reported values by $0.02\text{--}0.03 \text{ cm}^{-1}$.

Finally, there seemed to be a larger than expected difference in some of the reported FTIR emission transition frequencies (Maj89, Maj94) when compared to the theory and laser absorption experiments. Some of the lines that we were unable to assign (Table 4) from Maj89 and Maj94 were within 0.2 cm^{-1} of theoretically predicted lines that should be very strong. One difficulty with the FTIR emission experiments is the ubiquitous Rydberg H_2 emission. While these background features were identified by their strong pressure dependence (the Rydbergs are quenched at higher pressure), the apparent H_3^+ line position could be displaced if it were on the side of a strong H_2 signal. It is also possible that some of the lines attributed to H_3^+ are in fact H_2 lines which happen to increase in intensity with pressure.

It became apparent during our work that many of the assignments in Din97 were based on frequency alone. While the differences between the calculated and observed transition

³ Note that in the paper of Neale, Miller, and Tennyson (56), the upper and lower values of J are switched in the equations relating transition probabilities to the Einstein A -coefficients. In their equations (2) and (3), each J' should be changed to J'' , and vice versa.

TABLE 4
Remaining Unassigned Transitions

Frequency (cm ⁻¹)	Ref ^a	Frequency (cm ⁻¹)	Ref ^a	Frequency (cm ⁻¹)	Ref ^a
1980.367	Maj94	2702.321*	Baw90	3124.264*	Lin01
2028.198	Maj94	2708.432	Baw90	3128.912*	Xu92
2134.607	Maj94	2708.778	Baw90	3137.325	Xu92
2174.478	Maj94	2716.843*	Baw90	3161.895*	Xu92
2405.031	Baw90	2754.319	Baw90	3175.891	Maj94
2483.977	Baw90	2807.248*	Baw90	3177.467*	Maj94
2579.828	Baw90	2882.795*	Baw90	3180.420*	Xu92
2611.471*	Baw90	2915.872	Xu92	3182.593	Xu92
2612.538	Baw90	2918.157	Xu92	3182.605*	Lin01
2614.022*	Baw90	2932.711	Baw90	3188.562*	Maj94
2622.894*	Baw90	2942.920*	Maj94	3205.732*	Maj94
2623.274*	Baw90	2950.516	Maj94	3206.893*	Xu92
2626.289	Baw90	2958.735	Xu92	3235.521*	Lin01
2630.492	Baw90	2958.899	Xu92	3241.009	Maj94
2630.603*	Baw90	2965.791*	Xu92	3249.591*	Lin01
2653.290*	Baw90	2987.381	Maj94	3357.525*	Lin01
2653.559*	Baw90	2990.280*	Maj94	4394.944	Maj94
2653.692*	Baw90	2995.601*	Xu92	4587.373	Maj89
2672.862	Baw90	3005.898	Xu92	4756.345	Maj89
2673.229	Baw90	3022.332	Xu92	4788.544	Maj89
2674.344*	Baw90	3023.904*	Maj94	4823.315	Xu90
2680.330	Baw90	3104.125*	Lin01	4823.348	Maj89
2680.485	Baw90	3120.826	Xu92	4823.892	Maj94
2699.334*	Baw90	3121.475*	Xu92	4942.862	Maj89

Note. Some of these lines were previously assigned but have been ‘unassigned’ during our analysis. Transitions marked with asterisks do not have any reasonable assignment and are likely not due to H₃⁺. Lines without an asterisk had one or more candidate assignments whose frequency and/or intensity difference from theory was too large to make a confident assignment.

^a Reference from which the transition frequency was taken. Labels used in this column are defined in Table 1.

frequencies were usually small, sometimes assignments were made to transitions predicted to have intensities orders of magnitude weaker than the experimental sensitivity. This may partly be due to the fact that observed intensities are rarely published in the literature, and we urge experimenters to publish this information in the future. Frequently our reassignment of such lines increased the frequency difference from theory but ultimately made a much more reasonable assignment.

Once all of the assignments were made, we verified many of them by checking for combinations of other transitions that led to the same energy differences (combination differences). A program was written to search for all possible combinations of transitions that created closed ‘‘loops’’ of up to 6 transitions. The frequency of every verified transition agreed within 1.5 times the uncertainty in the frequency calculated by a combination of other transitions.

To label the transitions, we have extended the energy level notation from Section III.1 using the band symbol

$$v'_1 v_1 + v'_2 v_2^{|\ell'|} \leftarrow v''_1 v_1 + v''_2 v_2^{|\ell''|} \quad [7]$$

or more compactly

$$v'_1 v_2^{|\ell'|} \leftarrow v''_1 v_2^{|\ell''|} \quad [8]$$

and the branch symbol

$$\{n|t|\pm 6|\pm 9|\dots\} \{P|Q|R\} \{J'', G''\}_{\{u|l\}} \quad [9]$$

where *P*, *Q*, and *R* correspond to the usual $\Delta J = -1, 0, +1$. As was done for the level labels, *u* and *l* are appended to the end of the symbol, when appropriate, as a superscript and/or subscript referring to the upper and lower states in the transition, respectively. The preceding superscript specifies ΔG when it is not 0. For overtone and forbidden bands ΔG can equal ± 3 (or ± 1),⁴ signified by *t* and *n* for the + and -, respectively. For highly mixed levels $|\Delta G| > 3$ are possible and these are labeled by $\pm 6, \pm 9$, etc.

A total of 895 unique transition frequencies have been reported in the literature, and we were able to assign 823 of them to transitions of H₃⁺. Table 5 lists the adopted frequency, estimated uncertainty, assignment, and literature reference for each assigned transition. The assignments of 486 of these transitions were verified by their combination differences and are denoted by asterisks. For transitions that have been reported multiple times, we used the least uncertain measurement for the frequency. In cases where more than one equally accurate measurement was available, we chose the earliest measurement to include in this table. Many of the previous assignments have been changed due to both the new labeling scheme of energy levels and the reassignment of lines to different transitions. Surprisingly, we found that fewer than 4% of the lines were assigned incorrectly upon their initial observation. Most of the assignment conflicts were in the lines that were not initially assigned in Baw00 and Xu92. An expanded version of this table is available in electronic form online and includes the calculated lower state energies and Einstein *A* coefficients. This version also credits the first reported observation and first correct assignment of each line. An online intensity calculator is also available at the authors’ Web site (<http://h3plus.uchicago.edu>).

The remaining 72 unassigned transitions are listed in Table 4 and should be considered carefully before being assigned in the future. Some of these that had been previously assigned are no longer assigned. Many of them (marked with an asterisk) had no reasonable theoretically predicted lines of sufficient intensity within ~ 1 cm⁻¹ of the reported transition, and are likely

⁴ This ‘‘rule’’ is somewhat misleading and deserves more explanation. The signed *G*, denoted $g \equiv k - \ell$, carries the selection rule of $\Delta g = 0, \pm 3, \pm 6, \dots$ due to the parity and nuclear spin selection rules. The confusion begins when *g* goes from a positive to a negative value or vice versa. Take for example an overtone transition where $k'' = \pm 1, \ell'' = 0$ and $k' = 0, \ell' = \pm 2$. In this case $g'' = \pm 1, g' = \mp 2, G'' = 1$, and $G' = 2$. The transition $\Delta g = \mp 3$ is clearly allowed but ΔG appears to be a misleading +1. Both transitions are properly labeled with an *n*; a label of *t* would denote the transition $g'' = \pm 1$ to $g' = \pm 4$ where $\Delta g = \pm 3$ and $\Delta G = +3$.

TABLE 5
Observed and Assigned Laboratory Transitions of H₃⁺

Frequency ^a (cm ⁻¹)	Assignment ^b Label Band	Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b Label Band	Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b Label Band	Ref ^c			
1546.901 (10)	<i>P</i> (12, 12)	01 ¹ ← 00 ⁰	Joo00	2134.241 (10)*	<i>Q</i> (7, 6) _u	02 ⁰ ← 01 ¹	Maj94	2395.500 (10)*	<i>Q</i> (8, 3) _l	01 ¹ ← 00 ⁰	Baw90
1798.396 (02)*	<i>P</i> (9, 9)	01 ¹ ← 00 ⁰	Maj87	2134.922 (10)*	<i>P</i> (5, 5)	01 ¹ ← 00 ⁰	Maj87	2397.911 (10)*	<i>Q</i> (9, 5) _l	01 ¹ ← 00 ⁰	Maj94
1826.160 (02)	<i>P</i> (9, 8)	01 ¹ ← 00 ⁰	Maj87	2137.039 (10)*	<i>P</i> (5, 3) _u	02 ² ← 01 ¹	Maj94	2398.519 (10)*	<i>Q</i> (8, 7) _l	01 ¹ ← 00 ⁰	Maj87
1843.560 (10)*	<i>P</i> (10, 7) ^u	01 ¹ ← 00 ⁰	Maj94	2140.348 (10)*	<i>P</i> (5, 4)	01 ¹ ← 00 ⁰	Maj87	2399.749 (10)*	<i>Q</i> (1, 1)	11 ¹ ← 10 ⁰	Baw90
1865.199 (10)*	<i>P</i> (9, 7) ^u	01 ¹ ← 00 ⁰	Maj94	2142.328 (10)*	<i>P</i> (5, 4) _u	02 ² ← 01 ¹	Maj94	2402.621 (10)*	<i>Q</i> (6, 0)	02 ² ← 01 ¹	Baw90
1867.905 (10)	<i>P</i> (11, 6) ^u	01 ¹ ← 00 ⁰	Maj94	2152.615 (10)	<i>P</i> (3, 0)	11 ¹ ← 10 ⁰	Maj94	2403.350 (20)*	<i>Q</i> (2, 3)	12 ² ← 11 ¹	Baw90
1868.703 (10)*	<i>P</i> (9, 10)	02 ² ← 01 ¹	Maj94	2152.887 (10)*	<i>P</i> (5, 3) ^u	01 ¹ ← 00 ⁰	Maj87	2406.029 (10)*	<i>Q</i> (2, 1) ^u	11 ¹ ← 10 ⁰	Baw90
1876.392 (10)*	<i>P</i> (9, 9)	02 ² ← 01 ¹	Maj94	2160.320 (10)*	<i>P</i> (5, 5)	02 ² ← 01 ¹	Maj94	2408.730 (10)*	<i>Q</i> (8, 6) _l	01 ¹ ← 00 ⁰	Maj87
1882.985 (10)	<i>P</i> (8, 8)	01 ¹ ← 00 ⁰	Maj94	2164.278 (10)*	<i>P</i> (5, 2) ^u	01 ¹ ← 00 ⁰	Maj87	2411.518 (10)*	<i>Q</i> (8, 5) _l	01 ¹ ← 00 ⁰	Maj87
1883.755 (10)*	<i>P</i> (10, 6) ^u	01 ¹ ← 00 ⁰	Maj94	2168.349 (10)*	<i>P</i> (5, 6)	02 ² ← 01 ¹	Baw90	2412.859 (10)*	<i>Q</i> (2, 2)	11 ¹ ← 10 ⁰	Baw90
1904.235 (10)*	<i>P</i> (8, 7)	01 ¹ ← 00 ⁰	Maj94	2168.698 (10)*	<i>P</i> (3, 3)	11 ¹ ← 10 ⁰	Baw90	2413.314 (10)*	<i>Q</i> (5, 1) _u	02 ² ← 01 ¹	Baw90
1905.488 (10)*	<i>P</i> (9, 6) ^u	01 ¹ ← 00 ⁰	Maj94	2172.815 (10)*	<i>P</i> (5, 1) ^u	01 ¹ ← 00 ⁰	Maj87	2413.922 (10)*	<i>R</i> (1, 1)	02 ⁰ ← 01 ¹	Baw90
1916.714 (10)	<i>P</i> (10, 5) ^u	01 ¹ ← 00 ⁰	Maj94	2175.780 (10)*	<i>P</i> (5, 0)	01 ¹ ← 00 ⁰	Maj87	2416.289 (10)*	<i>R</i> (1, 0)	21 ¹ ← 20 ⁰	Baw90
1921.286 (10)*	<i>P</i> (8, 6) _u	02 ² ← 01 ¹	Maj94	2182.348 (10)*	<i>P</i> (4, 2) _u	02 ² ← 01 ¹	Maj94	2417.764 (10)*	<i>Q</i> (7, 4) _l	01 ¹ ← 00 ⁰	Maj87
1925.254 (10)	<i>P</i> (11, 4) ^u	01 ¹ ← 00 ⁰	Maj94	2197.743 (10)*	<i>P</i> (4, 3) _u	02 ² ← 01 ¹	Maj94	2418.899 (10)*	<i>Q</i> (7, 0)	01 ¹ ← 00 ⁰	Baw90
1927.291 (10)	<i>P</i> (11, 0)	01 ¹ ← 00 ⁰	Maj94	2202.691 (10)	<i>P</i> (4, 6)	03 ³ ← 02 ²	Maj94	2419.558 (30)	<i>Q</i> (7, 1) _l	01 ¹ ← 00 ⁰	Baw90
1927.792 (10)	<i>P</i> (7, 2) ^u	11 ¹ ← 10 ⁰	Maj94	2217.451 (10)*	<i>P</i> (4, 4)	01 ¹ ← 00 ⁰	Wat84	2420.207 (10)	<i>Q</i> (3, 2) ^u	11 ¹ ← 10 ⁰	Baw90
1933.653 (10)*	<i>P</i> (4, 3) _l	02 ⁰ ← 01 ¹	Maj94	2218.129 (10)*	<i>P</i> (4, 3)	01 ¹ ← 00 ⁰	Wat84	2420.728 (10)	<i>Q</i> (7, 3) _l	01 ¹ ← 00 ⁰	Baw90
1935.714 (10)*	<i>P</i> (8, 6) _l	01 ¹ ← 00 ⁰	Maj94	2223.965 (10)*	<i>P</i> (4, 2) ^u	01 ¹ ← 00 ⁰	Wat84	2421.888 (10)*	<i>Q</i> (7, 2) ^u	01 ¹ ← 00 ⁰	Baw90
1937.873 (10)*	<i>P</i> (8, 7) _u	02 ² ← 01 ¹	Maj94	2229.895 (10)*	<i>P</i> (4, 1) ^u	01 ¹ ← 00 ⁰	Wat84	2422.983 (10)*	<i>Q</i> (3, 1) _u	02 ² ← 01 ¹	Baw90
1939.934 (10)*	<i>P</i> (9, 5) ^u	01 ¹ ← 00 ⁰	Maj94	2229.912 (10)*	<i>P</i> (4, 4)	02 ² ← 01 ¹	Maj94	2423.646 (10)*	<i>Q</i> (7, 6) _l	01 ¹ ← 00 ⁰	Maj87
1944.087 (10)*	<i>P</i> (8, 9)	02 ² ← 01 ¹	Maj94	2241.077 (10)*	<i>R</i> (2, 2)	11 ¹ ← 10 ⁰	Baw90	2423.675 (20)*	<i>Q</i> (4, 1) _u	02 ² ← 01 ¹	Baw90
1945.254 (10)	<i>P</i> (10, 4) ^u	01 ¹ ← 00 ⁰	Maj94	2241.347 (10)*	<i>P</i> (4, 5)	02 ² ← 01 ¹	Baw90	2424.797 (10)*	<i>Q</i> (3, 3)	11 ¹ ← 10 ⁰	Baw90
1947.467 (10)*	<i>P</i> (8, 8)	02 ² ← 01 ¹	Maj94	2250.525 (10)	<i>Q</i> (7, 0)	02 ⁰ ← 01 ¹	Maj94	2431.821 (10)*	<i>Q</i> (7, 5) _l	01 ¹ ← 00 ⁰	Maj87
1958.420 (10)	<i>P</i> (10, 1) ^u	01 ¹ ← 00 ⁰	Maj94	2253.633 (10)*	<i>Q</i> (1, 0)	02 ⁰ ← 01 ¹	Maj94	2433.901 (10)*	<i>Q</i> (4, 3) ^u	11 ¹ ← 10 ⁰	Baw90
1959.957 (50)*	<i>P</i> (3, 1) _u	02 ⁰ ← 01 ¹	Maj94	2260.480 (10)	<i>Q</i> (11, 3) _l	01 ¹ ← 00 ⁰	Maj94	2436.653 (10)*	^t <i>Q</i> (6, 3) _u	11 ¹ ← 01 ¹	Baw90
1967.450 (02)*	<i>P</i> (7, 7)	01 ¹ ← 00 ⁰	Maj87	2260.480 (10)	<i>Q</i> (1, 1)	03 ¹ ← 02 ⁰	Maj94	2438.509 (10)*	<i>Q</i> (4, 4)	11 ¹ ← 10 ⁰	Baw90
1968.800 (02)*	<i>P</i> (8, 5) ^u	01 ¹ ← 00 ⁰	Maj87	2265.551 (10)	<i>Q</i> (6, 2) _l	02 ⁰ ← 01 ¹	Maj94	2446.632 (10)*	<i>Q</i> (6, 5) _l	01 ¹ ← 00 ⁰	Maj87
1969.319 (10)*	<i>P</i> (9, 4) ^u	01 ¹ ← 00 ⁰	Maj94	2271.405 (10)*	<i>Q</i> (6, 3) _l	02 ⁰ ← 01 ¹	Maj94	2447.903 (10)*	<i>Q</i> (6, 1) _l	01 ¹ ← 00 ⁰	Maj87
1977.313 (10)	<i>P</i> (9, 2) ^u	01 ¹ ← 00 ⁰	Maj94	2274.262 (10)	<i>Q</i> (11, 0)	01 ¹ ← 00 ⁰	Maj94	2449.533 (10)*	<i>Q</i> (6, 2) _l	01 ¹ ← 00 ⁰	Maj87
1981.672 (10)*	<i>P</i> (7, 3) _u	02 ² ← 01 ¹	Maj94	2277.104 (10)*	<i>Q</i> (5, 3) _l	02 ⁰ ← 01 ¹	Maj94	2449.800 (10)*	<i>Q</i> (4, 0)	02 ² ← 01 ¹	Baw90
1982.486 (10)*	<i>P</i> (6, 6)	11 ¹ ← 10 ⁰	Maj94	2279.406 (30)*	<i>Q</i> (3, 1) _l	02 ⁰ ← 01 ¹	Maj94	2449.885 (10)	<i>P</i> (1, 2)	02 ² ← 01 ¹	Baw90
1982.874 (02)*	<i>P</i> (7, 6)	01 ¹ ← 00 ⁰	Maj87	2279.406 (30)*	<i>Q</i> (2, 2) _l	02 ⁰ ← 01 ¹	Maj94	2452.718 (10)*	<i>Q</i> (6, 3) _l	01 ¹ ← 00 ⁰	Maj87
1984.067 (10)*	<i>P</i> (7, 5) _u	02 ² ← 01 ¹	Maj94	2279.632 (10)*	<i>Q</i> (3, 0)	02 ⁰ ← 01 ¹	Maj94	2453.408 (10)*	<i>Q</i> (6, 4) _l	01 ¹ ← 00 ⁰	Maj87
1990.807 (10)*	<i>P</i> (6, 0)	02 ² ← 01 ¹	Maj94	2279.913 (10)	<i>Q</i> (5, 0)	02 ⁰ ← 01 ¹	Maj94	2454.417 (10)*	^t <i>Q</i> (7, 4)	10 ⁰ ← 00 ⁰	Baw90
1996.884 (10)*	<i>P</i> (8, 4) ^u	01 ¹ ← 00 ⁰	Maj94	2280.547 (10)	<i>Q</i> (8, 1) _l	02 ⁰ ← 01 ¹	Maj94	2456.273 (20)*	<i>Q</i> (4, 2) _l	02 ² ← 01 ¹	Baw90
1997.172 (10)	<i>P</i> (9, 1) ^u	01 ¹ ← 00 ⁰	Maj94	2284.000 (10)*	<i>Q</i> (4, 2) _l	02 ⁰ ← 01 ¹	Maj94	2457.290 (05)	<i>P</i> (1, 1)	01 ¹ ← 00 ⁰	McK98
2001.479 (10)	<i>P</i> (9, 0)	01 ¹ ← 00 ⁰	Maj94	2284.333 (10)	<i>Q</i> (4, 1) _l	02 ⁰ ← 01 ¹	Maj94	2457.613 (10)*	<i>Q</i> (5, 5)	11 ¹ ← 10 ⁰	Baw90
2002.045 (10)*	<i>P</i> (9, 3) ^u	01 ¹ ← 00 ⁰	Maj94	2295.577 (10)*	<i>P</i> (3, 1) ^u	01 ¹ ← 00 ⁰	Wat84	2457.912 (10)*	<i>R</i> (1, 0)	03 ¹ ← 02 ⁰	Baw90
2006.615 (10)*	<i>P</i> (7, 6) _u	02 ² ← 01 ¹	Maj94	2295.947 (10)*	<i>P</i> (3, 2)	01 ¹ ← 00 ⁰	Wat84	2458.850 (10)	<i>R</i> (1, 1) ^u	03 ¹ ← 02 ⁰	Baw90
2007.290 (10)*	<i>P</i> (7, 5) ^u	01 ¹ ← 00 ⁰	Maj87	2295.980 (10)*	<i>P</i> (3, 0)	01 ¹ ← 00 ⁰	Wat84	2464.652 (10)*	<i>R</i> (2, 3)	02 ⁰ ← 01 ¹	Baw90
2011.400 (10)*	<i>P</i> (6, 3) _u	02 ² ← 01 ¹	Maj94	2298.930 (10)*	<i>P</i> (3, 3)	01 ¹ ← 00 ⁰	Wat84	2467.553 (10)*	<i>Q</i> (5, 4) _l	01 ¹ ← 00 ⁰	Baw90
2018.029 (10)*	<i>P</i> (8, 3) ^u	01 ¹ ← 00 ⁰	Maj94	2304.343 (10)*	<i>P</i> (3, 3)	02 ² ← 01 ¹	Maj94	2469.235 (10)	<i>Q</i> (5, 3) _u	03 ³ ← 02 ²	Baw90
2018.760 (10)*	<i>P</i> (7, 7)	02 ² ← 01 ¹	Maj94	2312.918 (10)*	<i>P</i> (3, 4)	02 ² ← 01 ¹	Maj94	2470.605 (10)	^t <i>Q</i> (8, 4)	10 ⁰ ← 00 ⁰	Baw90
2019.376 (10)*	<i>P</i> (7, 8)	02 ² ← 01 ¹	Maj94	2314.681 (10)	<i>Q</i> (10, 4) _l	01 ¹ ← 00 ⁰	Maj94	2471.384 (10)	<i>Q</i> (3, 3) _l	03 ¹ ← 02 ⁰	Baw90
2020.914 (10)*	<i>P</i> (5, 4)	11 ¹ ← 10 ⁰	Maj94	2324.698 (10)	<i>Q</i> (10, 3) _l	01 ¹ ← 00 ⁰	Maj94	2471.923 (10)	<i>Q</i> (5, 0)	01 ¹ ← 00 ⁰	Baw90
2022.011 (10)*	<i>P</i> (5, 3) ^u	11 ¹ ← 10 ⁰	Maj94	2331.823 (10)	<i>Q</i> (11, 9) _l	01 ¹ ← 00 ⁰	Maj94	2472.325 (10)*	<i>Q</i> (5, 1) _l	01 ¹ ← 00 ⁰	Maj87
2023.165 (10)	<i>P</i> (8, 2) ^u	01 ¹ ← 00 ⁰	Maj94	2333.983 (10)	<i>Q</i> (5, 0)	11 ¹ ← 10 ⁰	Maj94	2472.846 (10)*	<i>Q</i> (5, 3) _l	01 ¹ ← 00 ⁰	Maj87
2032.182 (10)*	<i>P</i> (5, 5)	11 ¹ ← 10 ⁰	Maj94	2334.544 (10)	<i>Q</i> (5, 1) _l	11 ¹ ← 10 ⁰	Maj94	2473.238 (10)*	<i>Q</i> (5, 2) _l	01 ¹ ← 00 ⁰	Maj87
2033.318 (10)*	<i>P</i> (7, 4) ^u	01 ¹ ← 00 ⁰	Maj87	2335.567 (10)*	<i>Q</i> (5, 3) _l	11 ¹ ← 10 ⁰	Maj94	2474.054 (10)*	<i>Q</i> (2, 0)	02 ² ← 01 ¹	Baw90
2036.291 (10)	<i>P</i> (8, 1) ^u	01 ¹ ← 00 ⁰	Maj94	2341.498 (10)	<i>Q</i> (10, 9) _l	01 ¹ ← 00 ⁰	Maj94	2477.797 (10) [†]	<i>Q</i> (4, 2) _u	03 ³ ← 02 ²	Baw90
2051.510 (10)*	<i>P</i> (6, 6)	01 ¹ ← 00 ⁰	Maj87	2348.355 (10)*	<i>Q</i> (9, 3) _l	01 ¹ ← 00 ⁰	Maj94	2483.553 (10)*	<i>Q</i> (3, 1) _l	02 ² ← 01 ¹	Baw90
2054.047 (10)*	<i>P</i> (6, 4) _u	02 ² ← 01 ¹	Maj94	2350.775 (10)	<i>Q</i> (4, 1) _l	11 ¹ ← 10 ⁰	Maj94	2486.559 (05)*	<i>Q</i> (4, 3) _l	01 ¹ ← 00 ⁰	McK98
2057.444 (10)*	<i>P</i> (2, 0)	02 ⁰ ← 01 ¹	Maj94	2351.639 (10)	<i>Q</i> (9, 0)	01 ¹ ← 00 ⁰	Maj94	2486.844 (10)*	<i>R</i> (2, 2)	02 ⁰ ← 01 ¹	Baw90
2060.200 (10)*	<i>P</i> (7, 3) ^u	01 ¹ ← 00 ⁰	Maj87	2353.250 (10)	<i>Q</i> (9, 2) _l	01 ¹ ← 00 ⁰	Maj94	2491.745 (05)*	<i>Q</i> (4, 2) _l	01 ¹ ← 00 ⁰	McK98
2061.680 (10)*	<i>P</i> (6, 5)	01 ¹ ← 00 ⁰	Maj87	2354.125 (10)*	<i>Q</i> (9, 4) _l	01 ¹ ← 00 ⁰	Maj94	2491.906 (10)*	<i>Q</i> (6, 6)	11 ¹ ← 10 ⁰	Baw90
2067.366 (10)*	<i>P</i> (7, 2) ^u	01 ¹ ← 00 ⁰	Maj87	2357.951 (10)*	<i>Q</i> (10, 7) _l	01 ¹ ← 00 ⁰	Maj94	2491.976 (10)*	<i>R</i> (2, 1) _u	02 ⁰ ← 01 ¹	Baw90
2073.951 (10)*	<i>P</i> (6, 5) _u	02 ² ← 01 ¹	Maj94	2360.957 (10)*	<i>Q</i> (10, 6) _l	01 ¹ ← 00 ⁰	Maj94	2492.537 (05)*	<i>Q</i> (4, 1) _l	01 ¹ ← 00 ⁰	McK98
2077.500 (10)*	<i>P</i> (7, 1) ^u	01 ¹ ← 00 ⁰	Maj87	2362.676 (10)							

TABLE 5—Continued

Frequency ^a			Ref ^c	Frequency ^a			Ref ^c	Frequency ^a			Ref ^c
(cm ⁻¹)	Label	Band		(cm ⁻¹)	Label	Band		(cm ⁻¹)	Label	Band	
2520.677 (10)*	Q(4, 1) _f	02 ² ← 01 ¹	Baw90	2617.809 (10)*	Q(5, 6)	03 ³ ← 02 ²	Baw90	2769.393 (10)*	R(3, 3) ^u	11 ¹ ← 10 ⁰	Baw90
2529.724 (05)	Q(1, 0)	01 ¹ ← 00 ⁰	McK98	2620.589 (10)*	Q(8, 6) ^u	01 ¹ ← 00 ⁰	Baw90	2769.863 (10)	R(3, 1) ^u	11 ¹ ← 10 ⁰	Baw90
2532.253 (10)	R(3, 4)	02 ⁰ ← 01 ¹	Baw90	2621.514 (10)	R(4, 4)	02 ⁰ ← 01 ¹	Baw90	2770.196 (10) [†]	R(7, 8)	02 ⁰ ← 01 ¹	Baw90
2534.922 (10)*	Q(5, 2) _u	02 ² ← 01 ¹	Baw90	2624.967 (10)*	Q(6, 3) ^u	01 ¹ ← 00 ⁰	Baw90	2770.940 (10)	R(3, 0)	11 ¹ ← 10 ⁰	Baw90
2536.931 (10)*	Q(4, 2) _u	02 ² ← 01 ¹	Baw90	2626.220 (10)	Q(6, 7)	02 ² ← 01 ¹	Baw90	2771.586 (10)	R(6, 3) _u	02 ⁰ ← 01 ¹	Baw90
2538.253 (10)	R(1, 1) ^f	11 ¹ ← 10 ⁰	Baw90	2628.097 (30)*	Q(4, 2) _f	02 ² ← 01 ¹	Baw90	2783.325 (10)*	R(3, 1) _f	02 ² ← 01 ¹	Baw90
2539.451 (10)*	Q(1, 2)	02 ² ← 01 ¹	Baw90	2628.119 (20)*	Q(7, 7)	02 ² ← 01 ¹	Baw90	2783.417 (10)*	R(2, 1) _u	02 ² ← 01 ¹	Baw90
2539.744 (10)*	Q(5, 3) _u	02 ² ← 01 ¹	Baw90	2630.814 (10)	Q(10, 8) ^u	01 ¹ ← 00 ⁰	Baw90	2785.121 (10)*	R(3, 2) _f	02 ² ← 01 ¹	Baw90
2541.293 (10)*	Q(3, 2) _u	02 ² ← 01 ¹	Baw90	2639.806 (10)*	Q(7, 8)	02 ² ← 01 ¹	Baw90	2787.400 (10)	R(4, 0)	12 ² ← 11 ¹	Baw90
2541.433 (10)*	Q(3, 0)	02 ² ← 01 ¹	Baw90	2640.172 (10)*	Q(8, 8)	02 ² ← 01 ¹	Baw90	2789.736 (10)	R(6, 4) _u	02 ⁰ ← 01 ¹	Baw90
2542.467 (10)*	Q(2, 2)	02 ² ← 01 ¹	Baw90	2648.105 (10)*	R(2, 3)	12 ² ← 11 ¹	Baw90	2795.213 (10)	R(4, 3) ^u	03 ¹ ← 02 ⁰	Baw90
2545.420 (05)*	Q(1, 1)	01 ¹ ← 00 ⁰	McK98	2648.692 (10)*	Q(5, 0)	02 ² ← 01 ¹	Baw90	2798.620 (10)	R(7, 7)	02 ⁰ ← 01 ¹	Baw90
2552.988 (05)*	Q(2, 1) ^u	01 ¹ ← 00 ⁰	McK98	2649.315 (10)	R(4, 3) _f	12 ² ← 11 ¹	Baw90	2801.108 (10)	R(5, 4) _f	11 ¹ ← 10 ⁰	Baw90
2554.276 (10)*	Q(3, 3)	02 ² ← 01 ¹	Baw90	2650.561 (10)	⁻⁶ P(4, 6)	03 ³ ← 02 ²	Baw90	2809.767 (10)*	R(2, 2)	02 ² ← 01 ¹	Baw90
2554.475 (10)*	Q(4, 3)	03 ³ ← 02 ²	Baw90	2650.954 (10)*	Q(9, 9)	02 ² ← 01 ¹	Baw90	2810.597 (10)*	R(5, 3) _f	11 ¹ ← 10 ⁰	Baw90
2554.666 (05)*	Q(2, 2)	01 ¹ ← 00 ⁰	McK98	2653.095 (10)*	Q(8, 9)	02 ² ← 01 ¹	Baw90	2816.843 (10)*	R(2, 3)	02 ² ← 01 ¹	Baw90
2557.484 (10)	R(3, 3)	02 ⁰ ← 01 ¹	Baw90	2653.885 (10)*	R(5, 5) _f	21 ¹ ← 20 ⁰	Baw90	2817.349 (10)*	R(5, 6)	12 ⁰ ← 11 ¹	Baw90
2561.497 (05)*	Q(3, 3)	01 ¹ ← 00 ⁰	McK98	2657.652 (10)	R(5, 6)	02 ⁰ ← 01 ¹	Baw90	2818.072 (10)*	R(4, 2) _u	02 ² ← 01 ¹	Baw90
2564.418 (05)*	Q(3, 2) _u	01 ¹ ← 00 ⁰	McK98	2660.373 (10)*	Q(5, 3) _f	02 ² ← 01 ¹	Baw90	2818.196 (10)*	R(2, 1) _f	02 ² ← 01 ¹	Baw90
2566.904 (10)*	Q(2, 3)	02 ² ← 01 ¹	Baw90	2660.638 (10)*	^f Q(4, 2) _u	11 ¹ ← 01 ¹	Baw90	2821.518 (10)*	R(8, 9)	02 ⁰ ← 01 ¹	Baw90
2567.288 (05)*	Q(4, 4)	01 ¹ ← 00 ⁰	McK98	2664.213 (10)* [†]	R(1, 2)	03 ³ ← 02 ²	Baw90	2822.357 (30)	R(4, 2) _u	12 ² ← 11 ¹	Baw90
2568.708 (05)*	Q(3, 1) ^u	01 ¹ ← 00 ⁰	McK98	2665.729 (10)*	Q(4, 3) _f	02 ² ← 01 ¹	Baw90	2822.448 (30)	R(4, 3) _u	12 ² ← 11 ¹	Baw90
2569.726 (10)*	^f Q(6, 3)	10 ⁰ ← 00 ⁰	Xu92	2666.142 (10) [†]	R(7, 6) _f	03 ¹ ← 02 ⁰	Baw90	2822.730 (20)	R(7, 5) _u	02 ⁰ ← 01 ¹	Baw90
2570.858 (10)*	Q(3, 1) _f	02 ² ← 01 ¹	Baw90	2666.500 (10)	Q(9, 10)	02 ² ← 01 ¹	Baw90	2823.138 (05)*	R(2, 2) ^u	01 ¹ ← 00 ⁰	McK98
2570.987 (10)*	Q(4, 6)	03 ³ ← 02 ²	Baw90	2670.234 (10)*	R(1, 0)	02 ² ← 01 ¹	Baw90	2824.754 (10)	R(7, 4) _u	02 ⁰ ← 01 ¹	Baw90
2571.118 (05)*	Q(5, 5)	01 ¹ ← 00 ⁰	McK98	2671.142 (10)	R(2, 1) ^u	11 ¹ ← 10 ⁰	Baw90	2825.956 (10)*	R(4, 3) _f	02 ² ← 01 ¹	Baw90
2572.220 (10)	R(1, 0)	11 ¹ ← 10 ⁰	Baw90	2672.799 (10)	R(2, 2) ^u	11 ¹ ← 10 ⁰	Baw90	2826.117 (05)*	R(2, 1) ^u	01 ¹ ← 00 ⁰	McK98
2572.357 (10)*	Q(6, 4) _u	02 ² ← 01 ¹	Baw90	2672.958 (10)	R(3, 3) _f	11 ¹ ← 10 ⁰	Baw90	2829.925 (05)*	R(3, 3) _f	01 ¹ ← 00 ⁰	McK98
2573.057 (10)	Q(6, 3) _u	02 ² ← 01 ¹	Baw90	2679.487 (10)	R(4, 2) _u	02 ⁰ ← 01 ¹	Baw90	2831.340 (10)*	R(3, 1) _f	01 ¹ ← 00 ⁰	Wat84
2573.582 (10)*	Q(6, 6)	01 ¹ ← 00 ⁰	Maj87	2680.631 (10)	R(3, 2) _f	11 ¹ ← 10 ⁰	Baw90	2832.198 (05)*	R(3, 2) _f	01 ¹ ← 00 ⁰	McK98
2574.659 (05)*	Q(4, 3) ^u	01 ¹ ← 00 ⁰	McK98	2681.500 (10)	R(3, 1) _f	11 ¹ ← 10 ⁰	Baw90	2836.028 (10)*	ⁿ R(5, 5) _f	02 ² ← 10 ⁰	Baw90
2574.893 (10)*	Q(7, 7)	01 ¹ ← 00 ⁰	Baw90	2683.755 (10)	R(5, 5)	02 ⁰ ← 01 ¹	Baw90	2838.041 (10)	R(6, 6) _f	11 ¹ ← 10 ⁰	Baw90
2575.112 (30)*	Q(9, 9)	01 ¹ ← 00 ⁰	Baw90	2685.157 (10)*	R(4, 3) _u	02 ⁰ ← 01 ¹	Baw90	2841.148 (10)	^f Q(2, 0)	11 ¹ ← 01 ¹	Baw90
2575.112 (10)	R(1, 1) ^u	11 ¹ ← 10 ⁰	Baw90	2685.942 (10)*	^f Q(5, 2) _u	11 ¹ ← 01 ¹	Xu92	2842.191 (10)*	R(5, 3) _u	02 ² ← 01 ¹	Baw90
2575.312 (10)	Q(8, 8)	01 ¹ ← 00 ⁰	Baw90	2691.443 (05)*	R(1, 1) _f	01 ¹ ← 00 ⁰	McK98	2843.898 (20)*	R(3, 1) _f	02 ² ← 01 ¹	Baw90 [†]
2577.492 (10) [†]	Q(2, 3)	03 ³ ← 02 ²	Baw90	2695.420 (10)*	R(1, 1)	02 ² ← 01 ¹	Baw90	2844.464 (10)*	ⁿ R(5, 4) ^u	02 ² ← 10 ⁰	Baw90 [†]
2577.629 (10)*	Q(4, 3) _u	02 ² ← 01 ¹	Baw90	2696.110 (10)*	R(2, 1) _u	02 ² ← 01 ¹	Baw90	2851.433 (10)*	R(8, 8)	02 ⁰ ← 01 ¹	Baw90
2577.694 (10)*	R(1, 1)	03 ³ ← 02 ²	Baw90	2700.573 (10)*	R(3, 0)	12 ² ← 11 ¹	Baw90	2852.156 (10)	^f R(2, 0)	12 ⁰ ← 02 ²	Baw90
2579.390 (10)	R(2, 0)	03 ³ ← 02 ²	Baw90	2704.382 (10)	R(3, 2) ^u	03 ¹ ← 02 ⁰	Baw90	2853.598 (10)*	R(4, 1) _u	02 ² ← 01 ¹	Baw90
2579.672 (10)*	Q(5, 4) _u	02 ² ← 01 ¹	Baw90	2709.405 (10)*	ⁿ R(4, 4) ^u	02 ² ← 10 ⁰	Baw90	2854.191 (10)*	^f R(5, 4) _u	11 ¹ ← 01 ¹	Baw90
2579.748 (10)*	Q(3, 4)	02 ² ← 01 ¹	Baw90	2709.479 (10)	R(3, 1) _u	12 ² ← 11 ¹	Baw90	2862.151 (10)*	R(4, 4) ^u	11 ¹ ← 10 ⁰	Baw90
2581.184 (10)*	Q(5, 4) ^u	01 ¹ ← 00 ⁰	Maj87	2713.789 (10)	R(4, 5)	12 ⁰ ← 11 ¹	Baw90	2864.369 (10)	R(4, 2) ^u	11 ¹ ← 10 ⁰	Baw90
2582.909 (10)*	Q(4, 2) ^u	01 ¹ ← 00 ⁰	Baw90	2715.559 (10)	R(6, 7)	02 ⁰ ← 01 ¹	Baw90	2868.040 (10)	R(4, 1) ^u	11 ¹ ← 10 ⁰	Baw90
2583.155 (10)*	Q(4, 4)	02 ² ← 01 ¹	Baw90	2715.827 (10)	R(3, 3) ^u	03 ¹ ← 02 ⁰	Baw90 [†]	2868.404 (10)*	R(3, 1) _u	02 ² ← 01 ¹	Baw90
2586.985 (10)*	Q(6, 5) ^u	01 ¹ ← 00 ⁰	Maj87	2718.262 (10)*	R(1, 2)	02 ² ← 01 ¹	Baw90	2869.535 (10)	R(9, 10)	02 ⁰ ← 01 ¹	Baw90
2589.541 (10)*	Q(4, 1) ^u	01 ¹ ← 00 ⁰	Baw90	2719.437 (10)* [†]	⁻⁶ Q(2, 4)	03 ³ ← 02 ²	Baw90	2870.890 (10)*	R(3, 0)	02 ² ← 01 ¹	Baw90
2590.071 (10)*	R(2, 0)	12 ² ← 11 ¹	Baw90	2724.058 (10)*	R(3, 2) _u	02 ² ← 01 ¹	Baw90	2884.148 (10)*	^f Q(3, 0)	11 ¹ ← 01 ¹	Xu92
2590.315 (10)*	Q(6, 5) _u	02 ² ← 01 ¹	Baw90	2725.342 (10)*	R(3, 0)	03 ¹ ← 02 ⁰	Baw90	2889.052 (10)*	R(4, 1) _f	01 ¹ ← 00 ⁰	Baw90
2591.323 (10)*	Q(7, 6) ^u	01 ¹ ← 00 ⁰	Baw90	2725.898 (05)*	R(1, 0)	01 ¹ ← 00 ⁰	McK98	2890.993 (10)*	^f R(4, 3) _u	21 ¹ ← 11 ¹	Baw90
2593.460 (10)*	Q(5, 3) ^u	01 ¹ ← 00 ⁰	Maj87	2726.220 (05)*	R(1, 1) ^u	01 ¹ ← 00 ⁰	McK98	2891.867 (10)*	R(4, 2) _f	01 ¹ ← 00 ⁰	Wat84
2594.477 (10)*	Q(8, 7) ^u	01 ¹ ← 00 ⁰	Baw90	2730.887 (10)*	R(2, 1) _f	02 ² ← 01 ¹	Baw90	2893.103 (10)	R(5, 0)	12 ² ← 11 ¹	Baw90
2595.880 (10) [†]	R(6, 6) _f	03 ¹ ← 02 ⁰	Baw90	2733.639 (10)*	^f R(8, 7) _u	11 ¹ ← 01 ¹	Baw90	2893.369 (10)*	R(5, 4) _u	02 ² ← 01 ¹	Baw90
2596.520 (20)	R(4, 5)	02 ⁰ ← 01 ¹	Baw90	2734.526 (10)	R(5, 2) _u	02 ⁰ ← 01 ¹	Baw90	2894.488 (10)*	R(4, 4) _f	01 ¹ ← 00 ⁰	Oka81
2596.520 (20)	Q(4, 5)	02 ² ← 01 ¹	Baw90	2735.515 (10)	R(3, 2) _u	12 ² ← 11 ¹	Baw90	2894.610 (10)*	R(4, 3) _f	01 ¹ ← 00 ⁰	Oka81
2597.058 (10)*	R(4, 3) _u	02 ⁰ ← 01 ¹	Baw90	2737.851 (10)*	R(4, 3) _u	02 ² ← 01 ¹	Baw90	2895.600 (10)	R(7, 7) _f	11 ¹ ← 10 ⁰	Baw90
2597.702 (10)	Q(10, 9) ^u	01 ¹ ← 00 ⁰	Baw90	2740.568 (10)*	^f R(5, 4) _u	11 ¹ ← 01 ¹	Baw90	2895.874 (10)*	R(3, 2) _u	02 ² ← 01 ¹	Baw90
2599.268 (10)*	Q(5, 5)	02 ² ← 01 ¹	Baw90	2742.697 (10)*	R(6, 6)	02 ⁰ ← 01 ¹	Baw90	2896.161 (10)*	^f R(4, 3) _u	11 ¹ ← 01 ¹	Baw90
2600.886 (20)*	Q(8, 6) _u	02 ² ← 01 ¹	Baw90 [†]	2743.418 (10)	R(4, 4) _f	11 ¹ ← 10 ⁰	Baw90	2898.614 (10)*	R(7, 6) _f	11 ¹ ← 10 ⁰	Baw90
2602.367 (10)*	Q(3, 2) _u	02 ² ← 01 ¹	Baw90	2744.586 (10)	R(4, 2) _f	11 ¹ ← 10 ⁰	Baw90	2901.516 (10)	R(9, 9)	02 ⁰ ← 01 ¹	Baw90
2603.883 (10)*	R(0, 1)	02 ² ← 01 ¹	Baw90	2744.719 (10)	R(5, 4) _u	02 ⁰ ← 01 ¹	Baw90	2901.653 (10)	R(8, 8) ^u	11 ¹ ← 10 ⁰	Baw90
2605.063 (10)*	Q(6, 4) ^u	01 ¹ ← 00 ⁰	Baw90	2745.307 (10)	Q(6, 3) _f	02 ² ← 01 ¹	Baw90	2902.523 (10)	R(5, 4) _f	12 ² ← 11 ¹	Baw90
2605.763 (10)*	Q(4, 1) _f	02 ² ← 01 ¹	Baw90	2747.457 (10)*	R(4, 3) _f	11 ¹ ←					

TABLE 5—Continued

Frequency ^a (cm ⁻¹)	Assignment ^b		Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b		Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b		Ref ^c
	Label	Band			Label	Band			Label	Band	
2934.155 (10)*	R(3, 3)	02 ² ← 01 ¹	Baw90	3042.578 (10)	R(6, 4) ^u	11 ¹ ← 10 ⁰	Xu92	3152.951 (05)	R(6, 2) ^u	02 ² ← 01 ¹	Lin01
2934.355 (10)* [†]	R(3, 3)	03 ³ ← 02 ²	Baw90	3046.045 (05)	R(5, 1) ^l	02 ² ← 01 ¹	Lin01	3159.015 (05)	R(9, 8) ^l	01 ¹ ← 00 ⁰	Lin01
2938.491 (10)*	R(5, 1) ^l	01 ¹ ← 00 ⁰	Xu92	3050.552 (05)*	R(8, 4) ^l	01 ¹ ← 00 ⁰	Lin01	3160.236 (05)	R(10, 7) ^l	01 ¹ ← 00 ⁰	Lin01
2941.187 (10)*	R(7, 6) ^l	02 ² ← 01 ¹	Xu92	3051.407 (05)	R(6, 5) ^l	02 ² ← 01 ¹	Lin01	3162.430 (05)	R(10, 6) ^l	01 ¹ ← 00 ⁰	Lin01
2942.209 (10)*	R(5, 2) ^l	01 ¹ ← 00 ⁰	Maj87	3052.077 (05)	R(5, 1) ^u	03 ³ ← 02 ²	Lin01	3163.198 (05)	R(5, 1) ^l	11 ¹ ← 01 ¹	Lin01
2944.828 (10)*	R(4, 0)	02 ² ← 01 ¹	Baw90	3053.355 (05)	R(6, 1) ^u	11 ¹ ← 10 ⁰	Lin01	3167.596 (05)	R(9, 9) ^l	01 ¹ ← 00 ⁰	Lin01
2949.555 (10)*	R(5, 3) ^l	01 ¹ ← 00 ⁰	Maj87	3053.562 (05)	R(10, 9) ^l	02 ⁰ ← 01 ¹	Lin01	3172.045 (05)	R(8, 7) ^l	02 ² ← 01 ¹	Lin01
2950.605 (10)*	R(4, 1) ^l	02 ² ← 01 ¹	Xu92	3056.252 (05)*	R(7, 5) ^l	01 ¹ ← 00 ⁰	Lin01	3175.189 (05)	R(7, 5) ^l	02 ² ← 01 ¹	Lin01
2951.438 (20)	R(5, 3) ^u	11 ¹ ← 10 ⁰	Carbo	3059.381 (10)	R(9, 5) ^l	01 ¹ ← 00 ⁰	Xu92	3177.167 (05)*	R(10, 8) ^l	01 ¹ ← 00 ⁰	Lin01
2953.405 (10)*	R(4, 1) ^u	02 ² ← 01 ¹	Xu92	3059.512 (10)	ⁿ P(3, 4) ^l	11 ¹ ← 01 ¹	Xu92	3177.167 (05)*	R(7, 3) ^u	11 ¹ ← 10 ⁰	Lin01
2955.154 (10)*	R(5, 4) ^l	01 ¹ ← 00 ⁰	Uy94	3060.507 (05)*	R(5, 0)	02 ² ← 01 ¹	Lin01	3177.628 (05)	ⁿ R(7, 5) ^u	02 ² ← 10 ⁰	Lin01
2956.072 (10)*	R(5, 5) ^l	01 ¹ ← 00 ⁰	Oka81	3061.287 (10)*	R(4, 6)	03 ³ ← 02 ²	Xu92	3179.115 (05)*	^l R(6, 3) ^u	11 ¹ ← 01 ¹	Lin01
2956.222 (10)	⁻⁶ R(6, 2) ^u	01 ¹ ← 00 ⁰	Uy94	3062.110 (05)*	R(6, 6) ^u	11 ¹ ← 10 ⁰	Lin01	3179.998 (05)*	R(6, 4) ^u	02 ² ← 01 ¹	Lin01
2956.843 (30)	^l R(7, 5)	10 ⁰ ← 00 ⁰	Xu92	3062.813 (10)*	ⁿ R(0, 1)	11 ¹ ← 01 ¹	Xu92	3182.038 (05)	R(6, 6) ^u	01 ¹ ← 00 ⁰	Lin01
2956.947 (30)*	R(3, 2) ^l	02 ² ← 01 ¹	Xu92	3063.078 (10)	R(11, 6) ^l	11 ¹ ← 00 ⁰	Xu92	3182.281 (05)*	^l R(5, 2) ^u	11 ¹ ← 01 ¹	Lin01
2962.822 (10)*	R(5, 3) ^l	02 ² ← 01 ¹	Xu92	3063.273 (10)	R(6, 3) ^u	03 ¹ ← 02 ⁰	Xu92	3187.488 (05)	R(11, 8) ^l	01 ¹ ← 00 ⁰	Lin01
2964.705 (10)	R(5, 0)	11 ¹ ← 10 ⁰	Xu92	3063.935 (05)	^l R(6, 2) ^l	11 ¹ ← 01 ¹	Lin01	3188.423 (05)	R(6, 2) ^l	02 ² ← 01 ¹	Lin01
2964.987 (10)*	^l R(5, 3) _u	11 ¹ ← 01 ¹	Xu92	3064.356 (05)*	R(7, 6) ^l	01 ¹ ← 00 ⁰	Lin01	3193.232 (05)*	R(6, 5) ^u	01 ¹ ← 00 ⁰	Lin01
2966.864 (10)	R(6, 4) _u	02 ² ← 01 ¹	Xu92	3064.356 (05)*	R(5, 2) _l	02 ² ← 01 ¹	Lin01	3194.796 (05)*	^l R(6, 0) ^l	11 ¹ ← 01 ¹	Lin01
2974.534 (20)*	^l R(2, 1) _u	11 ¹ ← 01 ¹	Xu92	3065.578 (05)*	R(5, 2) _u	02 ² ← 01 ¹	Lin01	3199.631 (05)*	ⁿ R(7, 7) ^l	02 ² ← 10 ⁰	Lin01
2974.682 (20)*	ⁿ Q(1, 1)	11 ¹ ← 01 ¹	Xu92	3065.777 (05)*	R(5, 1) _u	02 ² ← 01 ¹	Lin01	3200.723 (05)	R(10, 9) ^l	01 ¹ ← 00 ⁰	Lin01
2975.656 (10)*	^l R(5, 3)	12 ⁰ ← 02 ²	Xu92	3066.565 (05)*	^l R(5, 3)	10 ⁰ ← 00 ⁰	Lin01	3201.386 (05)	R(7, 1) _l	02 ² ← 01 ¹	Lin01
2976.080 (10)	R(8, 7) _u	02 ² ← 01 ¹	Xu92	3067.733 (05)*	R(4, 2) _l	02 ² ← 01 ¹	Lin01	3201.672 (05)	R(6, 5) _u	02 ² ← 01 ¹	Lin01
2976.566 (10)*	R(4, 2) _u	02 ² ← 01 ¹	Xu92 [‡]	3069.176 (05)*	R(7, 7) ^l	01 ¹ ← 00 ⁰	Lin01	3202.174 (05)	^l R(5, 2)	10 ⁰ ← 00 ⁰	Lin01
2977.488 (10) [†]	R(8, 7) _l	02 ⁰ ← 01 ¹	Xu92	3076.175 (10)	R(5, 3) _u	03 ³ ← 02 ²	Lin01	3203.158 (10)*	^l R(4, 1) _l	11 ¹ ← 01 ¹	Xu92 [‡]
2979.325 (10)	^l R(6, 4)	11 ¹ ← 01 ¹	Xu92	3077.457 (10) [†]	^l R(6, 3)	20 ⁰ ← 10 ⁰	Lin01	3203.513 (10)	R(8, 6) ^u	11 ¹ ← 10 ⁰	Xu92
2979.507 (10)	R(6, 1) ^l	01 ¹ ← 00 ⁰	Xu92	3078.892 (05)	R(9, 3) ^l	01 ¹ ← 00 ⁰	Lin01	3205.308 (05)	R(6, 4) ^u	01 ¹ ← 00 ⁰	Lin01
2979.658 (10)*	ⁿ R(6, 6) ^l	02 ² ← 10 ⁰	Xu92	3085.617 (05)*	^l R(5, 3) _l	11 ¹ ← 01 ¹	Lin01	3209.072 (05)*	R(11, 9) ^l	01 ¹ ← 00 ⁰	Lin01
2980.327 (10)	R(4, 3) _u	02 ² ← 01 ¹	Carbo	3086.072 (05)*	^l R(4, 2) _l	11 ¹ ← 01 ¹	Lin01	3210.543 (05)	R(12, 9) ^l	01 ¹ ← 00 ⁰	Lin01
2984.082 (10)*	R(6, 2) ^l	01 ¹ ← 00 ⁰	Xu92	3091.891 (10)*	ⁿ Q(2, 2) _l	11 ¹ ← 01 ¹	Xu92	3210.801 (05)	R(10, 10) ^l	01 ¹ ← 00 ⁰	Lin01
2984.259 (10)*	^l R(4, 3) _l	11 ¹ ← 01 ¹	Xu92	3092.324 (10)*	ⁿ Q(1, 2)	11 ¹ ← 01 ¹	Xu92 [‡]	3212.252 (05)	R(6, 2) ^u	01 ¹ ← 00 ⁰	Lin01
2985.494 (10)*	R(6, 3) ^l	01 ¹ ← 00 ⁰	Xu92	3093.669 (05)	R(7, 7) ^u	11 ¹ ← 10 ⁰	Lin01	3214.612 (05)	R(6, 6)	02 ² ← 01 ¹	Lin01
2989.507 (20)*	R(6, 4) ^l	01 ¹ ← 00 ⁰	Xu92	3096.416 (05)*	R(5, 5) ^u	01 ¹ ← 00 ⁰	Lin01	3216.361 (05)*	R(6, 3) _u	01 ¹ ← 00 ⁰	Lin01
2989.507 (30)*	^l R(3, 2) _l	11 ¹ ← 01 ¹	Xu92	3096.665 (05)*	R(6, 3) _l	02 ² ← 01 ¹	Lin01	3219.108 (05)	^l R(7, 3)	10 ⁰ ← 00 ⁰	Lin01
2989.618 (10)	^l R(4, 0)	12 ⁰ ← 02 ²	Xu92	3097.259 (05)*	^l R(2, 0)	11 ¹ ← 01 ¹	Lin01	3220.181 (05)	R(7, 0)	02 ² ← 01 ¹	Lin01
2990.585 (10)* [†]	^l R(5, 2) _l	11 ¹ ← 01 ¹	Xu92	3097.985 (05)*	R(5, 4) _u	02 ² ← 01 ¹	Lin01	3220.816 (05)	R(6, 1) ^u	01 ¹ ← 00 ⁰	Lin01
2993.467 (10)*	R(7, 5) _u	02 ² ← 01 ¹	Xu92	3099.905 (05)*	R(8, 6) ^l	01 ¹ ← 00 ⁰	Lin01	3221.086 (05)	ⁿ Q(2, 3)	11 ¹ ← 01 ¹	Lin01
2994.903 (10)*	^l R(4, 2) _u	11 ¹ ← 01 ¹	Xu92	3100.131 (05)*	^l R(7, 3)	20 ⁰ ← 10 ⁰	Lin01	3221.214 (05)*	R(7, 1) _u	02 ² ← 01 ¹	Lin01
2998.347 (15)*	R(11, 7) _l	02 ² ← 01 ¹	Lin01	3100.871 (05)	ⁿ R(7, 6) ^l	02 ² ← 10 ⁰	Lin01	3222.022 (05)	R(5, 3) _l	02 ² ← 01 ¹	Lin01
3000.105 (10)	R(8, 6) _l	11 ¹ ← 01 ¹	Xu92	3101.397 (05)	R(5, 3) _u	02 ² ← 01 ¹	Lin01	3228.754 (05)*	^l R(3, 0) _u	11 ¹ ← 01 ¹	Lin01
3002.355 (10)	R(10, 10) ^l	11 ¹ ← 10 ⁰	Xu92	3102.368 (05)*	R(8, 5) ^l	01 ¹ ← 00 ⁰	Lin01	3235.574 (05)*	R(6, 7)	02 ² ← 01 ¹	Lin01
3002.750 (10)	ⁿ R(8, 8)	02 ⁰ ← 10 ⁰	Xu92	3102.736 (05)*	^l R(3, 1) _l	11 ¹ ← 01 ¹	Lin01	3235.813 (05)	R(12, 10) ^l	01 ¹ ← 00 ⁰	Lin01
3003.253 (05)	R(5, 3) _u	03 ¹ ← 02 ⁰	Lin01	3103.873 (05)*	R(6, 0)	02 ² ← 01 ¹	Lin01	3236.270 (05)	R(7, 4) _l	02 ² ← 01 ¹	Lin01
3006.996 (05)	R(5, 4) _l	01 ¹ ← 00 ⁰	Lin01	3106.804 (05)	R(5, 4) ^u	01 ¹ ← 00 ⁰	Lin01	3238.614 (05)*	R(11, 10) ^l	01 ¹ ← 00 ⁰	Lin01
3008.108 (05)*	R(4, 4) ^u	01 ¹ ← 00 ⁰	Lin01	3108.871 (05)*	R(7, 6) _l	02 ² ← 01 ¹	Lin01	3238.662 (05)	R(9, 8) _l	02 ² ← 01 ¹	Lin01
3009.317 (05)*	^l R(2, 1) _l	11 ¹ ← 01 ¹	Lin01	3110.877 (10)*	ⁿ P(5, 6) ^l	11 ¹ ← 01 ¹	Lin01	3240.385 (05)*	R(8, 6) _l	02 ² ← 01 ¹	Lin01
3011.509 (05)*	R(6, 5) ^l	01 ¹ ← 00 ⁰	Lin01	3113.838 (10)	ⁿ R(9, 9)	02 ⁰ ← 10 ⁰	Lin01	3247.272 (05)	^l R(8, 2) _u	11 ¹ ← 01 ¹	Lin01
3014.364 (05)*	R(6, 6) ^l	01 ¹ ← 00 ⁰	Lin01	3113.532 (05)*	R(8, 7) ^l	01 ¹ ← 00 ⁰	Lin01	3247.694 (05)	R(7, 2) _u	02 ² ← 01 ¹	Lin01
3015.241 (05)*	R(4, 3) ^u	01 ¹ ← 00 ⁰	Lin01	3115.615 (05)*	R(5, 5)	02 ² ← 01 ¹	Lin01	3247.800 (05)	^l R(6, 1) _l	11 ¹ ← 01 ¹	Lin01
3017.629 (05)	R(5, 0)	03 ³ ← 02 ²	Lin01	3118.511 (05)	R(6, 4) _l	02 ² ← 01 ¹	Lin01	3247.891 (05)	R(7, 4) _u	02 ² ← 01 ¹	Lin01
3018.586 (05)	R(9, 8) _l	02 ⁰ ← 01 ¹	Lin01	3120.210 (05)	^l R(4, 2)	10 ⁰ ← 00 ⁰	Lin01	3249.704 (05)	R(11, 11) ^l	01 ¹ ← 00 ⁰	Lin01
3020.495 (05)	R(4, 4)	02 ² ← 01 ¹	Lin01	3120.321 (05)*	R(8, 8) ^l	01 ¹ ← 00 ⁰	Lin01	3249.794 (05)	R(7, 3) _u	02 ² ← 01 ¹	Lin01
3021.862 (05)*	R(6, 4) _u	12 ² ← 11 ¹	Lin01	3121.216 (05)	^l R(4, 0) ^l	11 ¹ ← 01 ¹	Lin01	3259.835 (05)*	R(7, 3) _l	02 ² ← 01 ¹	Lin01
3022.424 (05)*	^l R(5, 1) _u	11 ¹ ← 01 ¹	Lin01	3121.814 (05)*	R(5, 3) ^u	01 ¹ ← 00 ⁰	Lin01	3261.336 (05)	R(8, 5) _u	02 ² ← 01 ¹	Lin01
3023.674 (10)	R(6, 3) ^u	11 ¹ ← 10 ⁰	Xu92	3122.252 (05)*	R(5, 2) ^u	01 ¹ ← 00 ⁰	Lin01	3265.138 (05)	R(7, 7) ^u	01 ¹ ← 00 ⁰	Lin01
3024.439 (10)	^l R(8, 5)	10 ⁰ ← 00 ⁰	Lin01	3128.068 (05)*	R(5, 1) ^u	01 ¹ ← 00 ⁰	Lin01	3265.304 (05)	R(8, 1) _u	02 ² ← 01 ¹	Lin01
3024.558 (10)*	R(4, 2) ^u	01 ¹ ← 00 ⁰	Lin01	3128.296 (05)	R(7, 2) ^u	11 ¹ ← 10 ⁰	Lin01	3266.017 (05)	R(7, 5) _u	02 ² ← 01 ¹	Lin01
3025.941 (05)*	R(7, 4) ^l	01 ¹ ← 00 ⁰	Lin01	3129.516 (10)	R(6, 1) _l	02 ² ← 01 ¹	Lin01	3269.095 (05)	ⁿ R(8, 5) ^u	02 ² ← 10 ⁰	Lin01
3026.162 (05)*	^l R(6, 4)	10 ⁰ ← 00 ⁰	Lin01	3129.811 (05)	R(5, 0)	01 ¹ ← 00 ⁰	Lin01	3269.496 (05)	R(8, 3) ^u	11 ¹ ← 10 ⁰	Lin01
3028.543 (05)*	R(8, 6) _u	02 ² ← 01 ¹	Lin01	3130.216 (05)*	R(9, 6) ^l	01 ¹ ←					

TABLE 5—Continued

Frequency ^a (cm ⁻¹)	Assignment ^b Label	Band	Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b Label	Band	Ref ^c	Frequency ^a (cm ⁻¹)	Assignment ^b Label	Band	Ref ^c
3292.521 (05)	<i>R</i> (7, 2) _i	02 ² ← 01 ¹	Lin01	3473.764 (05)	<i>R</i> (10, 6) _u	02 ² ← 01 ¹	Lin01	4900.393 (10)*	[†] <i>R</i> (3, 3)	02 ² ← 00 ⁰	Xu90
3293.790 (05)	<i>R</i> (9, 6) _u	11 ¹ ← 10 ⁰	Lin01	3476.189 (05)	<i>R</i> (9, 4) _u	01 ¹ ← 00 ⁰	Lin01	4907.871 (10)*	[†] <i>Q</i> (1, 0)	02 ² ← 00 ⁰	Xu90
3296.014 (05)	<i>R</i> (9, 9) _u	11 ¹ ← 10 ⁰	Lin01	3486.049 (05)	<i>R</i> (9, 9)	02 ² ← 01 ¹	Lin01	4908.672 (20)*	[†] <i>P</i> (4, 3)	02 ² ← 00 ⁰	Xu90
3298.990 (05)	<i>R</i> (8, 3) _u	02 ² ← 01 ¹	Lin01	3497.971 (05)	<i>R</i> (9, 3) _u	01 ¹ ← 00 ⁰	Lin01	4914.248 (10)*	[†] <i>Q</i> (3, 0)	02 ² ← 00 ⁰	Xu90
3300.111 (10)*	[†] <i>R</i> (5, 0) _u	11 ¹ ← 01 ¹	Lin01	3498.764 (05)	⁻⁶ <i>R</i> (9, 5) _i	01 ¹ ← 00 ⁰	Lin01	4930.981 (20)*	[†] <i>P</i> (6, 5) _i	02 ² ← 00 ⁰	Maj89
3301.694 (05)	⁻⁶ <i>R</i> (8, 5) _i	02 ⁰ ← 01 ¹	Lin01	3499.417 (05)	<i>R</i> (10, 10) _u	01 ¹ ← 00 ⁰	Lin01	4931.596 (20)*	[†] <i>R</i> (6, 5)	02 ² ← 00 ⁰	Xu90
3302.423 (05)*	<i>R</i> (7, 4) _u	01 ¹ ← 00 ⁰	Lin01	3503.306 (10)	<i>R</i> (10, 9) _u	01 ¹ ← 00 ⁰	Lin01	4936.000 (20)*	[†] <i>R</i> (2, 2)	02 ² ← 00 ⁰	Xu90
3303.093 (05)*	<i>R</i> (13, 12) _i	01 ¹ ← 00 ⁰	Lin01	3513.541 (05)	<i>R</i> (10, 8) _u	01 ¹ ← 00 ⁰	Lin01	4955.991 (10)*	[†] <i>P</i> (2, 8) _u	02 ² ← 00 ⁰	Xu90
3305.935 (05)	<i>R</i> (7, 1) _u	01 ¹ ← 00 ⁰	Lin01	3516.951 (05)	<i>R</i> (9, 2) _u	01 ¹ ← 00 ⁰	Lin01	4966.838 (20)*	[†] <i>R</i> (5, 4)	02 ² ← 00 ⁰	Xu90
3307.150 (05)	<i>R</i> (10, 9) _i	02 ² ← 01 ¹	Lin01	3521.044 (05)	<i>R</i> (9, 10)	02 ² ← 01 ¹	Lin01	4968.272 (10)*	[†] <i>R</i> (1, 1)	02 ² ← 00 ⁰	Xu90
3308.650 (10)	<i>R</i> (9, 7) _i	02 ² ← 01 ¹	Lin01	3523.742 (05)	[†] <i>R</i> (8, 1)	10 ⁰ ← 00 ⁰	Lin01	4971.561 (10)*	[†] <i>P</i> (3, 3)	02 ² ← 00 ⁰	Xu90
3308.685 (05)	<i>R</i> (7, 0)	01 ¹ ← 00 ⁰	Lin01	3523.998 (05)	<i>R</i> (10, 7) _u	01 ¹ ← 00 ⁰	Lin01	4975.338 (20)*	[†] <i>P</i> (6, 6) _i	02 ² ← 00 ⁰	Xu90
3309.924 (05)	<i>R</i> (7, 7)	02 ² ← 01 ¹	Lin01	3527.047 (05)	<i>R</i> (10, 9) _u	02 ² ← 01 ¹	Lin01	5000.499 (10)*	[†] <i>R</i> (4, 3)	02 ² ← 00 ⁰	Xu90
3311.009 (05)	<i>R</i> (8, 0)	02 ² ← 01 ¹	Lin01	3531.279 (05)	<i>R</i> (10, 6) _u	01 ¹ ← 00 ⁰	Lin01	5023.496 (10)*	[†] <i>Q</i> (1, 1)	02 ² ← 00 ⁰	Xu90
3313.752 (05)	<i>R</i> (13, 13) _i	01 ¹ ← 00 ⁰	Lin01	3546.576 (05)	<i>R</i> (10, 5) _u	01 ¹ ← 00 ⁰	Lin01	5029.071 (10)*	[†] <i>Q</i> (2, 1)	02 ² ← 00 ⁰	Xu90
3317.786 (05)	[†] <i>R</i> (8, 1) _u	11 ¹ ← 01 ¹	Lin01	3551.579 (15)	<i>R</i> (10, 3) _u	01 ¹ ← 00 ⁰	Lin01	5032.447 (10)*	[†] <i>R</i> (3, 2)	02 ² ← 00 ⁰	Xu90
3321.010 (05)	[†] <i>R</i> (7, 3) _u	01 ¹ ← 00 ⁰	Lin01	3552.313 (15)	<i>R</i> (10, 4) _u	01 ¹ ← 00 ⁰	Lin01	5054.742 (100)*	[†] <i>Q</i> (4, 1) _u	02 ² ← 00 ⁰	Maj89
3325.674 (10)	[†] <i>R</i> (5, 1)	10 ⁰ ← 00 ⁰	Lin01	3553.705 (15)	<i>R</i> (11, 0)	01 ¹ ← 00 ⁰	Lin01	5061.882 (20)*	[†] <i>R</i> (2, 1)	02 ² ← 00 ⁰	Xu90
3328.773 (05)	<i>R</i> (8, 3) _i	02 ² ← 01 ¹	Lin01	3571.295 (15)	<i>R</i> (11, 10) _u	01 ¹ ← 00 ⁰	Lin01	5094.218 (20)*	[†] <i>R</i> (1, 0)	02 ² ← 00 ⁰	Xu90
3329.924 (05)	<i>R</i> (14, 13) _i	01 ¹ ← 00 ⁰	Lin01	3572.419 (15)	<i>R</i> (11, 11) _u	01 ¹ ← 00 ⁰	Lin01	6806.665 (70)*	<i>P</i> (3, 1) _u	03 ¹ ← 00 ⁰	Ven94
3331.374 (05)	<i>R</i> (8, 4) _i	02 ² ← 01 ¹	Lin01	3574.750 (15)	[†] <i>R</i> (7, 0)	10 ⁰ ← 00 ⁰	Lin01	6807.297 (70)*	<i>P</i> (3, 3)	03 ¹ ← 00 ⁰	Ven94
3331.571 (05)	<i>R</i> (8, 5) _u	02 ² ← 01 ¹	Lin01	3579.301 (15)	<i>R</i> (11, 9) _u	01 ¹ ← 00 ⁰	Lin01	6807.724 (70)*	<i>P</i> (3, 2)	03 ¹ ← 00 ⁰	Ven94
3332.520 (05)	<i>R</i> (7, 8)	02 ² ← 01 ¹	Lin01	3586.139 (15)	<i>R</i> (10, 2) _u	01 ¹ ← 00 ⁰	Lin01	6811.218 (200)*	<i>P</i> (3, 0)	03 ¹ ← 00 ⁰	Ven94
3338.534 (05)	<i>R</i> (14, 14) _i	01 ¹ ← 00 ⁰	Lin01	3588.381 (15)	<i>R</i> (11, 8) _u	01 ¹ ← 00 ⁰	Lin01	6865.731 (70)*	<i>P</i> (2, 1)	03 ¹ ← 00 ⁰	Ven94
3343.327 (05)	[†] <i>R</i> (9, 3)	10 ⁰ ← 00 ⁰	Lin01	3596.217 (15)	<i>R</i> (11, 7) _u	01 ¹ ← 00 ⁰	Lin01	6866.340 (70)*	<i>Q</i> (5, 0)	03 ¹ ← 00 ⁰	Ven94
3345.710 (05)	<i>R</i> (8, 8) _u	01 ¹ ← 00 ⁰	Lin01	3642.547 (10)	<i>R</i> (12, 12) _u	01 ¹ ← 00 ⁰	Maj94	6877.546 (70)*	<i>P</i> (2, 2)	03 ¹ ← 00 ⁰	Ven94
3348.845 (05)	<i>R</i> (9, 6) _i	02 ² ← 01 ¹	Lin01	4434.861 (10)*	[†] <i>Q</i> (5, 4)	02 ² ← 00 ⁰	Maj94	6883.091 (70)*	<i>Q</i> (5, 3) _i	03 ¹ ← 00 ⁰	Ven94
3355.517 (05)*	<i>R</i> (8, 7) _u	01 ¹ ← 00 ⁰	Lin01	4465.095 (10)*	[†] <i>Q</i> (6, 4)	02 ² ← 00 ⁰	Maj94	6891.619 (70)*	<i>Q</i> (4, 3) _i	03 ¹ ← 00 ⁰	Ven94
3356.747 (05)*	<i>R</i> (8, 2) _u	01 ¹ ← 00 ⁰	Lin01	4539.759 (20)*	[†] <i>P</i> (5, 0)	02 ² ← 00 ⁰	Maj89	7144.212 (70)*	<i>R</i> (1, 1) _i	03 ¹ ← 00 ⁰	Ven94
3358.400 (05)	<i>R</i> (15, 15) _i	01 ¹ ← 00 ⁰	Lin01	4553.340 (10)*	[†] <i>Q</i> (4, 3) _u	03 ³ ← 01 ¹	Maj94	7192.908 (70)*	<i>R</i> (2, 2) _i	03 ¹ ← 00 ⁰	Ven94
3362.256 (10)	[†] <i>R</i> (7, 2)	10 ⁰ ← 00 ⁰	Lin01	4557.020 (20)*	[†] <i>Q</i> (4, 3)	02 ² ← 00 ⁰	Xu90	7234.957 (70)*	<i>R</i> (3, 3) _i	03 ¹ ← 00 ⁰	Ven94
3368.118 (05)*	<i>R</i> (8, 6) _u	01 ¹ ← 00 ⁰	Lin01	4557.731 (10)*	[†] <i>P</i> (7, 1) _u	02 ² ← 00 ⁰	Maj94	7237.285 (70)*	<i>R</i> (1, 1) _u	03 ¹ ← 00 ⁰	Ven94
3368.560 (05)*	<i>R</i> (8, 7) _u	02 ² ← 01 ¹	Lin01	4578.735 (20)*	[†] <i>Q</i> (5, 3)	02 ² ← 00 ⁰	Xu90	7241.245 (70)*	<i>R</i> (1, 0)	03 ¹ ← 00 ⁰	Ven94
3369.664 (05)	<i>R</i> (9, 5) _i	02 ² ← 01 ¹	Lin01	4607.205 (20)*	[†] <i>Q</i> (6, 3)	02 ² ← 00 ⁰	Maj89	7265.882 (70)*	<i>R</i> (4, 4) _i	03 ¹ ← 00 ⁰	Ven94
3375.003 (05)	<i>R</i> (8, 6) _u	02 ² ← 01 ¹	Lin01	4637.992 (50)*	[†] <i>P</i> (5, 1) _u	02 ² ← 00 ⁰	Maj89	7785.233 (10)*	[†] <i>Q</i> (3, 0)	12 ² ← 00 ⁰	McC00
3376.775 (05)	<i>R</i> (11, 10) _u	02 ² ← 01 ¹	Lin01	4638.331 (10)	[†] <i>R</i> (9, 1)	02 ² ← 00 ⁰	Xu90	7785.701 (10)	[†] <i>Q</i> (1, 0)	12 ² ← 00 ⁰	McC00
3377.047 (05)	[†] <i>R</i> (10, 3)	10 ⁰ ← 00 ⁰	Lin01	4641.987 (20)*	[†] <i>Q</i> (7, 3)	02 ² ← 00 ⁰	Maj89	7789.878 (10)	[†] <i>R</i> (3, 3)	12 ² ← 00 ⁰	McC00
3380.010 (05)	<i>R</i> (8, 5) _u	01 ¹ ← 00 ⁰	Lin01	4661.576 (10)*	[†] <i>P</i> (7, 3)	02 ² ← 00 ⁰	Maj94	7805.893 (10) [†]	[†] <i>P</i> (1, 1)	12 ² ← 00 ⁰	McC00
3381.399 (05)	<i>R</i> (8, 1) _u	01 ¹ ← 00 ⁰	Lin01	4664.306 (10)*	[†] <i>P</i> (3, 0)	02 ² ← 00 ⁰	Xu90	7820.239 (10)*	[†] <i>P</i> (2, 2)	12 ² ← 00 ⁰	McC00
3388.155 (05)	<i>R</i> (8, 2) _i	02 ² ← 01 ¹	Lin01	4677.273 (15)*	[†] <i>Q</i> (3, 2)	02 ² ← 00 ⁰	Xu90	7822.375 (10)	[†] <i>R</i> (2, 2)	12 ² ← 00 ⁰	McC00
3389.119 (05)	<i>R</i> (9, 3) _u	02 ² ← 01 ¹	Lin01	4685.564 (10)	[†] <i>R</i> (8, 8)	02 ² ← 00 ⁰	Xu90	7826.739 (10)	[†] <i>P</i> (3, 3)	12 ² ← 00 ⁰	McC00
3392.547 (05)	<i>R</i> (8, 4) _u	01 ¹ ← 00 ⁰	Lin01	4691.962 (100)	[†] <i>Q</i> (4, 2)	02 ² ← 00 ⁰	Maj89	7833.249 (20)	[†] <i>P</i> (4, 4) _i	12 ² ← 00 ⁰	McC00
3395.752 (05)	[†] <i>R</i> (6, 1)	10 ⁰ ← 00 ⁰	Lin01	4700.139 (20)*	[†] <i>P</i> (8, 1)	02 ² ← 00 ⁰	Maj89	7850.959 (10)	[†] <i>R</i> (1, 1)	12 ² ← 00 ⁰	McC00
3399.510 (05)	<i>R</i> (8, 3) _u	01 ¹ ← 00 ⁰	Lin01	4712.282 (10)*	[†] <i>Q</i> (5, 2)	02 ² ← 00 ⁰	Maj94	7880.921 (10)	[†] <i>R</i> (4, 3)	12 ² ← 00 ⁰	McC00
3399.872 (05)	<i>R</i> (8, 8)	02 ² ← 01 ¹	Lin01	4721.019 (10)*	[†] <i>P</i> (4, 3) _i	03 ³ ← 01 ¹	Maj94	7894.711 (10)	[†] <i>Q</i> (1, 1)	12 ² ← 00 ⁰	McC00
3407.501 (05)	<i>R</i> (9, 6) _u	02 ² ← 01 ¹	Lin01	4732.041 (10)*	[†] <i>R</i> (7, 7)	02 ² ← 00 ⁰	Maj94	7898.371 (10)*	[†] <i>Q</i> (2, 1)	12 ² ← 00 ⁰	McC00
3408.984 (10)	<i>R</i> (10, 7) _i	02 ² ← 01 ¹	Lin01	4735.941 (100)*	[†] <i>Q</i> (6, 2)	02 ² ← 00 ⁰	Maj89	7905.717 (10)*	[†] <i>Q</i> (3, 1)	12 ² ← 00 ⁰	McC00
3411.415 (05)	<i>R</i> (9, 7) _u	02 ² ← 01 ¹	Lin01	4744.767 (10)*	[†] <i>P</i> (5, 2) _u	02 ² ← 00 ⁰	Maj94	7912.047 (10)	[†] <i>R</i> (3, 2)	12 ² ← 00 ⁰	McC00
3411.859 (05)	[†] <i>R</i> (9, 2)	10 ⁰ ← 00 ⁰	Lin01	4766.167 (100)*	[†] <i>P</i> (7, 4) _i	02 ² ← 00 ⁰	Maj94	7939.619 (10)	[†] <i>R</i> (2, 1)	12 ² ← 00 ⁰	McC00
3423.809 (05)	<i>R</i> (9, 9) _u	01 ¹ ← 00 ⁰	Lin01	4771.641 (100)*	[†] <i>P</i> (3, 1)	02 ² ← 00 ⁰	Maj89	7970.413 (10)*	[†] <i>R</i> (1, 0)	12 ² ← 00 ⁰	McC00
3427.667 (05)*	<i>R</i> (8, 9)	02 ² ← 01 ¹	Lin01	4777.226 (10)*	[†] <i>R</i> (6, 6)	02 ² ← 00 ⁰	Xu90	7998.890 (10)	[†] <i>Q</i> (2, 2)	12 ² ← 00 ⁰	McC00
3431.295 (05)	<i>R</i> (9, 8) _u	01 ¹ ← 00 ⁰	Lin01	4795.030 (10)*	[†] <i>Q</i> (2, 1)	02 ² ← 00 ⁰	Maj94	8005.582 (30)	[†] <i>R</i> (4, 2)	12 ² ← 00 ⁰	McC00
3439.825 (05)	<i>R</i> (11, 9) _i	02 ² ← 01 ¹	Lin01	4804.406 (50)*	[†] <i>Q</i> (3, 1)	02 ² ← 00 ⁰	Maj89	8007.410 (10)	[†] <i>Q</i> (3, 2) _u	12 ² ← 00 ⁰	McC00
3441.416 (05)	[†] <i>R</i> (8, 2)	10 ⁰ ← 00 ⁰	Lin01	4805.287 (20)*	[†] <i>P</i> (4, 2) _u	02 ² ← 00 ⁰	Maj89	8022.012 (20)	[†] <i>Q</i> (4, 2) _u	12 ² ← 00 ⁰	McC00
3443.148 (05)	<i>R</i> (9, 7) _u	01 ¹ ← 00 ⁰	Lin01	4814.521 (20)*	[†] <i>P</i> (6, 3)	02 ² ← 00 ⁰	Maj89	8027.840 (20)	[†] <i>R</i> (3, 1)	12 ² ← 00 ⁰	McC00
3443.466 (10)	[†] <i>R</i> (7, 0) _i	11 ¹ ← 01 ¹	Lin01	4816.353 (10)*	[†] <i>P</i> (5, 3)	02 ² ← 00 ⁰	Xu90	8037.673 (10)	[†] <i>R</i> (3, 1) _i	12 ² ← 00 ⁰	McC00
3445.702 (05)	<i>R</i> (9, 1) _u	01 ¹ ← 00 ⁰	Lin01	4818.901 (20)*	[†] <i>Q</i> (4, 1)	02 ² ← 00 ⁰	Maj89	8053.382 (10)*	<i>P</i> (6, 6)	21 ¹ ← 00 ⁰	McC00
3448.014 (05)	[†] <i>R</i> (5, 0)	10 ⁰ ← 00 ⁰									

due to a species other than H_3^+ . Lines without an asterisk had one or more candidate assignments whose frequency and/or intensity difference from theory was too large to allow a confident assignment.

III.3. Construction of Experimentally Determined Energy Levels

One of the goals of this work was to determine the energy levels from experimental transitions. Constructing all of the relationships between the levels is only possible by combining transitions from different bands. For example, the fundamental band (with the selection rule $\Delta G = 0$) can relate individual J levels within a G "stack" to one another with a combination of P , Q , and R transitions. Relating different G stacks requires transitions with a selection rule other than $\Delta G = 0$. This is the case for overtone and forbidden bands which have the selection rule $\Delta G = \pm 3$. Using a combination of the $\nu_2 \leftarrow 0$, $2\nu_2^2 \leftarrow \nu_2$, and $2\nu_2^2 \leftarrow 0$ transitions, Baw90 and Xu90 experimentally determined the first term values of the ground state in 1990.

We wrote a program to automatically extract from the transition data the relative energies of each level. First, combinations of $\nu_2 \leftarrow 0$, $2\nu_2^2 \leftarrow \nu_2$ and $2\nu_2^2 \leftarrow 0$ bands were used to determine as many ground vibrational state energy relationships as possible. Once this was done, the program examined every transition, searching for transitions whose upper or lower level had already been "determined." The other level in the transition was then calculated. This process was then iterated until no additional levels could be determined. Uncertainties in the transitions were added in quadrature and propagated through the calculation. We performed the entire process twice, once with all of the assigned transitions and once with only the transitions that had been confirmed by combination differences. Levels that were calculated in the first list but not in the second list were necessarily determined by only one transition and are susceptible to mistakes in transition assignments.

At this stage, only the relative values of the energies have been determined and an absolute standard is needed. Additionally, the energy differences between between *ortho* ($G = 3n$) and *para* ($G = 3n \pm 1$) levels are not determined because transitions between them are forbidden. In the past (21, 22), the relationship between *ortho* and *para* levels and the offset from the forbidden level (J, G) = (0, 0) were taken from theoretical calculations. To remain independent from calculations, we instead performed a fit on the ground vibrational state to determine the relationship between the *ortho* and *para* levels as well as their relationship to the (0, 0) level. To do this, we initially computed the absolute values of all of the *ortho* and *para* levels assuming that the lowest populated energy level in each set had zero energy. Next, we performed a least-squares fit of every determined energy level in the ground vibrational state to the following modified symmetric top energy level

expression:

$$\begin{aligned}
 E(J, G) = & -E_{1,1} - \delta_{G,3n} E_{o-p} + BJ(J+1) + (C-B)G^2 \\
 & - D_{JJ}J^2(J+1)^2 - D_{JG}J(J+1)G^2 - D_{GG}G^4 \\
 & - \delta_{G,3}(-1)^J h_3 \left\{ \frac{(J+3)!}{(J-3)!} \right\} + H_{JJJ}J^3(J+1)^3 \\
 & + H_{JJG}J^2(J+1)^2G^2 + H_{JGG}J(J+1)G^4 \\
 & + H_{GGG}G^6 + L_{JJJJ}J^4(J+1)^4 \\
 & + L_{JJJG}J^3(J+1)^3G^2 + L_{JJGG}J^2(J+1)^2G^4 \\
 & + L_{JGGG}J(J+1)G^6 + L_{GGGG}G^8. \quad [10]
 \end{aligned}$$

The first fitted parameter, $E_{1,1}$, gives the energy of the lowest populated *ortho* level (1,1) relative to the forbidden (0,0) level, and the second parameter, E_{o-p} , gives the energy separation between the (1,1) and the (1,0) levels (the relationship between *ortho* and *para* levels). Each energy level was weighted by the inverse of its uncertainty for the fit. The results of the fit and the 2σ uncertainties in the parameters are listed in Table 6. With this information, we adjusted the absolute value of the $G = 3n$ and $G = 3n \pm 1$ levels by the fit parameters, defining the nonphysical (0,0) level as zero energy. Please note that expression 10 does not behave properly outside of the energy levels used in the fit. As pointed out in Watson *et al.* (17), the effective Hamiltonian converges very slowly, making extrapolation difficult. One may be able to overcome this problem by using a

TABLE 6
Determined Molecular Constants for the Ground
Vibrational State of H_3^+ ^a

$E_{1,1}^\dagger$	64.1214 (116)
E_{o-p}^\dagger	22.8389 (56)
B	43.5605 (16)
C	20.6158 (20)
D_{JJ}	$4.1400 (63) \times 10^{-2}$
D_{JG}	$-0.7496 (14) \times 10^{-1}$
D_{GG}	$0.3700 (14) \times 10^{-1}$
h_3	$-0.4846 (26) \times 10^{-5}$
H_{JJJ}	$0.6745 (86) \times 10^{-4}$
H_{JJG}	$-0.2919 (28) \times 10^{-3}$
H_{JGG}	$0.4145 (38) \times 10^{-3}$
H_{GGG}	$-0.1942 (29) \times 10^{-3}$
L_{JJJJ}	$-0.1015 (36) \times 10^{-6}$
L_{JJJG}	$0.0769 (15) \times 10^{-5}$
L_{JJGG}	$-0.1964 (27) \times 10^{-5}$
L_{JGGG}	$0.1934 (31) \times 10^{-5}$
L_{GGGG}	$-0.0594 (22) \times 10^{-5}$

^a All values are in units of cm^{-1} . The numbers in parentheses are the 2σ uncertainties in the last digits. See text for a warning about the use of these values.

[†] Coefficients used to adjust the absolute energy of the experimental energy levels. These terms in Eq. (10) should be set to zero when the energy structure is simulated with the other coefficients.

Padè-type expression (57–60) as done in Ref. (17). The energy levels included in the fit, however, do behave properly, justifying our use of expression 10 to determine $E_{1,1}$ and E_{o-p} .

The values of the determined energy levels are listed in Table 3. Levels that were determined using only transitions verified with combination differences are marked with an asterisk. The values in parentheses correspond to the 2σ uncertainty (in the last digits) in the energy of each level due to the uncertainties in the transition frequencies used to construct the level. This can be thought of as the relative uncertainty for each level. There is an additional uncertainty in the systematic shift that must be considered when comparing the absolute energy of each level. The error in the value of the fit parameter $E_{1,1}$ must be included in the uncertainty for every level. The energy values for levels with $G = 3n$ also depend on the fit parameter E_{o-p} which adds an additional uncertainty which must be accounted for. However, the uncertainties in $E_{1,1}$ and E_{o-p} do not affect the calculations of transitions using these energy levels.

IV. APPLICATION OF RESULTS

The comprehensive list of assigned transitions and observed energy levels presented here will find many applications in the theoretical, laboratory, and astrophysical spectroscopy of H₃⁺. In this section, we briefly outline two such applications: the search for the “forbidden” rotational spectrum and the evaluation of theoretical energy level calculations.

IV.1 Forbidden Rotational Transitions

At its potential minimum, H₃⁺ is a perfect equilateral triangle with no dipole moment and consequently does not possess an allowed pure rotational spectrum. However, as pointed out by Pan and Oka (61), the centrifugal distortion of the molecule due to rotation will break its C_3 symmetry and induce a small dipole moment in the plane of the molecule. The resulting dipole moment will give rise to a weak rotational spectrum which obeys the selection rules $\Delta J = 0, \pm 1$ and $\Delta G = \pm 3$. The general theory of forbidden rotational transitions in polyatomic molecules was developed by Watson (62). In the case of a nonpolar molecule like H₃⁺, the rotational transition $|J, G + 3\rangle \leftarrow |J - 1, G\rangle$ can be thought of as arising from a mixing between $|J, G + 3\rangle$ in the ground state and $|J, G\rangle$ in the ν_2 state, which leads to an intensity borrowing from the allowed rovibrational transition $R(J - 1, G)$ of the fundamental band.

The transition dipole of such rotational transitions is proportional to the derivative of the dipole moment with respect to the ν_2 coordinate. This quantity is much larger for H₃⁺ than for other molecules—in fact, the line strengths of H₃⁺ transitions are orders of magnitude larger than those of CH₄, which have been observed in the laboratory. Although the transition dipole moments are small by the usual standards of rotational spectroscopy (most ranging between 1 and 30 mD), they approach the infrared transition moment (158 mD) at higher J levels. With the rapid development of quantum cascade lasers in the far in-

frared, the rotational transitions of H₃⁺ may soon be detected in the laboratory.

These rotational transitions are also of fundamental importance in the relaxation of H₃⁺ in the interstellar medium, where their spontaneous emission lifetimes are shorter than collision times. Black (63) has pointed out that the flux of such H₃⁺ transitions would be orders of magnitude lower than the thermal continuum from warm dust grains, making their detection infeasible. Draine and Woods (64) have suggested that H₃⁺ rotational transitions may be observable in X-ray heated regions such as the starburst galaxy NGC 6240. Black (63, 65) has further suggested that, under the right conditions, the ${}^1R(3, 1)$ transition could become an astrophysical maser.

In order to enable (laboratory and astronomical) searches for the rotational spectrum of H₃⁺, we have estimated the transition frequencies using our experimentally determined energy levels from Table 3. These are given in Table 7, along with the most recent intensity calculations by Neale *et al.* (56).

IV.2. Evaluation of Theoretical Calculations

Variational calculations of H₃⁺ have substantially improved in recent years with the introduction of adiabatic, relativistic, and nonadiabatic corrections to the theory. The experimentally determined energy levels provide a powerful tool to diagnose the behavior of these calculations, and to compare and contrast the different computational approaches. Before doing so, we give a brief overview of the development of the most recent H₃⁺ theoretical calculations.

IV.2.1. Computational Overview

The first calculations to effectively account for non-Born–Oppenheimer behavior did so by taking an *ab initio* potential energy surface (PES) and adjusting its fitting parameters to better match the experimental values. This semi-empirical approach was used by Watson (55) using the Meyer–Botschwina–Burton PES (67) and similarly by Dinelli *et al.* (68) using the Lie and Frye (69) PES. Later, Dinelli, Polyanski, and Tennyson (DPT) (53) introduced a slightly different approach: a new semiempirical surface is built by adding a purely *ab initio* Born–Oppenheimer PES to another surface (which they call the “adiabatic surface”) of the same functional form whose parameters are determined from the fit to experimental data. In their work the PES of Röhse–Kutzelnigg–Jaquet–Klopper (RKJK) (70) was used as the Born–Oppenheimer surface. Energy level calculations using Watson’s spectroscopically determined potential were reported by Majewski *et al.* (Maj94) (27), and Neale *et al.* (56) calculated energy levels using the DVR3D (71) suite from the DPT surface. The transitions calculated from these energy levels proved to be invaluable in the assignment of laboratory spectra.

The first attempt to calculate the adiabatic effects *ab initio* was by Dinelli *et al.* (Din95) (72), who added a mass-dependent function to the RKJK surface, which accounts for the diagonal

TABLE 7

Pure Rotational Transition Frequencies in the Ground Vibrational State Determined from Experimentally Determined Energy Levels

Label ^a	Frequency ^b (cm ⁻¹)	$ \mu_{ij} ^c$ (mD)	A_{ij}^c (s ⁻¹)	Label ^a	Frequency ^b (cm ⁻¹)	$ \mu_{ij} ^c$ (mD)	A_{ij}^c (s ⁻¹)	Label ^a	Frequency ^b (cm ⁻¹)	$ \mu_{ij} ^c$ (mD)	A_{ij}^c (s ⁻¹)
^t R(3, 1)	7.255 (10)	4.23 [†]	2.78×10^{-9}	ⁿ Q(5, 3)	190.756 (13)	17.7	6.80×10^{-4}	ⁿ R(5, 2)	553.791 (19)	9.24	5.37×10^{-3}
^t R(6, 3)	9.261 (13)	14.7 [†]	6.22×10^{-8}	ⁿ Q(3, 3)	201.524 (09)	5.37	7.40×10^{-5}	ⁿ Q(7, 6)	555.500 (14)	13.6	9.93×10^{-3}
ⁿ P(8, 7)	29.655 (18)	22.3	3.59×10^{-6}	^t R(7, 2)	220.891 (25)	22.7	1.98×10^{-3}	ⁿ R(7, 1)	568.013 (34)	20.0	2.59×10^{-2}
ⁿ P(5, 5)	39.453 (12)	8.33	1.10×10^{-6}	^t R(6, 1)	261.550 (21)	17.3	1.93×10^{-3}	ⁿ Q(6, 6)	581.450 (11)	6.99	3.00×10^{-3}
^t R(5, 2)	51.347 (16)	11.3	6.38×10^{-6}	ⁿ Q(8, 4)	286.320 (42)	29.3	6.35×10^{-3}	ⁿ R(4, 3)	612.525 (12)	5.86	3.03×10^{-3}
ⁿ Q(8, 2)	56.563 (47)	32.5	5.94×10^{-5}	ⁿ Q(7, 4)	298.423 (24)	22.1	4.05×10^{-3}	ⁿ R(6, 2)	621.074 (25)	13.1	1.47×10^{-2}
ⁿ Q(7, 2)	58.880 (28)	25.6	4.19×10^{-5}	^t R(5, 0)	306.088 (14)	17.3	3.18×10^{-3}	ⁿ Q(8, 7)	666.334 (20)	15.4	2.19×10^{-2}
ⁿ Q(6, 2)	61.101 (21)	19.4	2.68×10^{-5}	ⁿ Q(6, 4)	310.199 (19)	15.5	2.25×10^{-3}	ⁿ R(7, 2)	683.456 (44)	17.4	3.44×10^{-2}
ⁿ Q(5, 2)	63.197 (16)	13.9	1.53×10^{-5}	ⁿ Q(5, 4)	321.347 (15)	9.83	1.00×10^{-6}	ⁿ Q(7, 7)	700.315 (17)	7.95	6.80×10^{-3}
ⁿ Q(4, 2)	65.107 (13)	9.17	7.27×10^{-6}	ⁿ R(2, 2)	325.482 (09)	1.52	3.51×10^{-5}	ⁿ R(6, 3)	743.039 (18)	14.8	3.25×10^{-2}
ⁿ Q(3, 2)	66.758 (11)	5.33	2.64×10^{-6}	ⁿ Q(4, 4)	331.549 (12)	4.97	2.81×10^{-4}	ⁿ R(4, 4)	748.280 (13)	2.16	7.51×10^{-4}
ⁿ Q(2, 2)	68.062 (07)	2.40	5.66×10^{-7}	^t R(7, 1)	338.256 (26)	22.6	7.00×10^{-3}	ⁿ R(5, 4)	811.941 (18)	4.62	4.22×10^{-3}
ⁿ P(6, 6)	84.606 (10)	10.6	1.81×10^{-5}	ⁿ R(4, 1)	353.533 (15)	7.58	9.73×10^{-4}	ⁿ Q(8, 8)	815.622 (20)	8.93	1.35×10^{-2}
^t R(4, 1)	95.383 (14)	8.07	2.16×10^{-5}	ⁿ R(3, 2)	405.563 (12)	3.47	3.23×10^{-4}	ⁿ R(6, 4)	870.172 (23)	7.69	1.41×10^{-2}
^t R(7, 3)	100.112 (15)	21.5	1.65×10^{-4}	ⁿ Q(8, 5)	406.002 (31)	25.7	1.38×10^{-2}	ⁿ R(7, 4)	922.999 (41)	11.3	3.58×10^{-2}
ⁿ R(1, 1)	105.173 (04)	0.84	4.26×10^{-7}	ⁿ Q(7, 5)	423.844 (24)	18.3	8.00×10^{-3}	ⁿ R(5, 5)	950.783 (17)	2.34	1.75×10^{-3}
ⁿ P(7, 7)	128.566 (15)	13.1	9.85×10^{-5}	ⁿ R(5, 1)	429.493 (19)	11.1	3.62×10^{-3}	ⁿ R(6, 5)	1003.537 (23)	4.98	9.00×10^{-3}
^t R(6, 2)	138.350 (20)	16.8	2.70×10^{-4}	ⁿ Q(6, 5)	441.343 (19)	11.7	3.72×10^{-3}	ⁿ R(7, 5)	1050.737 (30)	8.22	2.57×10^{-2}
^t R(3, 0)	141.847 (10)	7.20	5.97×10^{-5}	^t R(7, 0)	455.294 (20)	30.5	3.15×10^{-2}	ⁿ R(6, 6)	1146.211 (12)	2.48	3.34×10^{-3}
ⁿ P(8, 8)	170.887 (18)	15.7	3.40×10^{-4}	ⁿ Q(5, 5)	458.093 (13)	6.00	1.08×10^{-3}	ⁿ R(7, 6)	1189.072 (16)	5.22	1.63×10^{-2}
ⁿ Q(7, 3)	178.278 (19)	34.6	2.13×10^{-3}	ⁿ R(4, 2)	481.837 (15)	6.05	1.57×10^{-3}	ⁿ R(7, 7)	1336.994 (19)	2.57	5.61×10^{-3}
^t R(5, 1)	180.395 (16)	12.4	3.36×10^{-4}	ⁿ R(6, 1)	501.093 (25)	15.2	1.05×10^{-2}				
ⁿ R(2, 1)	190.662 (09)	2.41	1.76×10^{-5}	ⁿ Q(8, 6)	533.460 (17)	21.0	2.09×10^{-2}				

^a Labels for pure rotational transitions using the transition notation defined in Section III.2.^b Transition frequencies using energy data from Table 3. Reported uncertainty in the last digits (in parenthesis) is the quadrature sum of the uncertainties from Table 3.^c Dipole moments and Einstein coefficients from Ref. (56) except when marked otherwise.[†] Dipole moment and Einstein coefficient taken from (66). The error in the reported values of A_{ij} and μ_{ij} have been corrected as pointed out in Ref. (19).

adiabatic contributions. Energy levels were calculated from the modified surface using the TRIATOM program suite (73). Results of these calculations gave the best *ab initio* values at the time, but were still inferior to the calculations using the fitted potentials.

Three years later Cencek, Rychlewski, Jaquet, and Kutzelnigg (CRJK) calculated a new *ab initio* PES (12), taking into account both the diagonal adiabatic and relativistic corrections, and claimed an accuracy of a few hundredths of a cm^{-1} . Jaquet *et al.* (Ja98) (74) then calculated energies from this surface using TRIATOM. Jaquet *et al.* considered the different choices of mass: the average mass (proton mass plus 2/3 electron mass denoted **NU23**), nuclear mass (**NU**), atomic mass (**AT**), and reduced mass (**RE**). Using the same PES and DVR3D, Polyansky and Tennyson (Pol99) (75) calculated energy levels but attempted to simulate the nonadiabatic effects by using a different mass for rotational and vibrational motion. The rotational masses in their work were set to the nuclear value and the vibrational masses were set to a scaled atomic mass. Similarly, Jaquet (15) (Ja99) calculated energies of the CRJK PES using NU23 masses for motion along the $R(\text{H}-\text{H}_2)$ and $r(\text{H}_2)$ coordinates and NU masses for motion along the $\theta(R, r)$ coordinate (Jacobi-type scattering coordinates), which he denotes as **NUVR**.

The coordinate systems used in Watson's and in TRIATOM and DVR3D calculations cannot handle the kinetic energy singularity that occurs at the barrier to linearity ($\sim 10,000 \text{ cm}^{-1}$) when using the usual Morse oscillator basis functions.⁵ In 1989, Whitnell and Light (79) introduced hyperspherical coordinates, which properly treat the linear regions of the potential, to the calculations of H_3^+ . Their methodology limited the calculations to $J = 0$ levels, but this limitation was later overcome by Bartlett and Howard (80). Initially, calculations using hyperspherical coordinates were performed on the MBB PES (79–82) but later used the more accurate RKJK surface (Ali95) (83) and very recently the CRJK surface (Ali01) (84). Ali95 and Ali01 were

⁵ At the barrier to linearity one of the moments of inertia vanishes, causing some of the terms in the kinetic energy Hamiltonian to become singular. The terms that become singular depend on the coordinate system used. These singularities impose boundary conditions on the basis functions which are not met when using the common coordinate systems with the convenient Morse oscillators. Instead, artificial barriers must be applied to the potential to keep the calculations from diverging (consequently these calculations are expected to give poor results at energies near and above the barrier to linearity) (52), or alternative basis functions such as spherical oscillators (which are much harder to make converge) are used (56). References (76–78) discuss this computational problem in more detail.

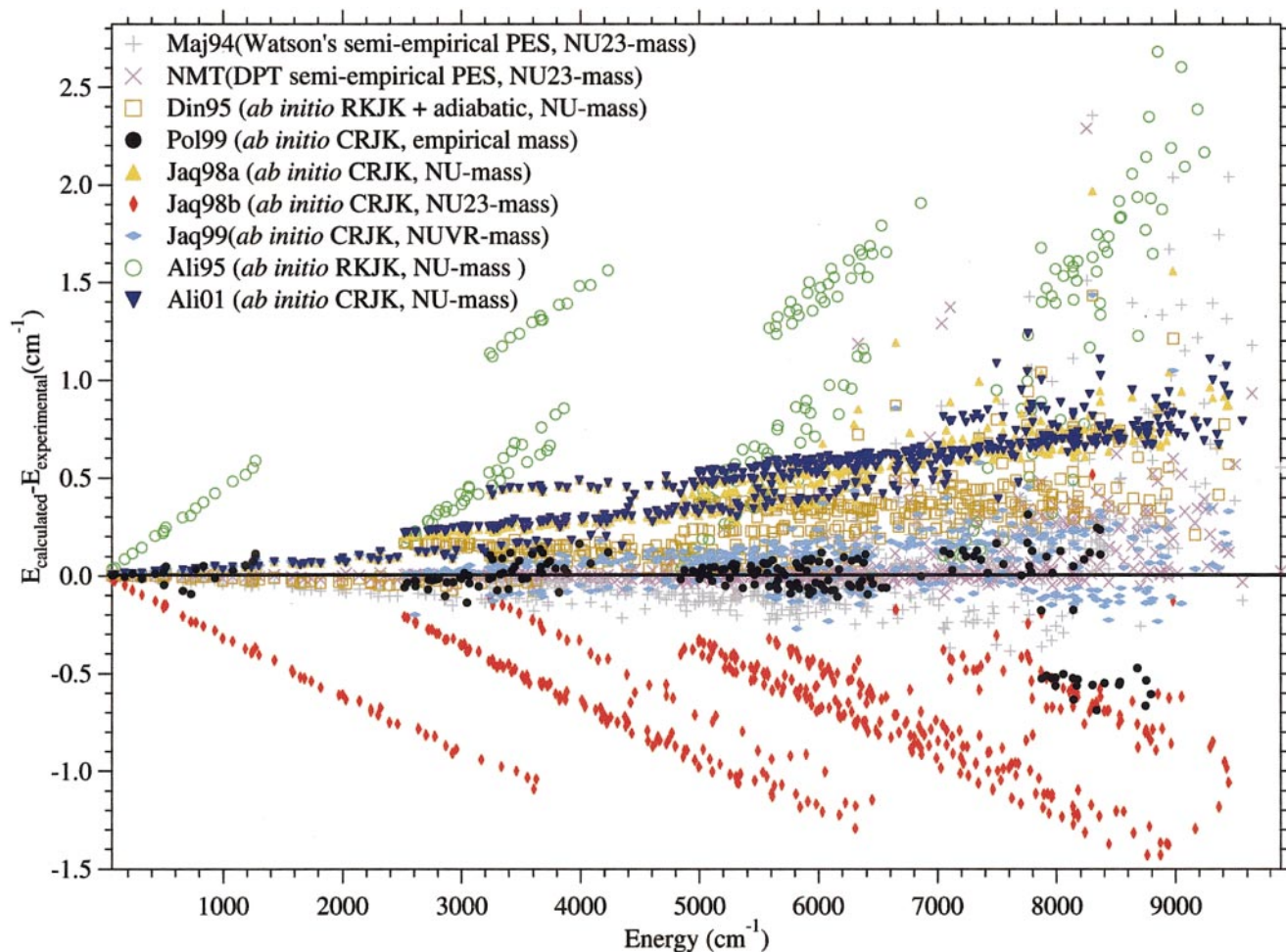


FIG. 3. Comparison of the latest calculated energy levels to experiment. Labels used in the legend are defined in Section IV.2. A printable color version of this figure is available online (39).

performed using nuclear masses, and in Ref. (13) the authors discuss the merits of this approach.

IV.2.2. Qualitative Comparison

To evaluate each of the calculations, we have plotted the difference in the calculated and experimental values for all experimentally determined levels (Fig. 3). This diagram clearly depicts the dependence of each calculation on vibrational state, rotational energy, and general scatter, which is useful in analyzing the effects of the various theoretical approaches. While a detailed analysis of each of these calculations is beyond the scope of this paper, we would like to make several qualitative remarks that are apparent from our comparison to experimental data:

1. Semiempirical vs ab initio approaches. The semiempirical calculations give the most accurate results at energies where data were available at the time of the fit. At higher energies, where experimental data was sparse, the agreement is consider-

ably poorer. In these cases, *ab initio* calculations perform better due to their more systematic residuals.

2. PES and non-BO corrections. There is considerable difference between Din95 and Ali01 (and Jaq98a) which both use NU-mass, but use different PESs. This suggests that the introduction of relativistic effects (as done in CRJK, but not in RKJK + adiabatic) may increase the energy residuals or that the diagonal adiabatic contribution is treated differently in the two calculations.

3. Choice of mass. The large rotational dependence of the residuals of Jaq98b implies that NU23 calculations produce too large a moment of inertia and consequently underestimates the rotational dependence of the energies. While the scaled mass systems (NUVR and empirical mass) give smaller residuals and a flatter rotational dependence, the NU-mass calculations seem to give much more systematic residuals. We cannot rule out the possibility that the scatter in Pol99 and Jaq99 is due to convergence problems.

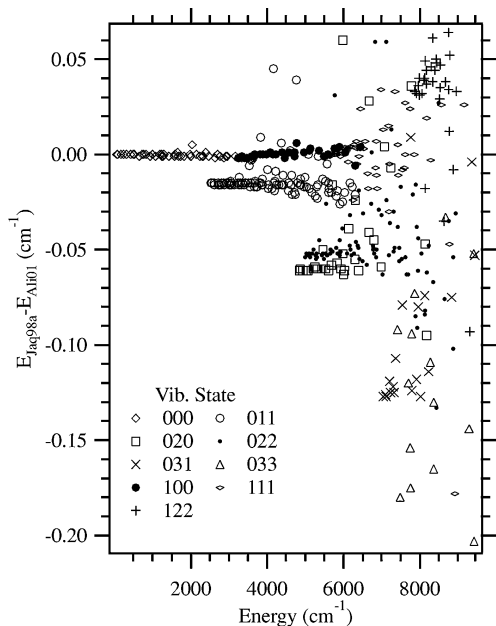


FIG. 4. Comparison of the energy level calculations of Ali95 and Jaq98a versus energy. Labels are defined in Section IV.2.

4. *Differences in similar approaches.* To verify that the observed differences in calculations are indeed due to the different approaches (and not simply due to variances from one group's calculation to another) we compared Ali01 and Jaq98a which use the same choice of mass and same PES. Their results compared quite nicely, confirming that the differences in each of the other calculations are significant. Upon close inspection, the differences between Ali01 and Jaq98a show a slight dependence on the vibrational state (see Fig. 4). This difference is on the order of 0.05 cm^{-1} and is significant when compared to the most accurate calculations of Pol99 and Jaq99. The source of this vibrational dependence is unclear, but may be due to the choice of coordinate system and/or basis set.

It is difficult to pinpoint the source of the differences in results between each computational approach due to the limited amount of computational data available. Further “experiments on the calculations” need to be performed—systematic comparisons of the energies calculated with different coordinate systems, basis sets, mass choices, and non-BO corrections will be necessary to iron out the remaining discrepancies. The experimentally determined energy levels will be instrumental in this endeavor as a powerful tool to probe the rotational, vibrational, and energy dependencies, as well as the general scatter of the various computational methodologies.

V. CONCLUSION

This work represents the end of a chapter in the laboratory spectroscopy study of H_3^+ . Energy levels for nearly every vibrational band below the barrier to linearity have been probed and

determined experimentally, many of them up to $J = 9$. Almost all of the observed lines have been assigned, and those that have not are probably due to species other than H_3^+ or have an error in the frequency measurement. While there are still transitions to be measured in this energy regime—higher J transitions in the $2\nu_2$ and $3\nu_2$ bands should be achievable with the better diode lasers and higher sensitivity available today—these will likely not lead to a better understanding of H_3^+ behavior at low energies or produce a qualitatively better diagnostic tool for theoretical calculations.

The next step for laboratory work is to make observations of states above the barrier to linearity, where some of the theoretical calculations are expected to break down. This is also the regime where the approximate quantum numbers begin to fail, and a new formalism may need to be developed to describe such levels. Such experiments are currently underway in Chicago where the $5\nu_2 \leftarrow 0$ band is being studied with a high-power Ti:Sapphire laser.

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